

# Starting your journey into neutrino physics

Mauricio Bustamante

Niels Bohr Institute, University of Copenhagen

CaféConCiencia  
August 07, 2025

UNIVERSITY OF  
COPENHAGEN



VILLUM FONDEN



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Who am I?

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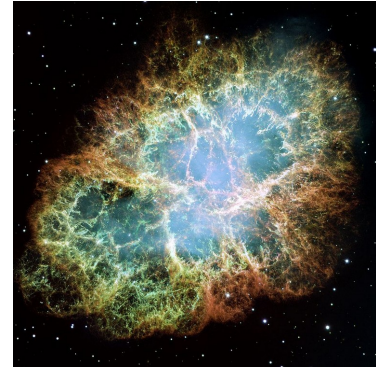
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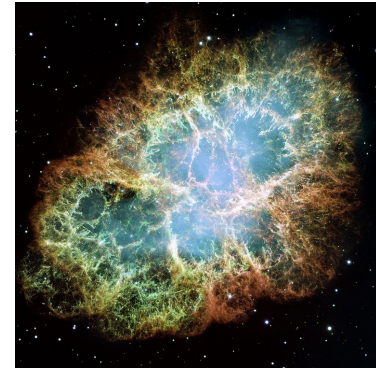
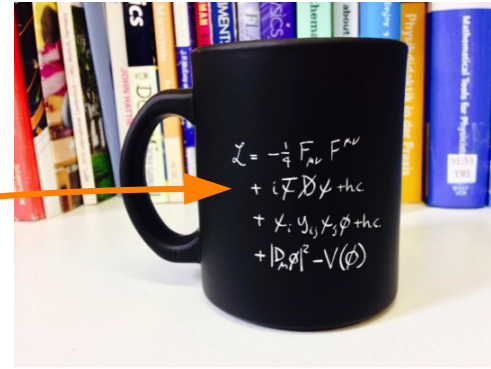
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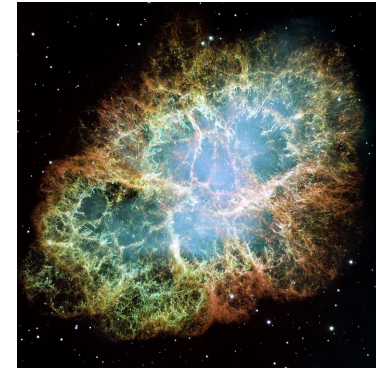
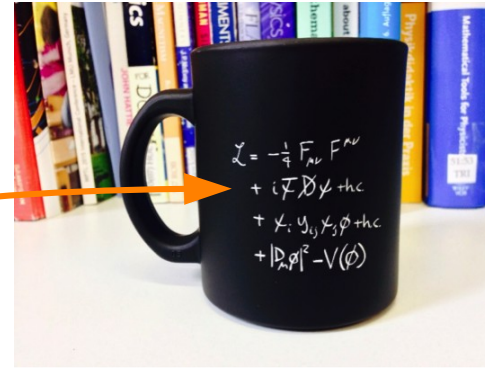
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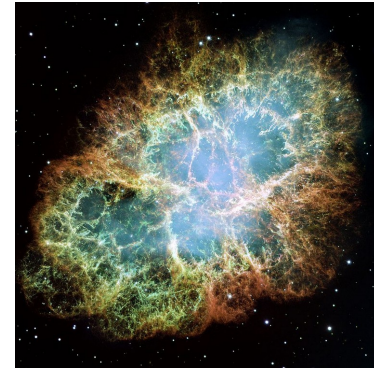
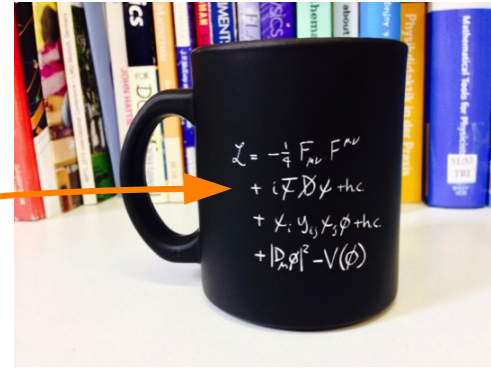
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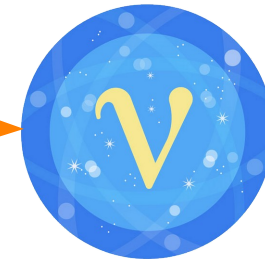
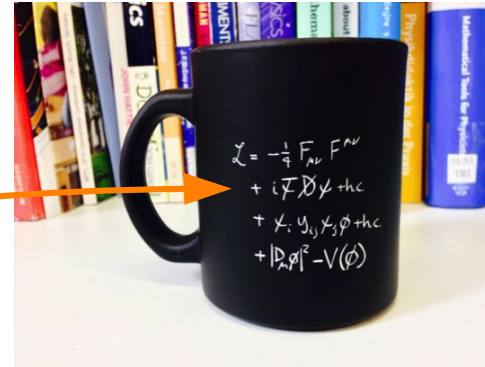
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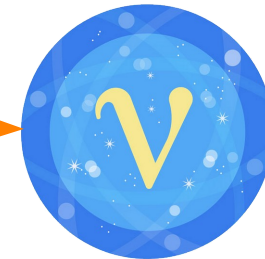
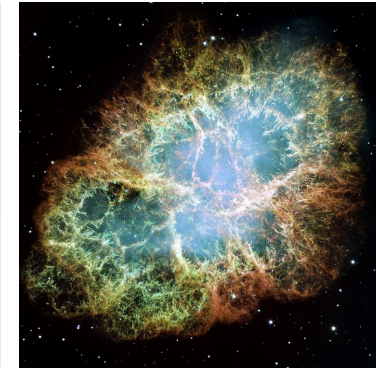
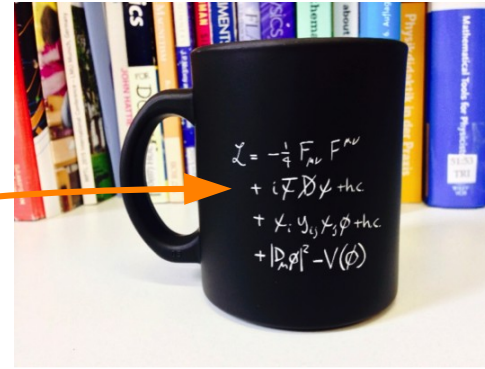
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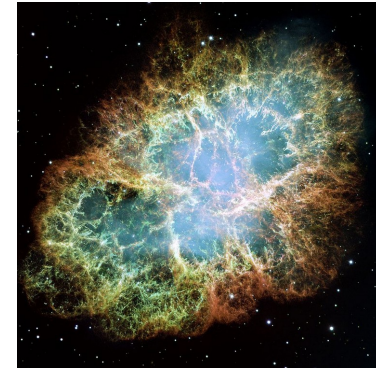
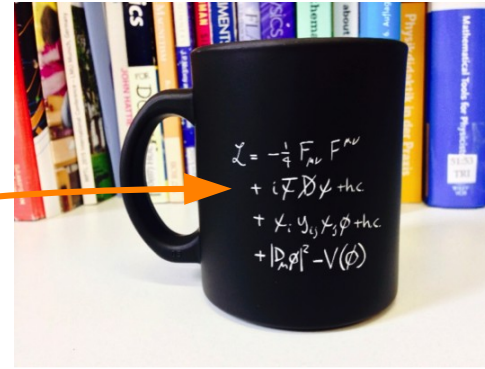
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- ▶ Sometimes I give public outreach talks



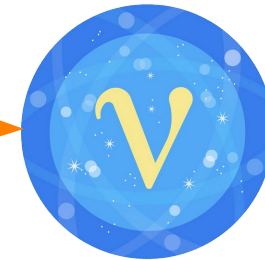
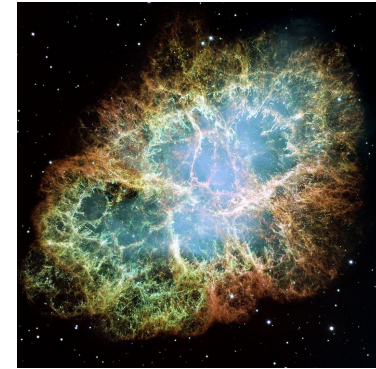
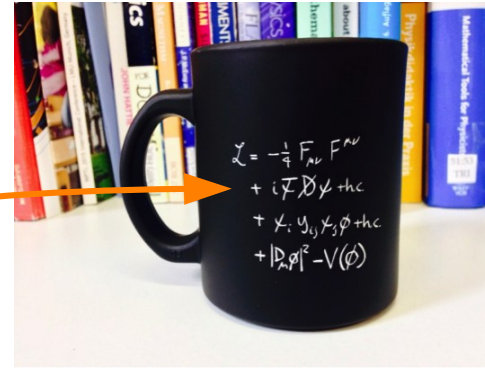
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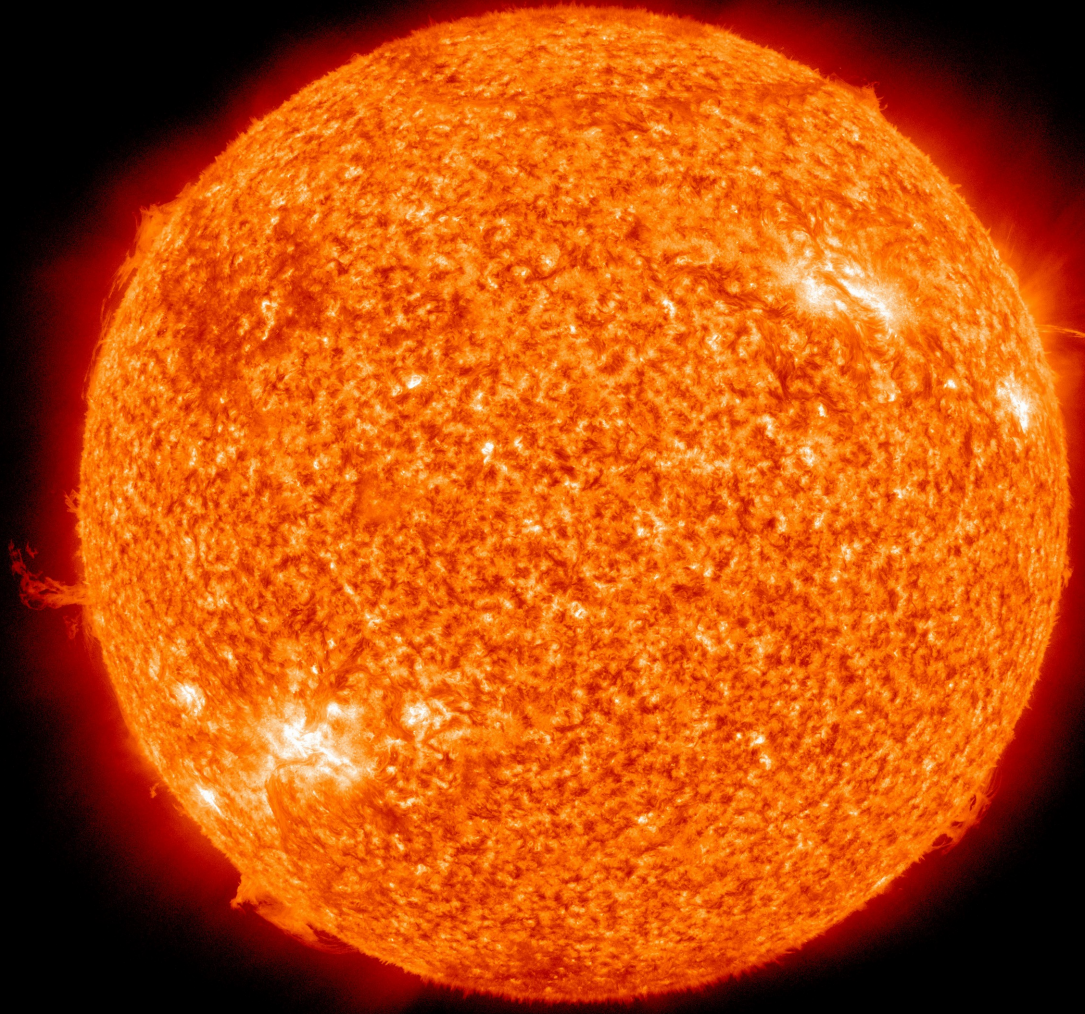
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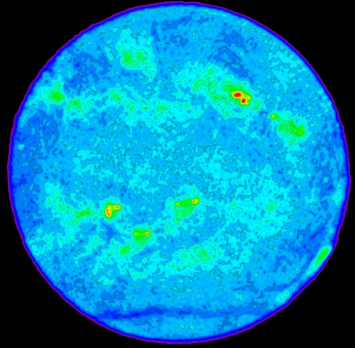
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- ▶ Professional webpage: [mbustamante.net](http://mbustamante.net)

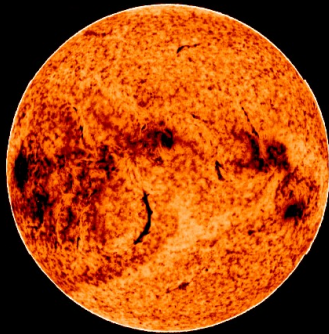




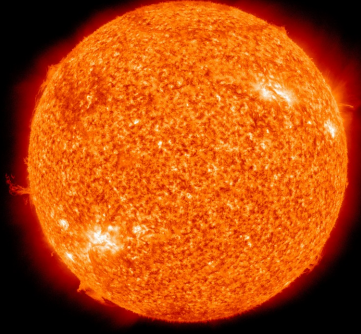
Light of longer wavelength (*i.e.*, less energetic)



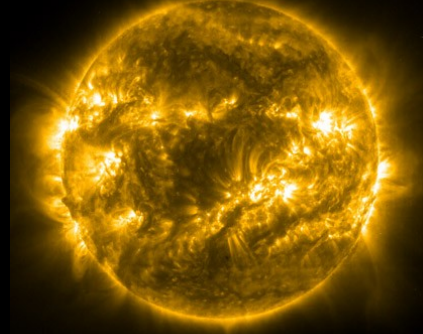
Radio



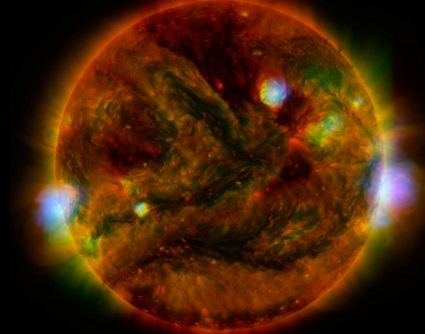
Infrared



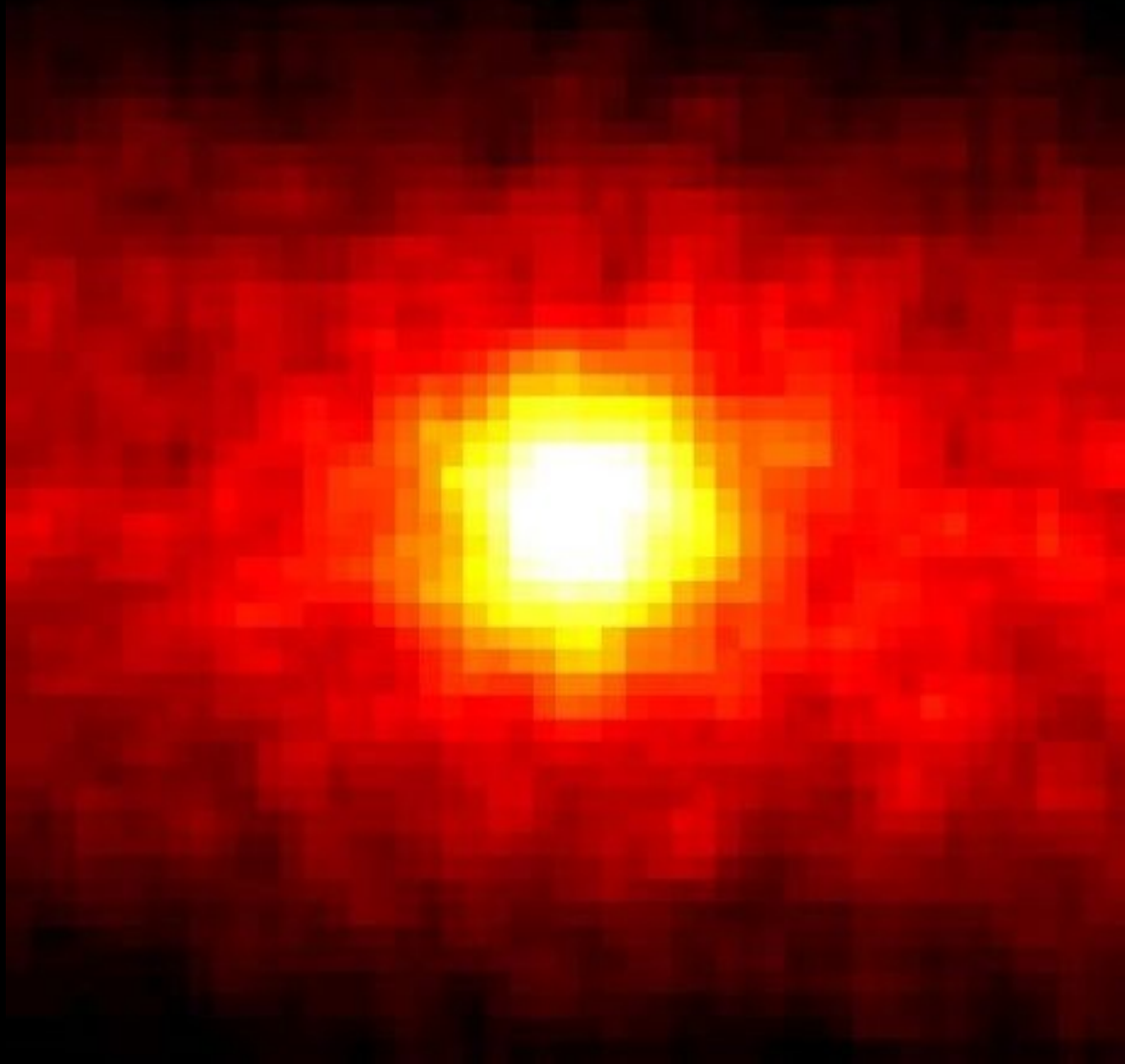
Optical



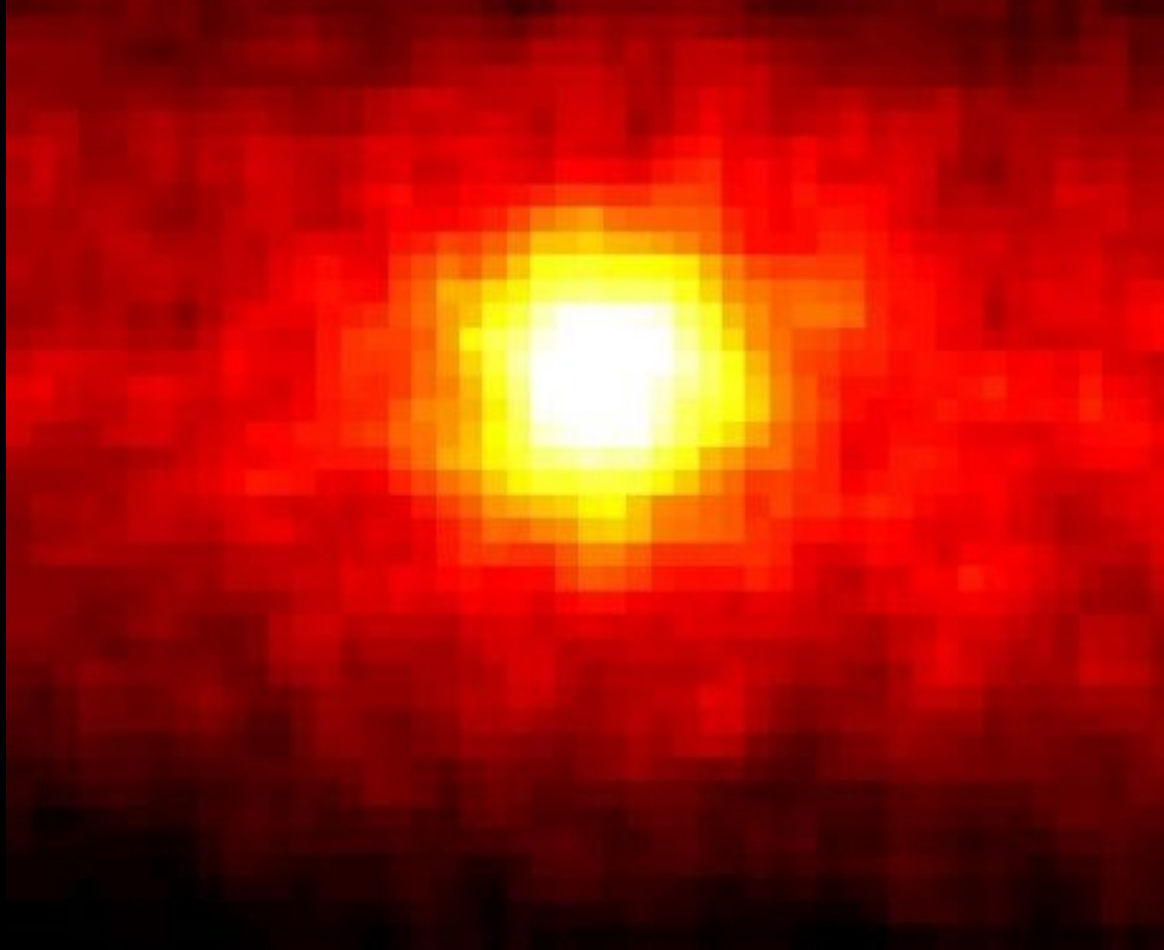
Ultraviolet



X-rays

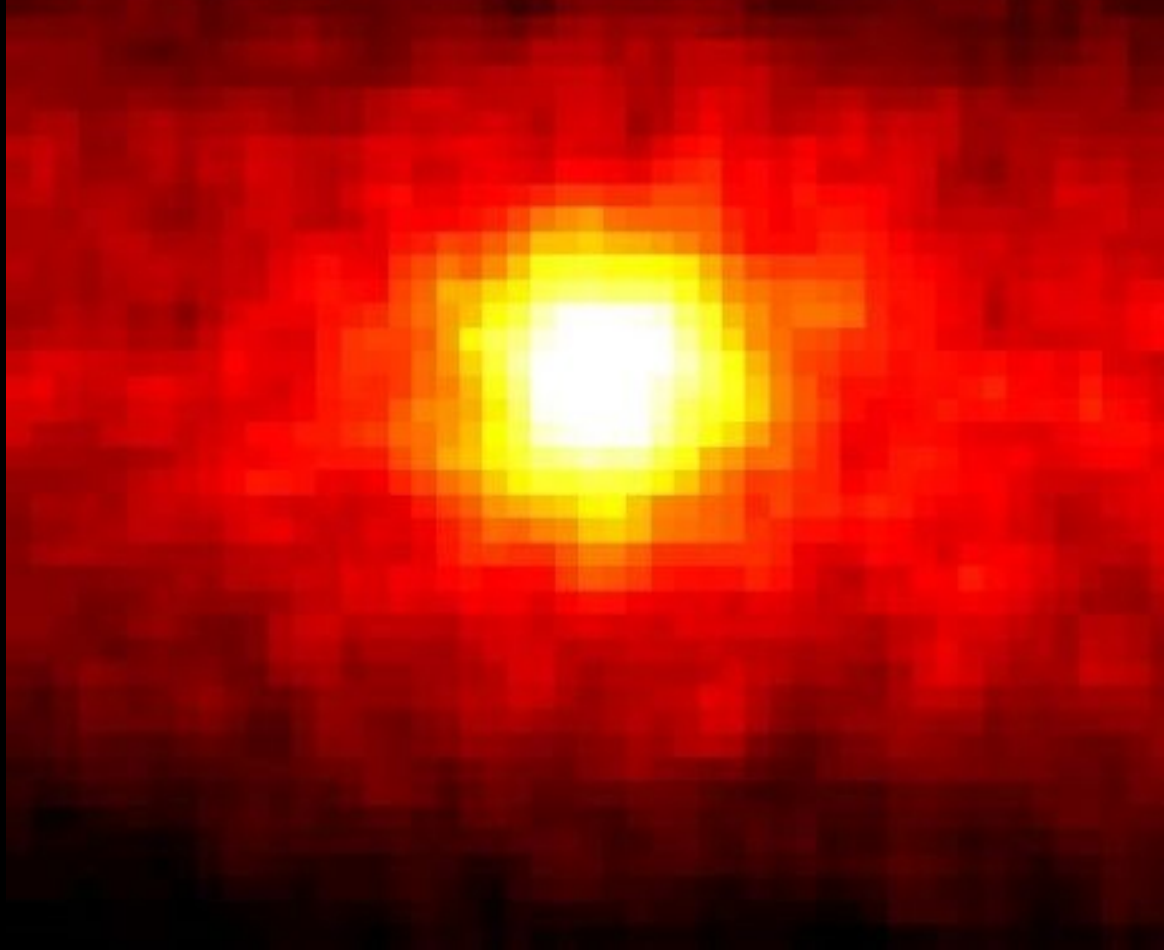


Also the Sun, but now seen in  
*neutrinos* ( $\nu$ ), not light!



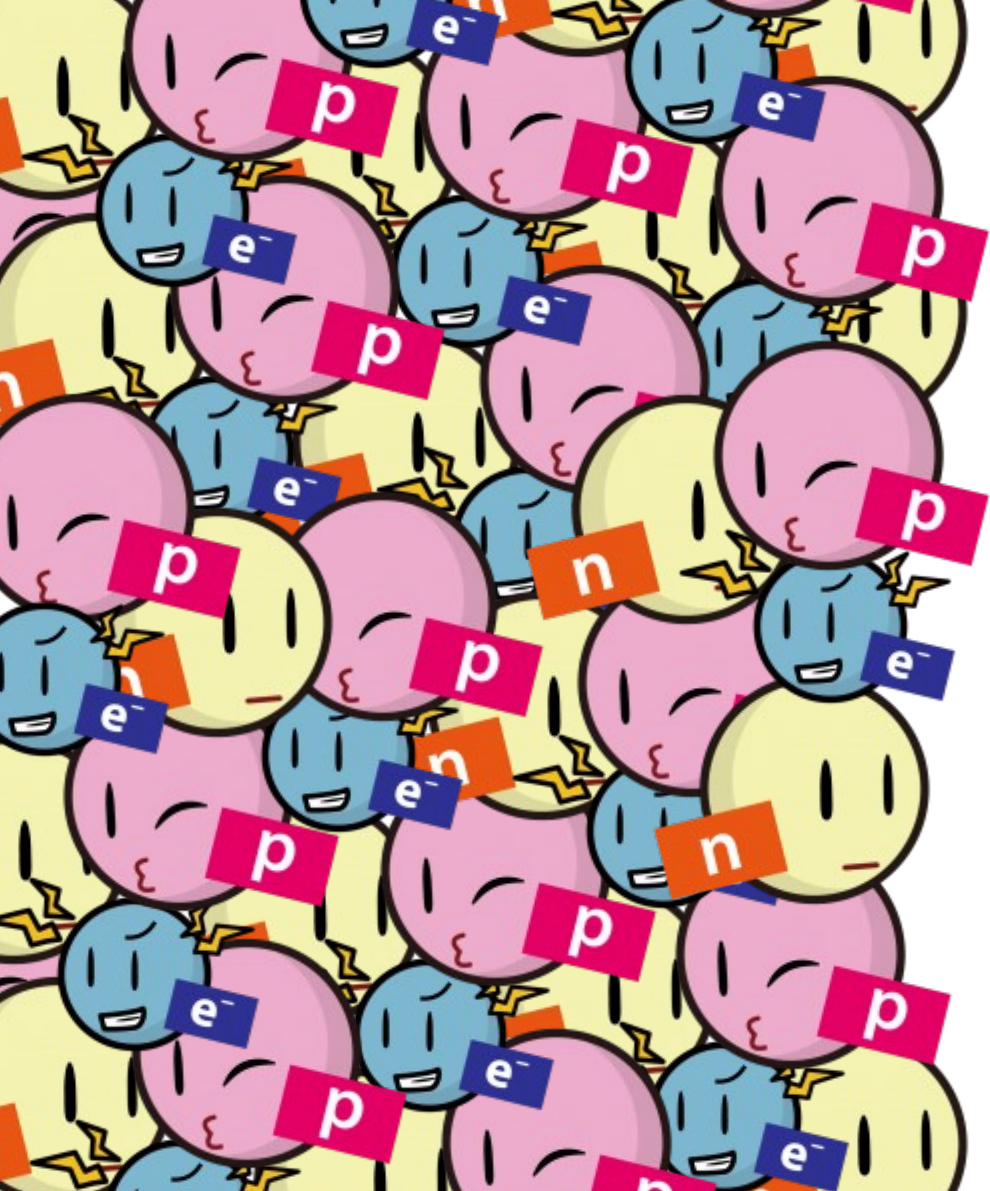
~100 billion  $\nu$   
cross every  $\text{cm}^2$   
of your body  
every second

Also the Sun, but now seen in  
*neutrinos* ( $\nu$ ), not light!



~100 billion  $\nu$   
cross every  $\text{cm}^2$   
of your body  
every second

... but only a  
handful collide  
with our  
detectors!



Normal matter is made of

neutrons,



protons,



and electrons



Neutrinos are *not* part of  
normal matter

Neutrinos are elementary particles,

electrically neutral,

very light,

and superbly antisocial

Neutrinos are elementary particles,  
*= indivisible*

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*= so light that we don't know their mass!*

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Neutrinos are **elementary particles**,

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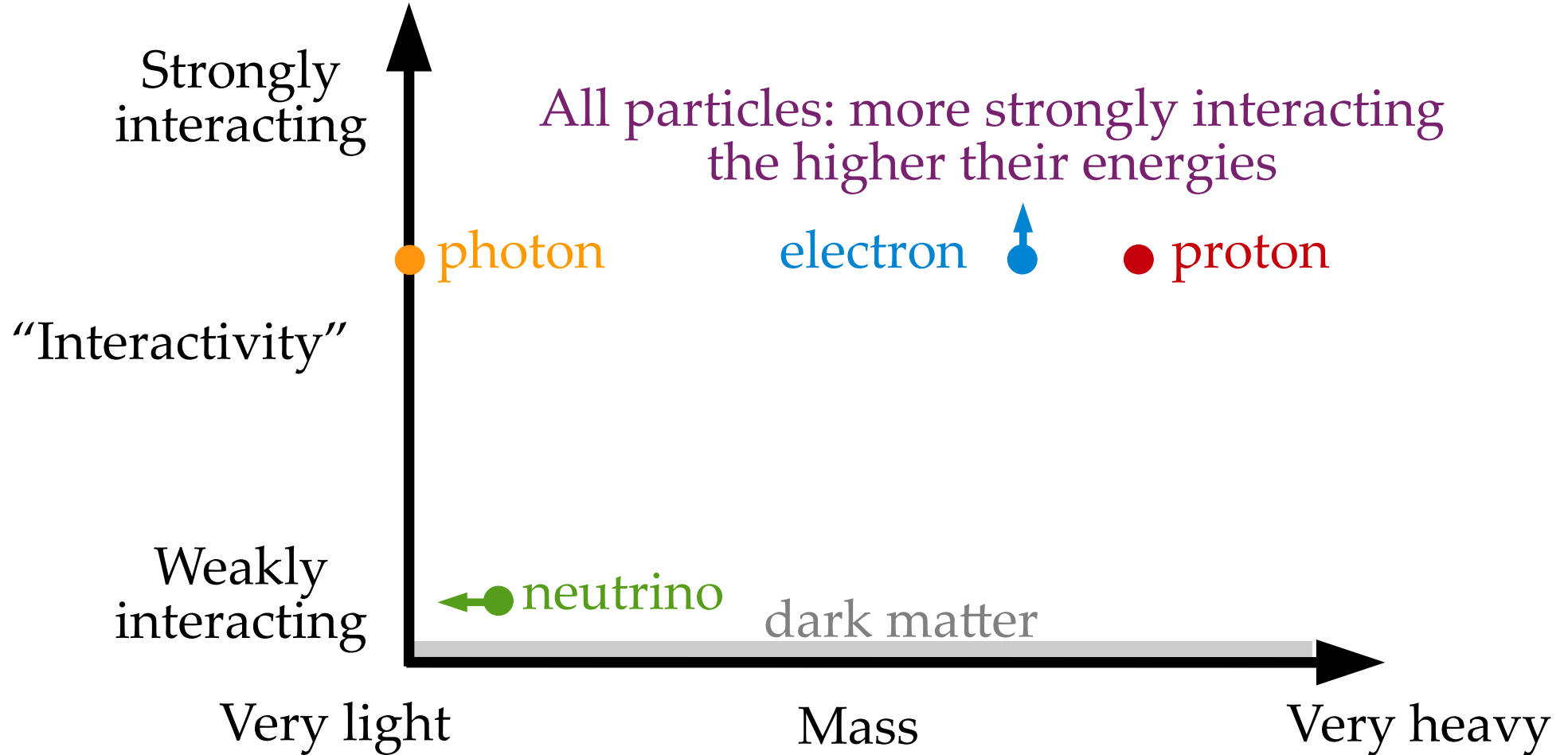
**very light**,

*= so light that we don't know their mass!*

and **superbly antisocial**

*= barely interact with matter*

# Neutrinos are *very* light and *very* anti-social



Just how weak is *weak*

Just how weak is *weak*

# Just how weak is *weak*

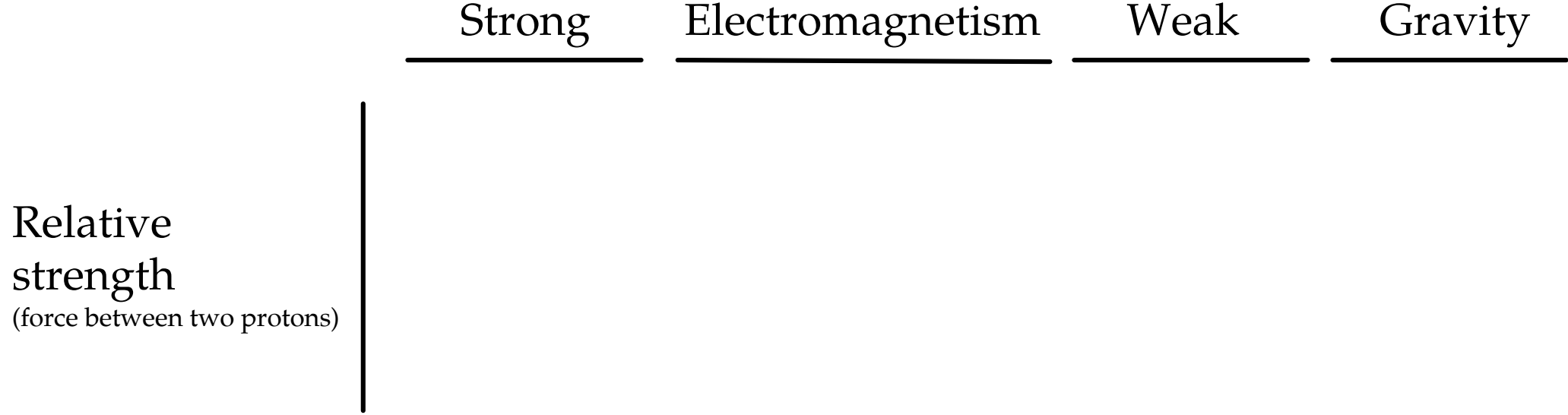
Strong

Electromagnetism

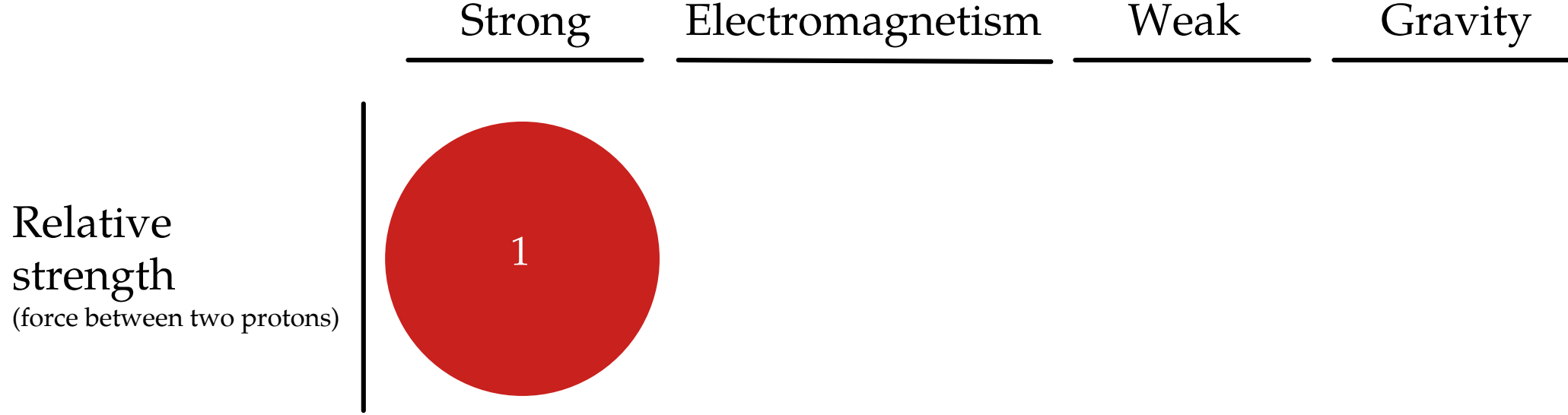
Weak

Gravity

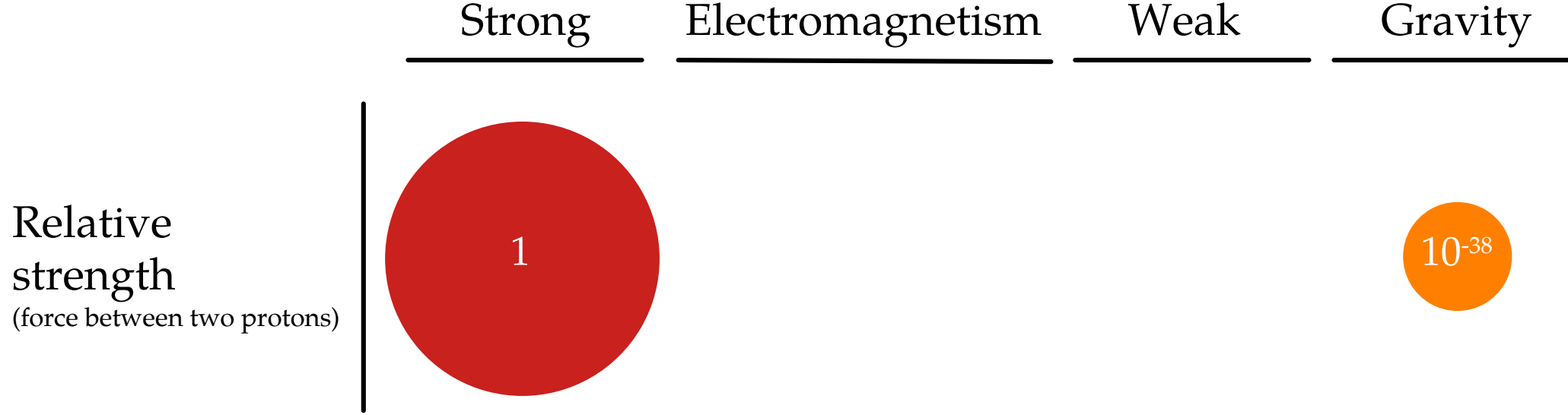
# Just how weak is *weak*



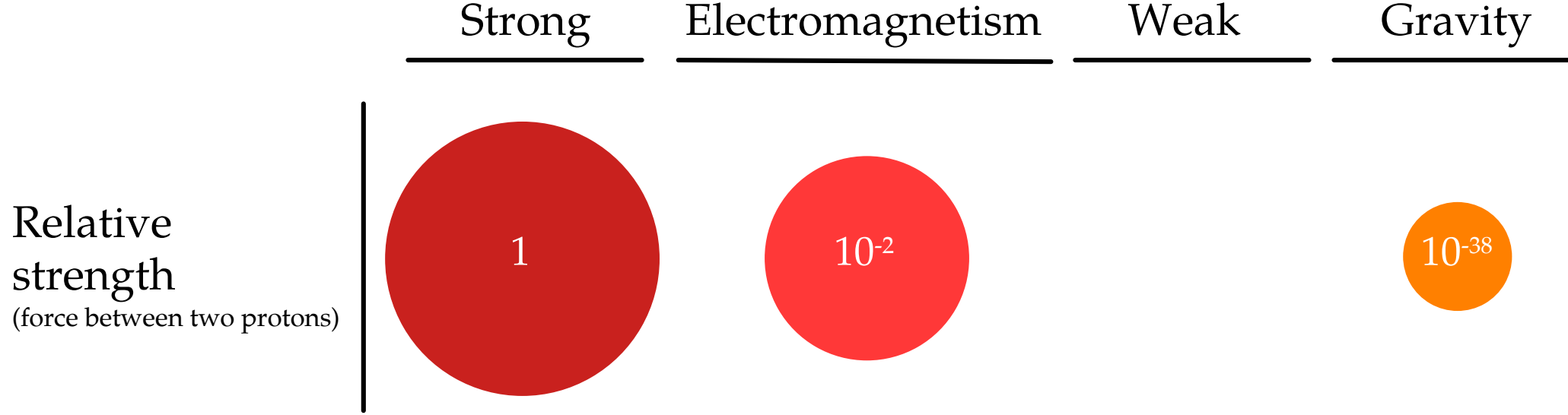
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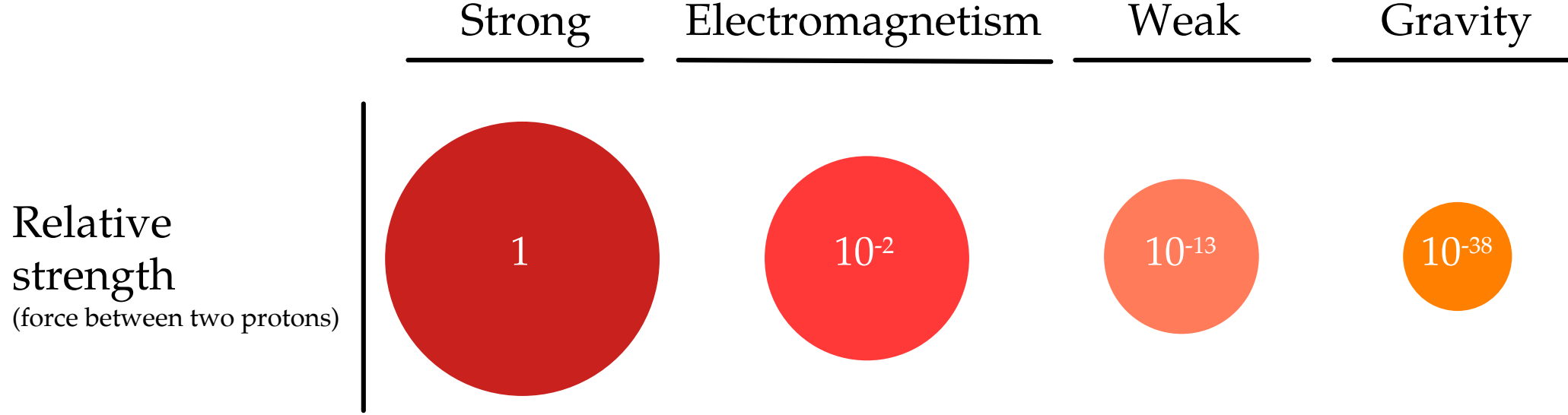
# Just how weak is *weak*



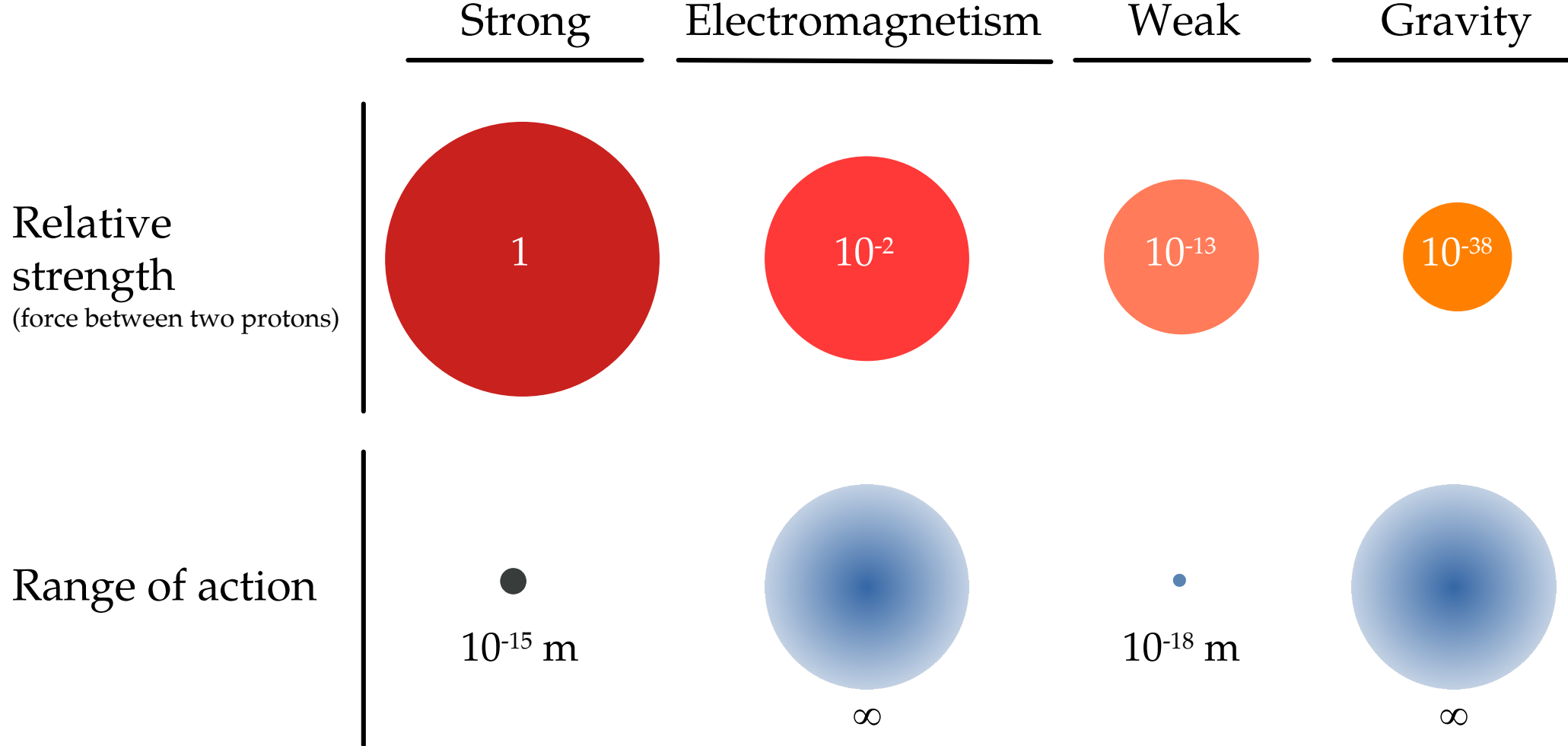
# Just how weak is *weak*



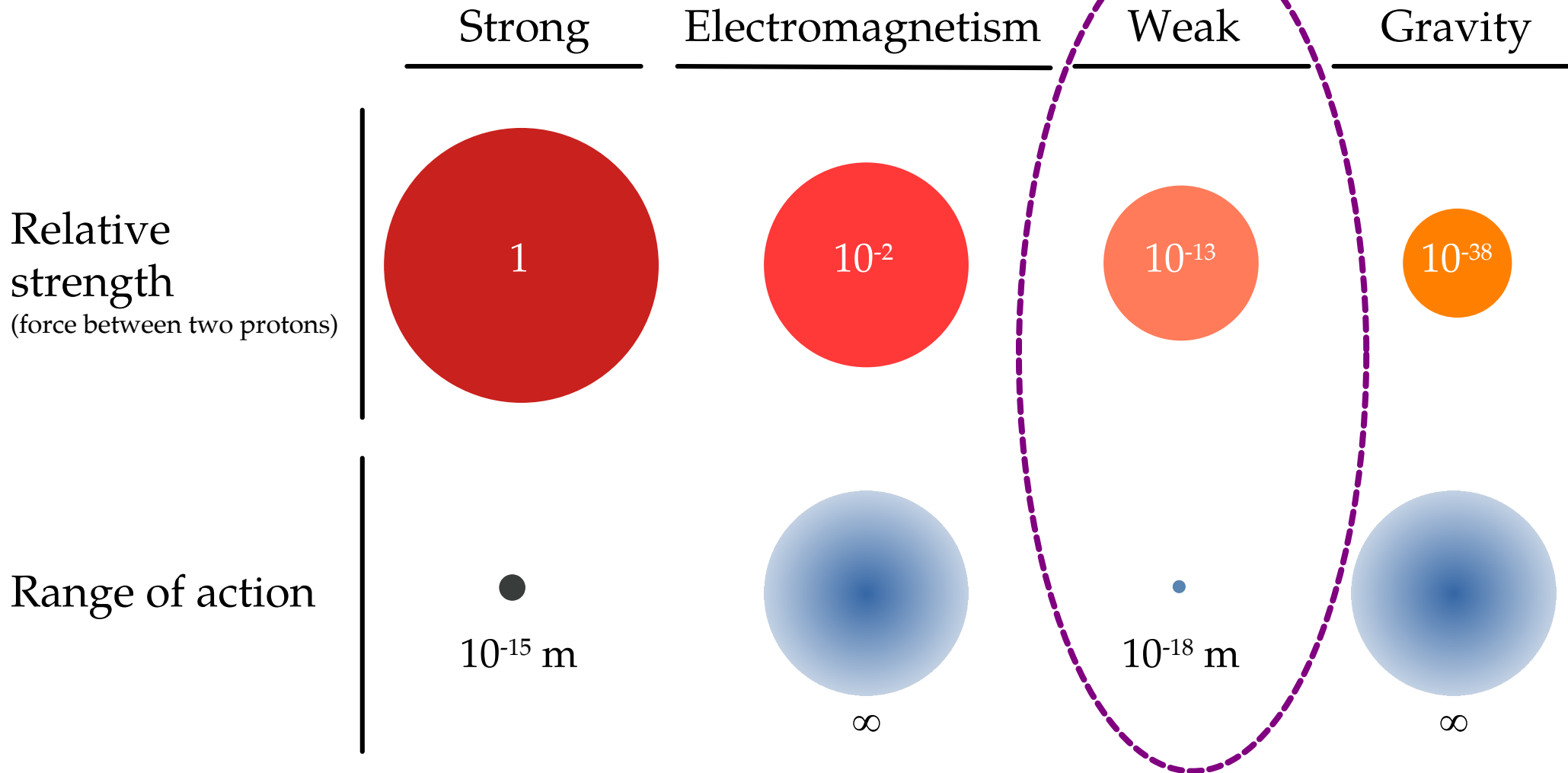
# Just how weak is *weak*



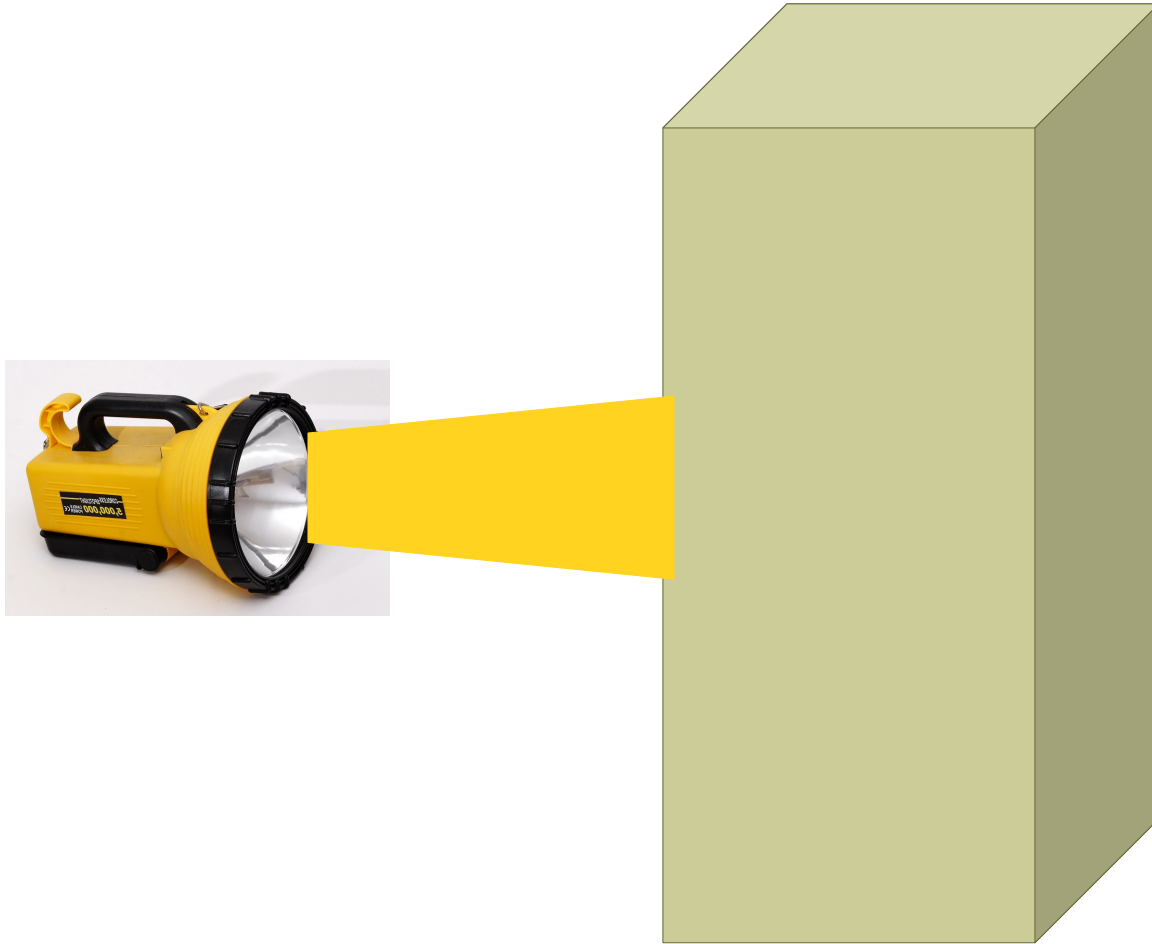
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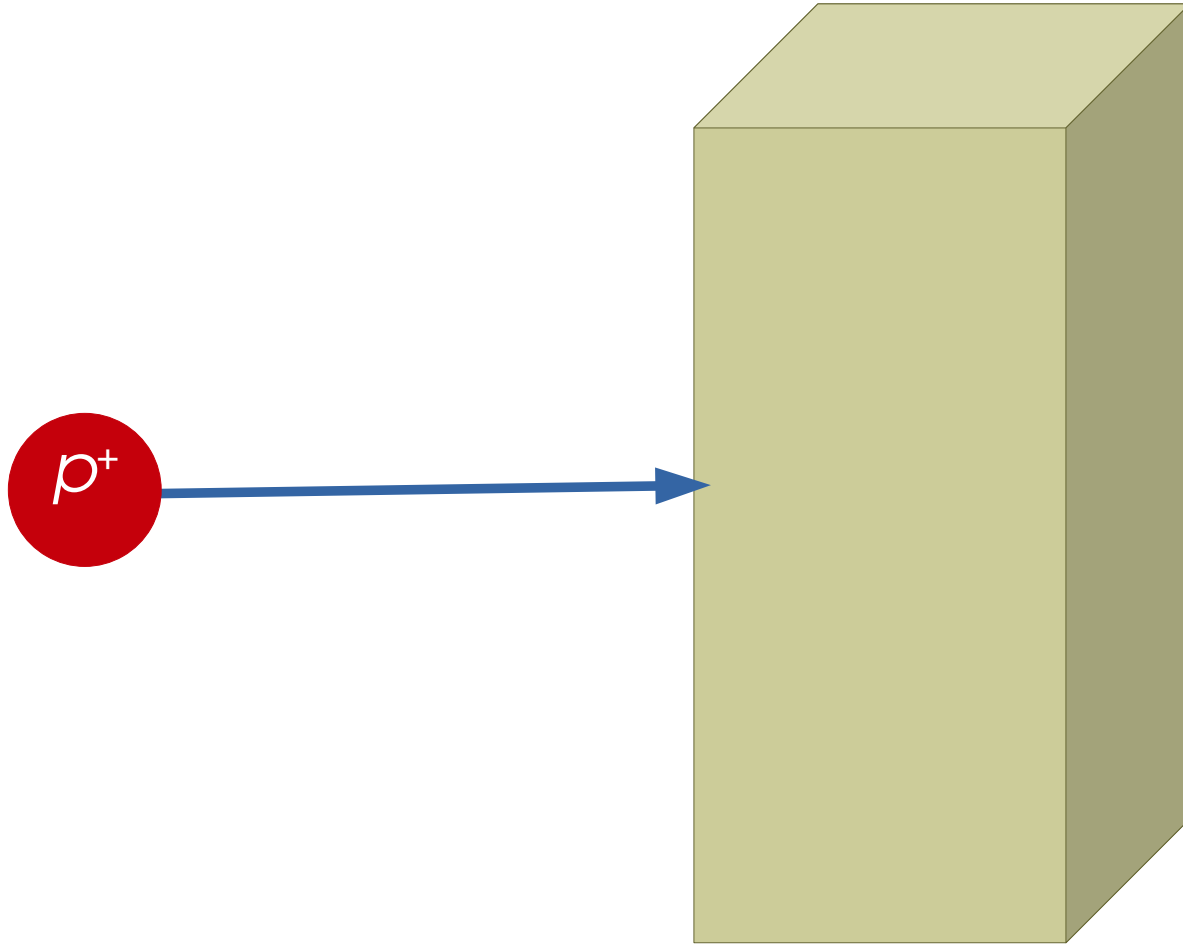
# Just how weak is *weak*



Stopping something that is *almost* not there

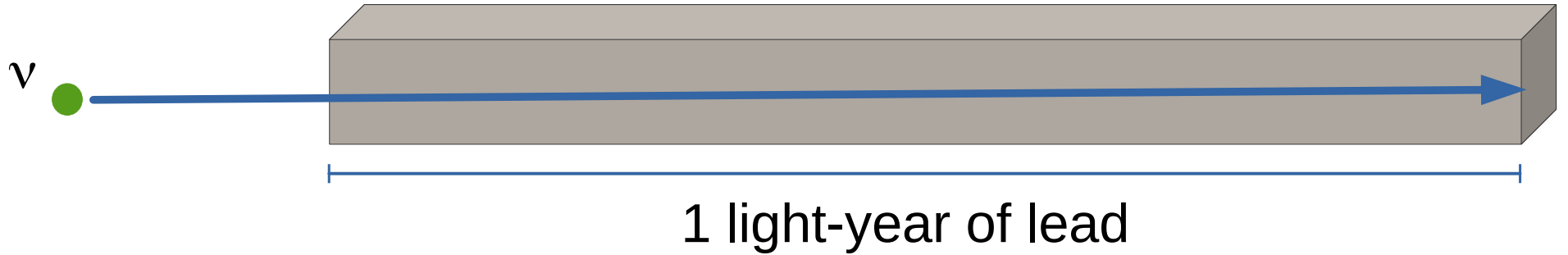


Stopping something that is *almost* not there



Stopping something that is *almost* not there

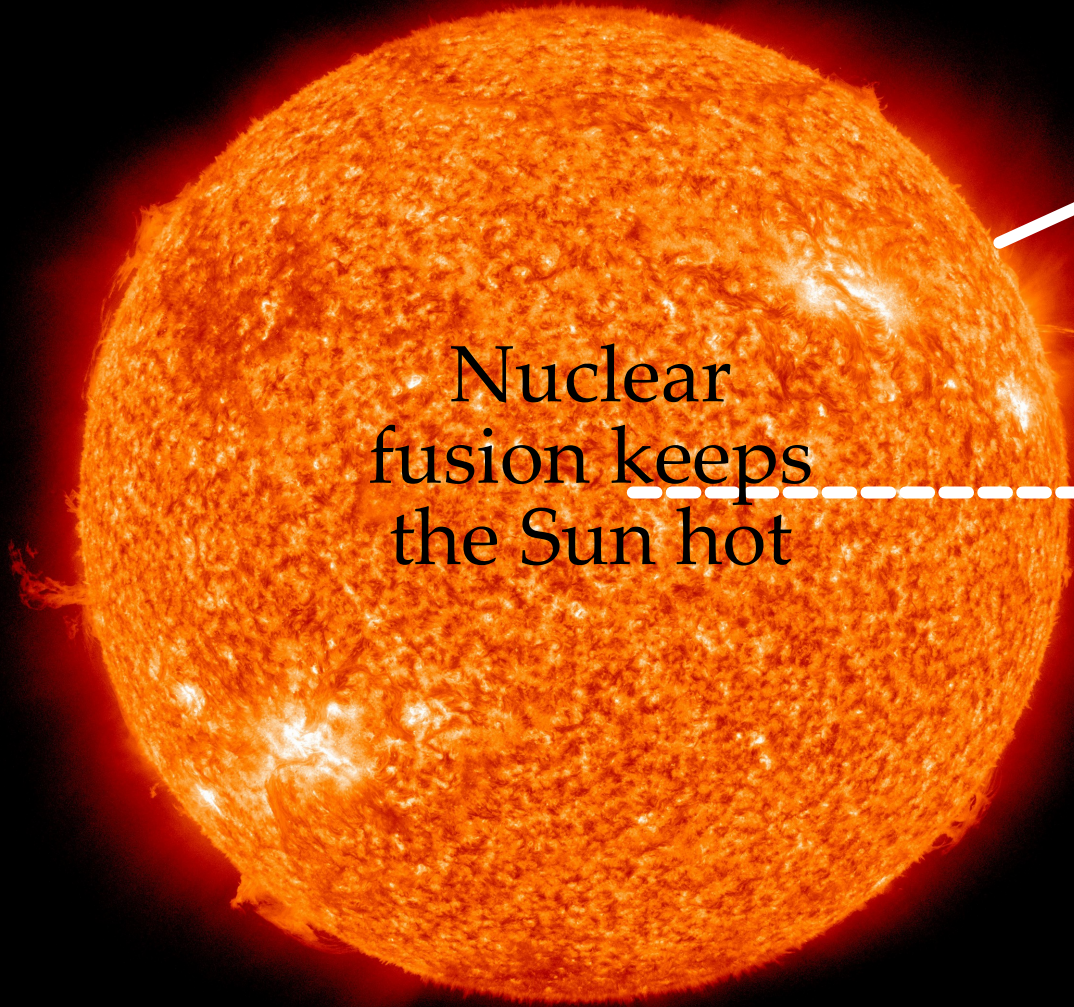
# Stopping something that is *almost* not there



Stopping something that is *almost* not there

High-energy  $\nu$





Light comes  
from the surface

Nuclear  
fusion keeps  
the Sun hot

Neutrinos ( $\nu$ ) come  
from deep within

# Neutrinos from the Sun

Nuclear fusion:



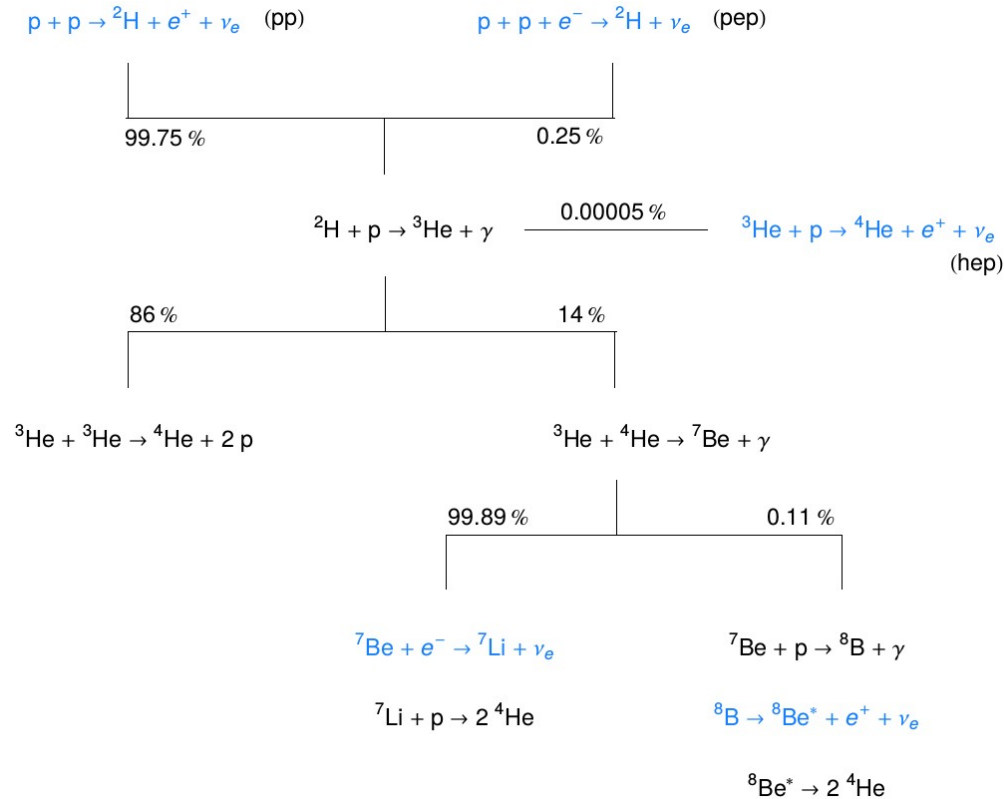
# Neutrinos from the Sun

Nuclear fusion:

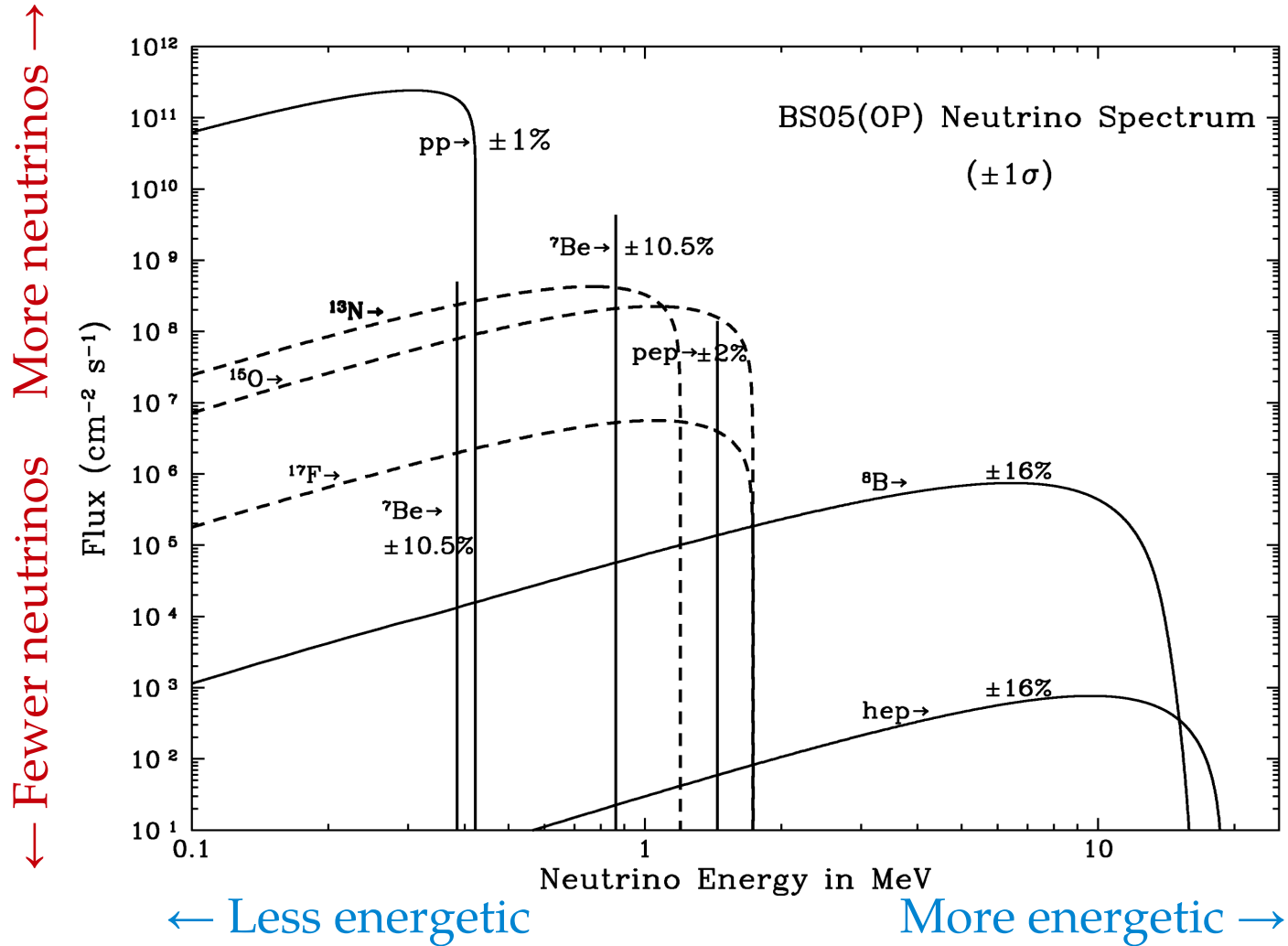


# Neutrinos from the Sun

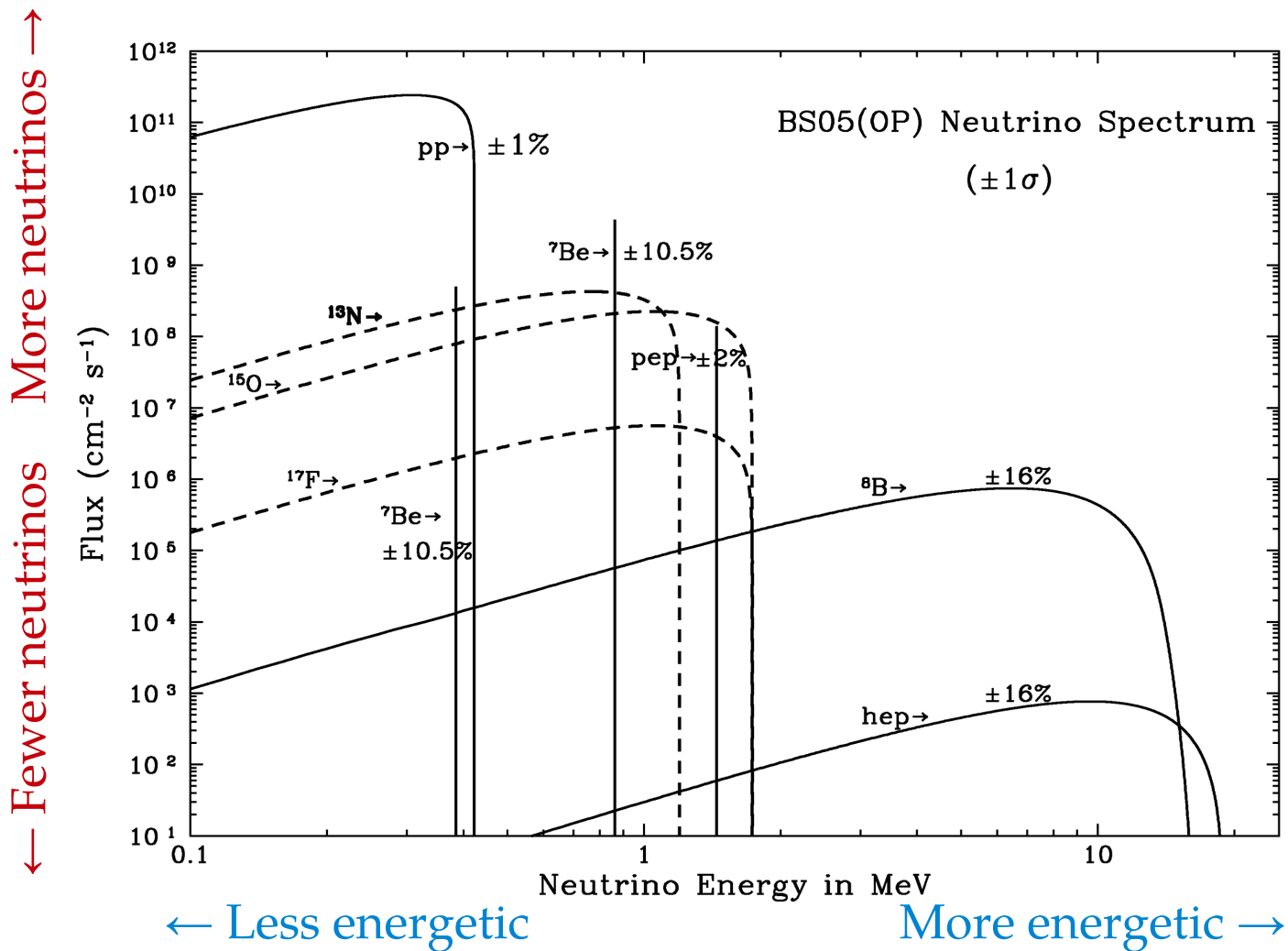
## Nuclear fusion:

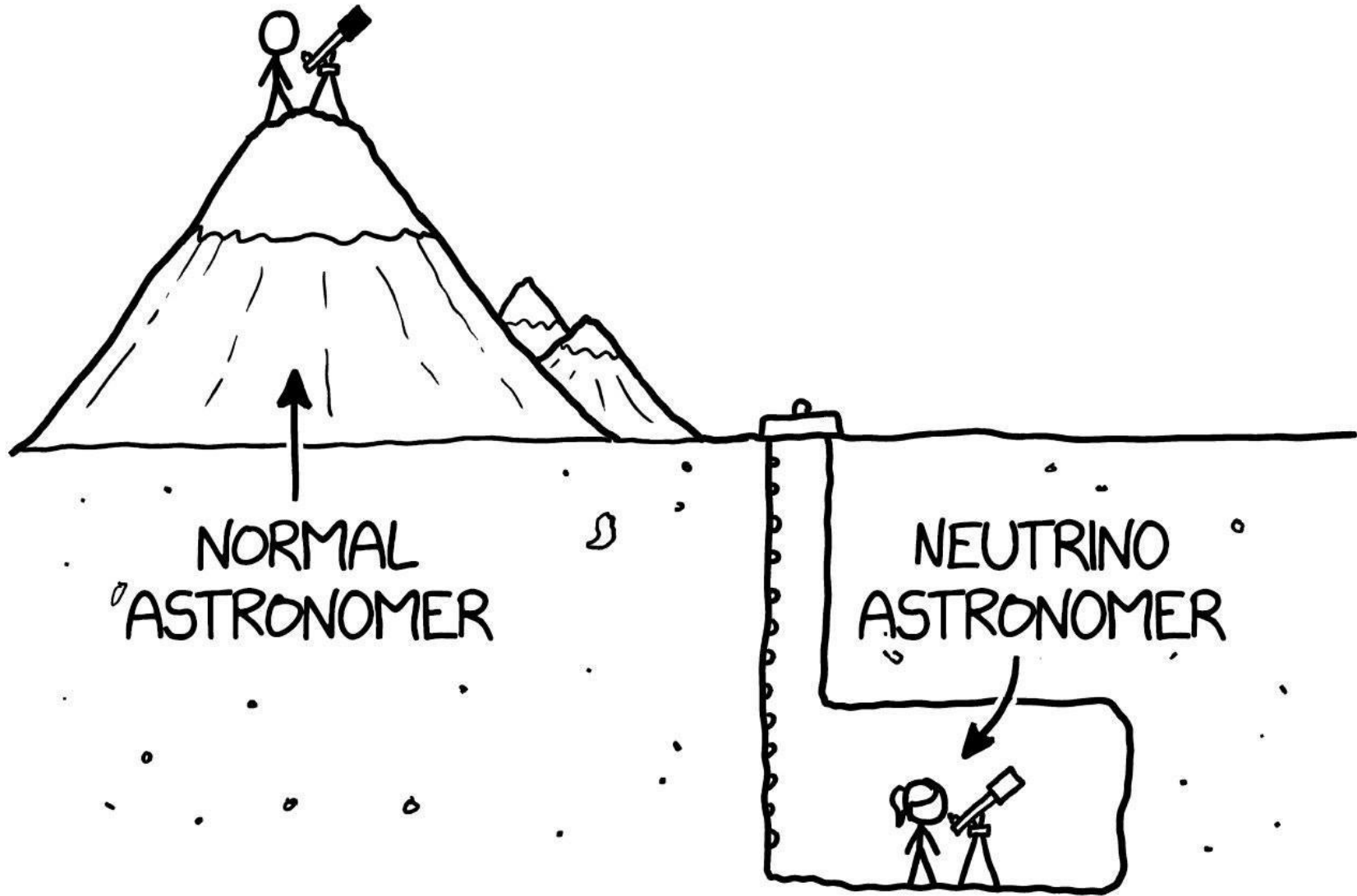


# Neutrinos from the Sun



# Neutrinos from the Sun





Space

Atmosphere

Space



Incoming cosmic ray



Atmosphere

Space

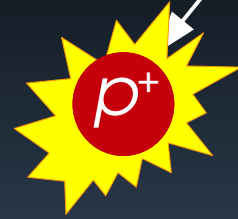
$p^+$  Incoming cosmic ray

$p^+$  Proton in the air

Atmosphere

Space

$p^+$  Incoming cosmic ray



Proton in the air

Atmosphere

Space

$p^+$  Incoming cosmic ray



Proton in the air

Pion  $\pi^+$

Neutron  $n$



Atmosphere

Space

$p^+$  Incoming cosmic ray



$p^+$  Proton in the air

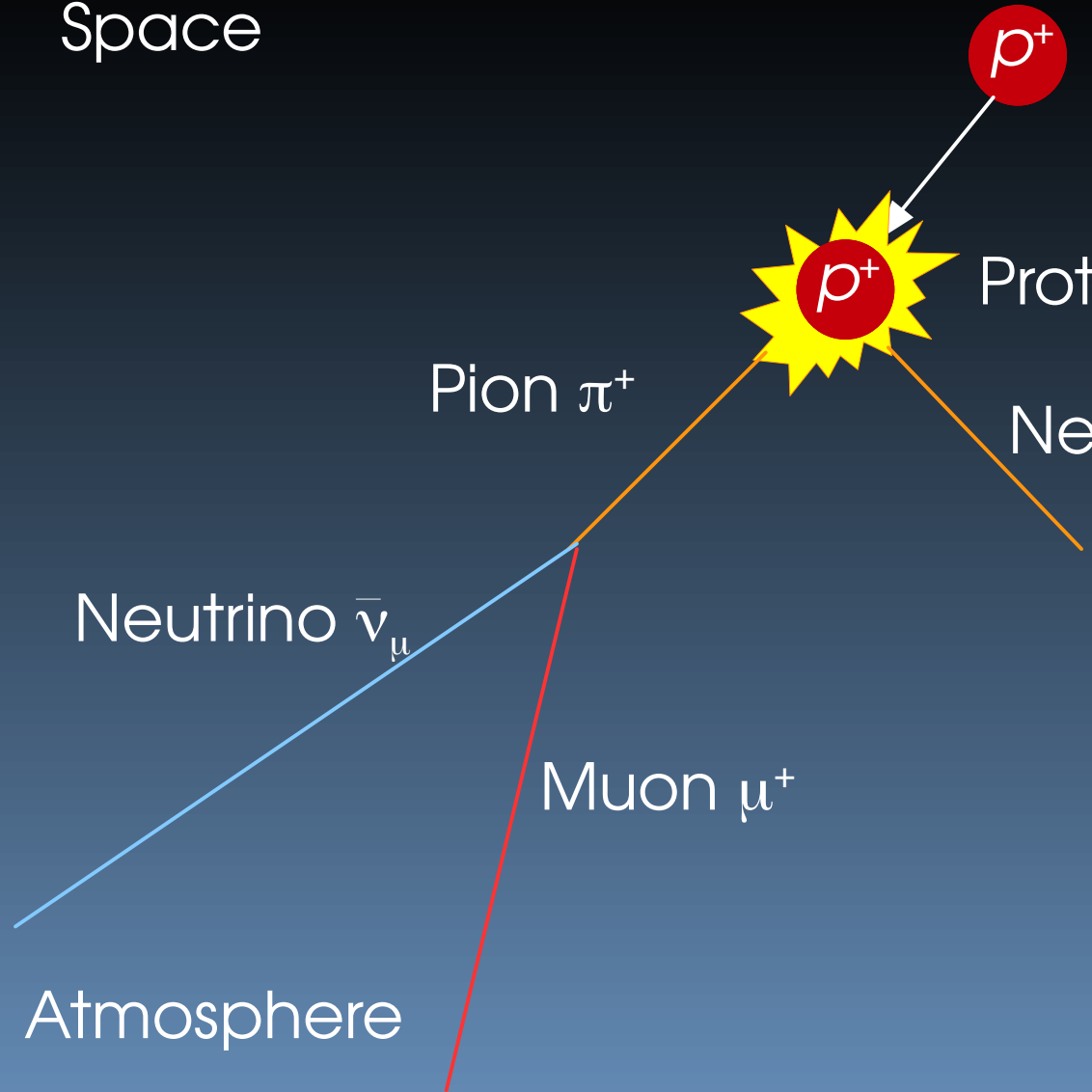
Pion  $\pi^+$

Neutron  $n$

Neutrino  $\bar{\nu}_\mu$

Muon  $\mu^+$

Atmosphere



Space

$p^+$  Incoming cosmic ray



$p^+$  Proton in the air

Pion  $\pi^+$

Neutron  $n$

Neutrino  $\bar{\nu}_\mu$

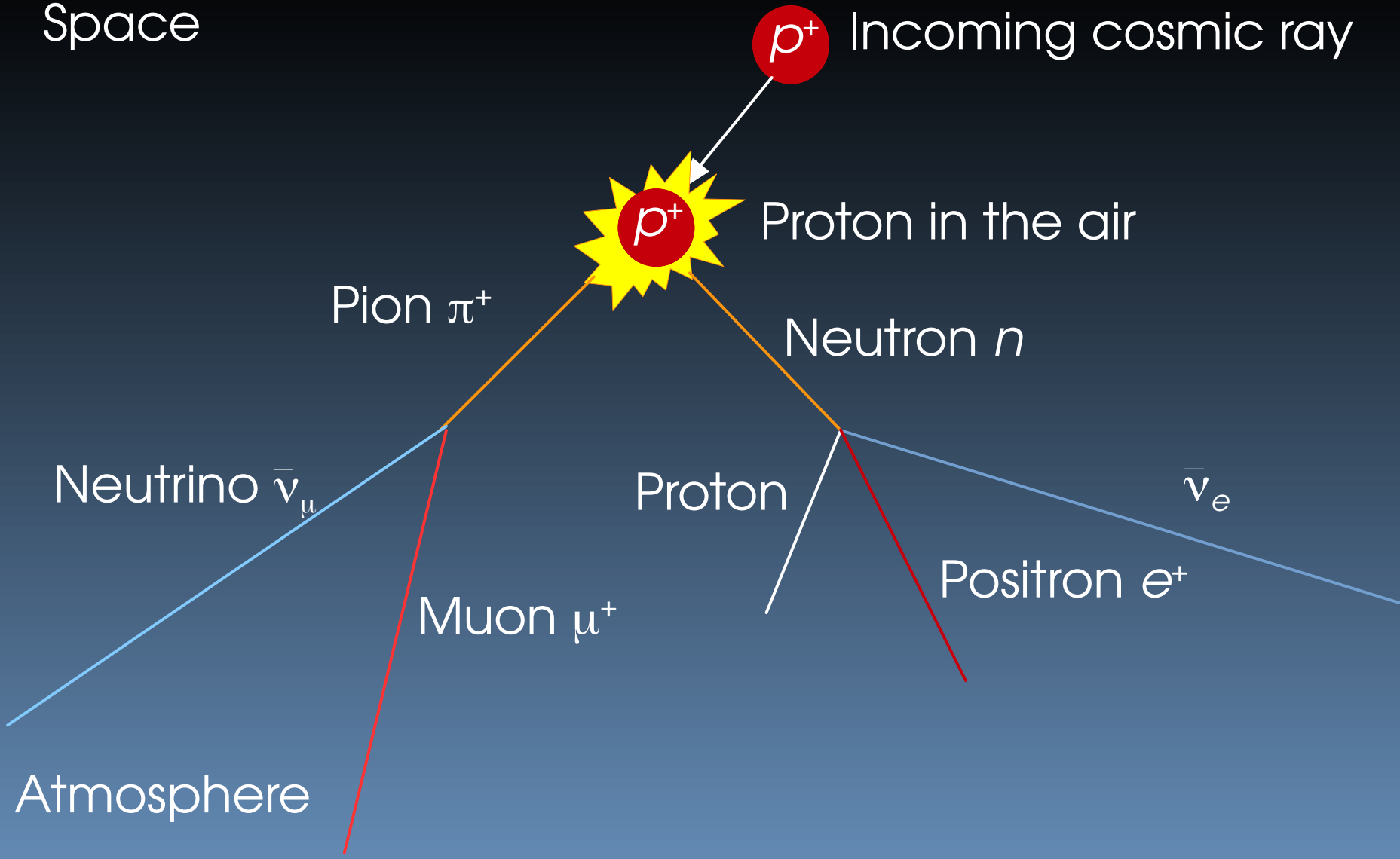
Proton

$\bar{\nu}_e$

Muon  $\mu^+$

Positron  $e^+$

Atmosphere



Space

$p^+$  Incoming cosmic ray



$p^+$  Proton in the air

Pion  $\pi^+$

Neutron  $n$

Neutrino  $\bar{\nu}_\mu$

Proton

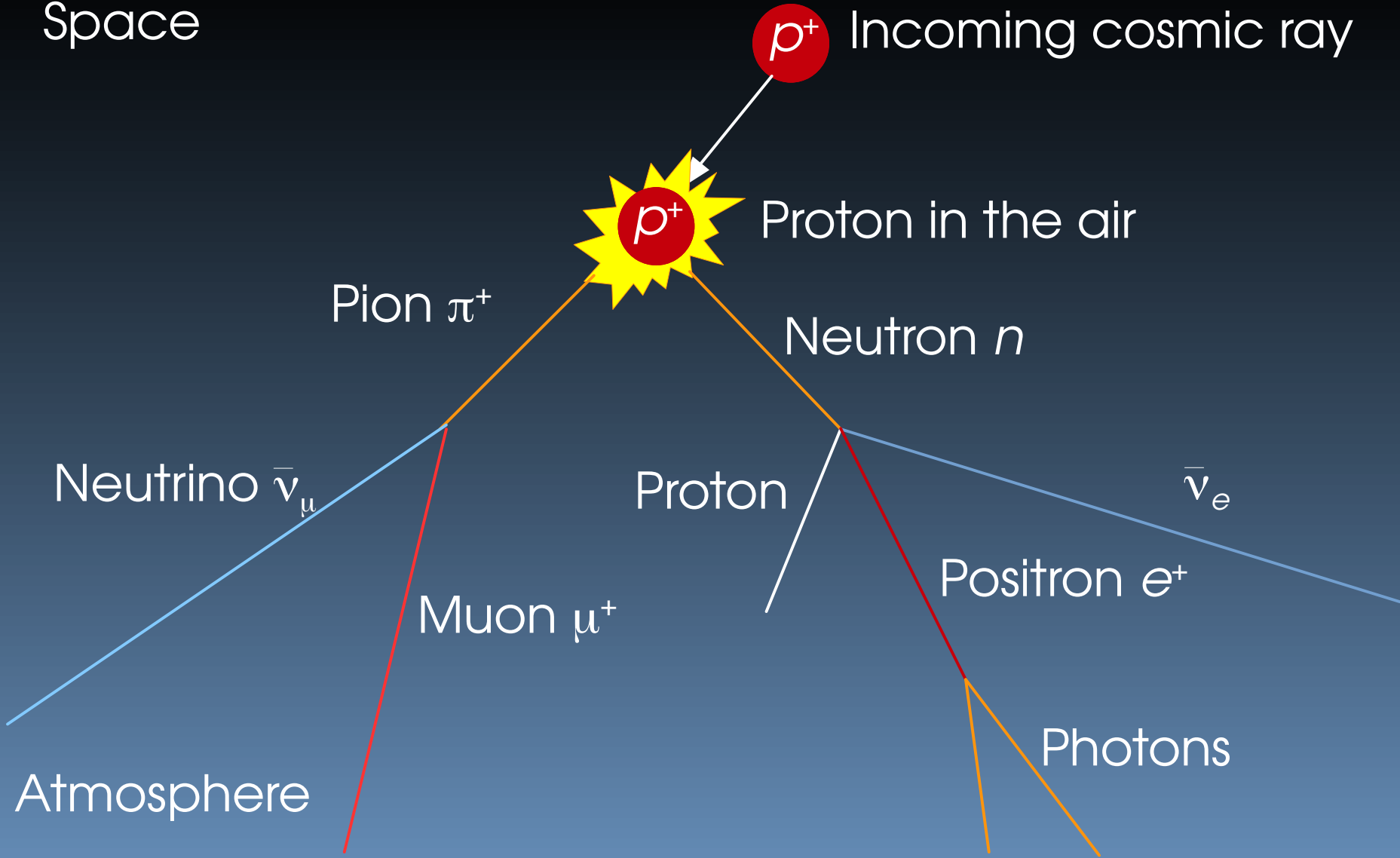
$\bar{\nu}_e$

Muon  $\mu^+$

Positron  $e^+$

Atmosphere

Photons



Space

$p^+$  Incoming cosmic ray



Proton in the air

Pion  $\pi^+$

Neutron  $n$

Neutrino  $\bar{\nu}_\mu$

Proton

$\bar{\nu}_e$

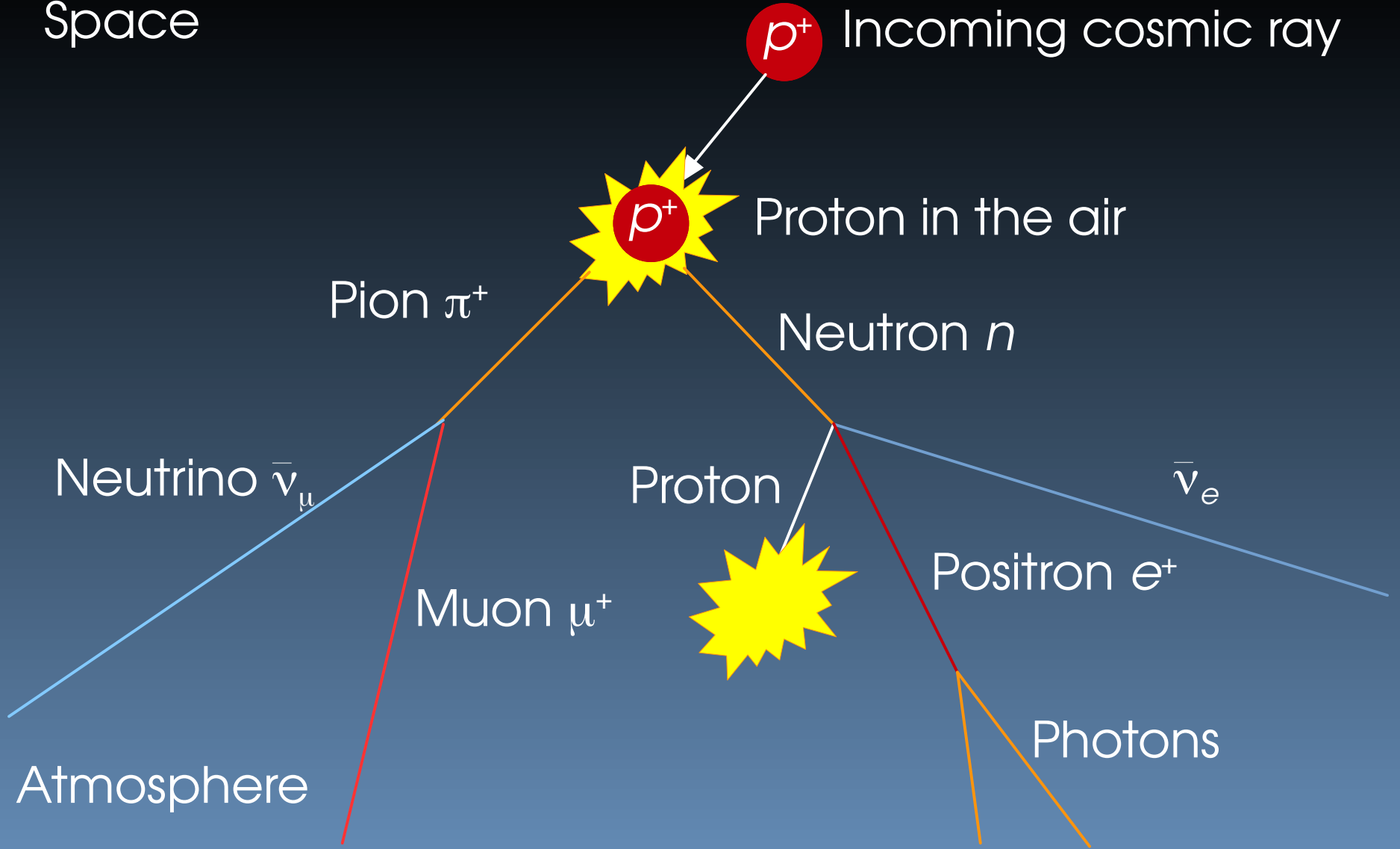
Muon  $\mu^+$



Positron  $e^+$

Atmosphere

Photons



Space

$p^+$  Incoming cosmic ray



$p^+$  Proton in the air

Pion  $\pi^+$

Neutron  $n$

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Proton

$\bar{\nu}_e$

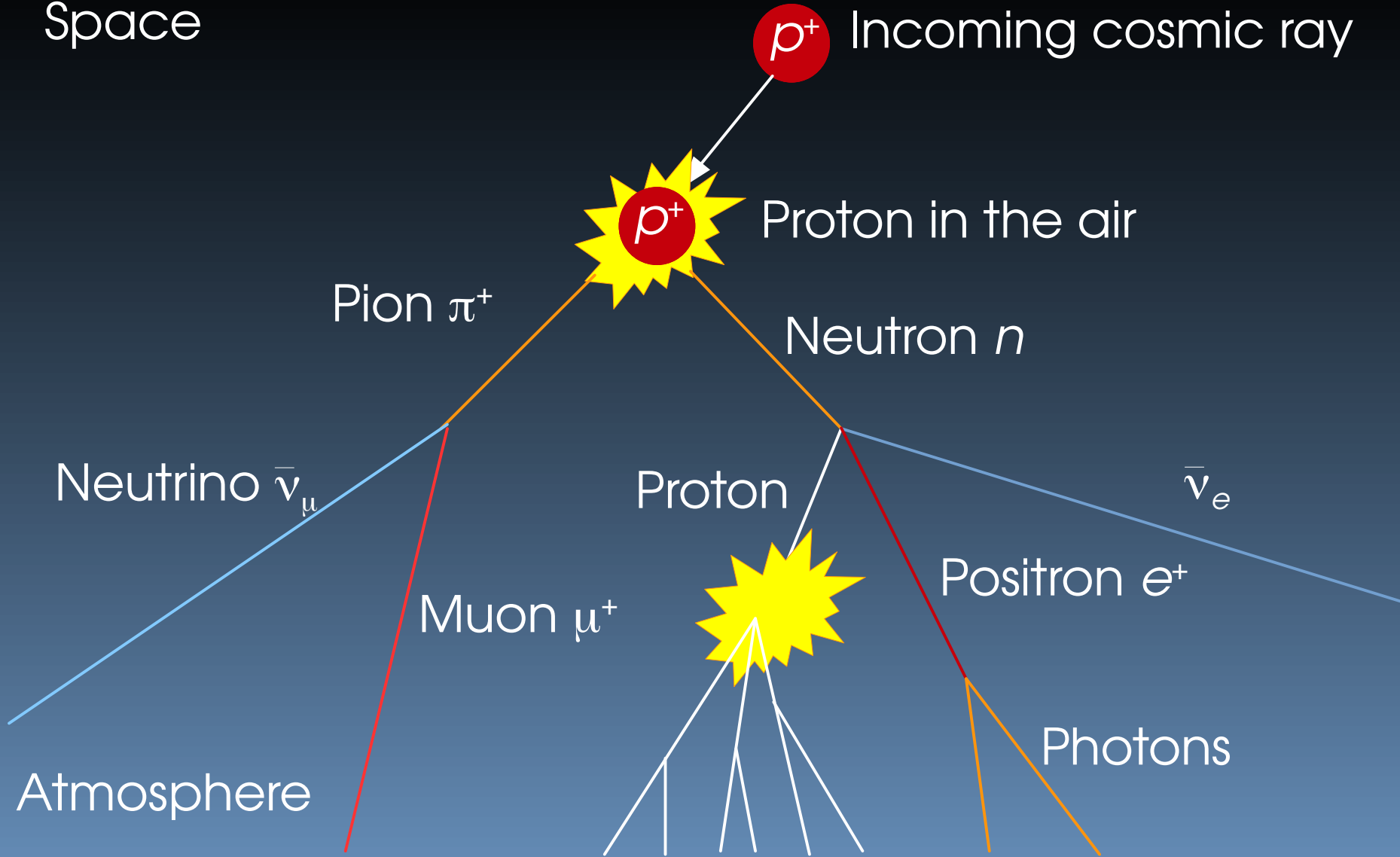
Muon  $\mu^+$



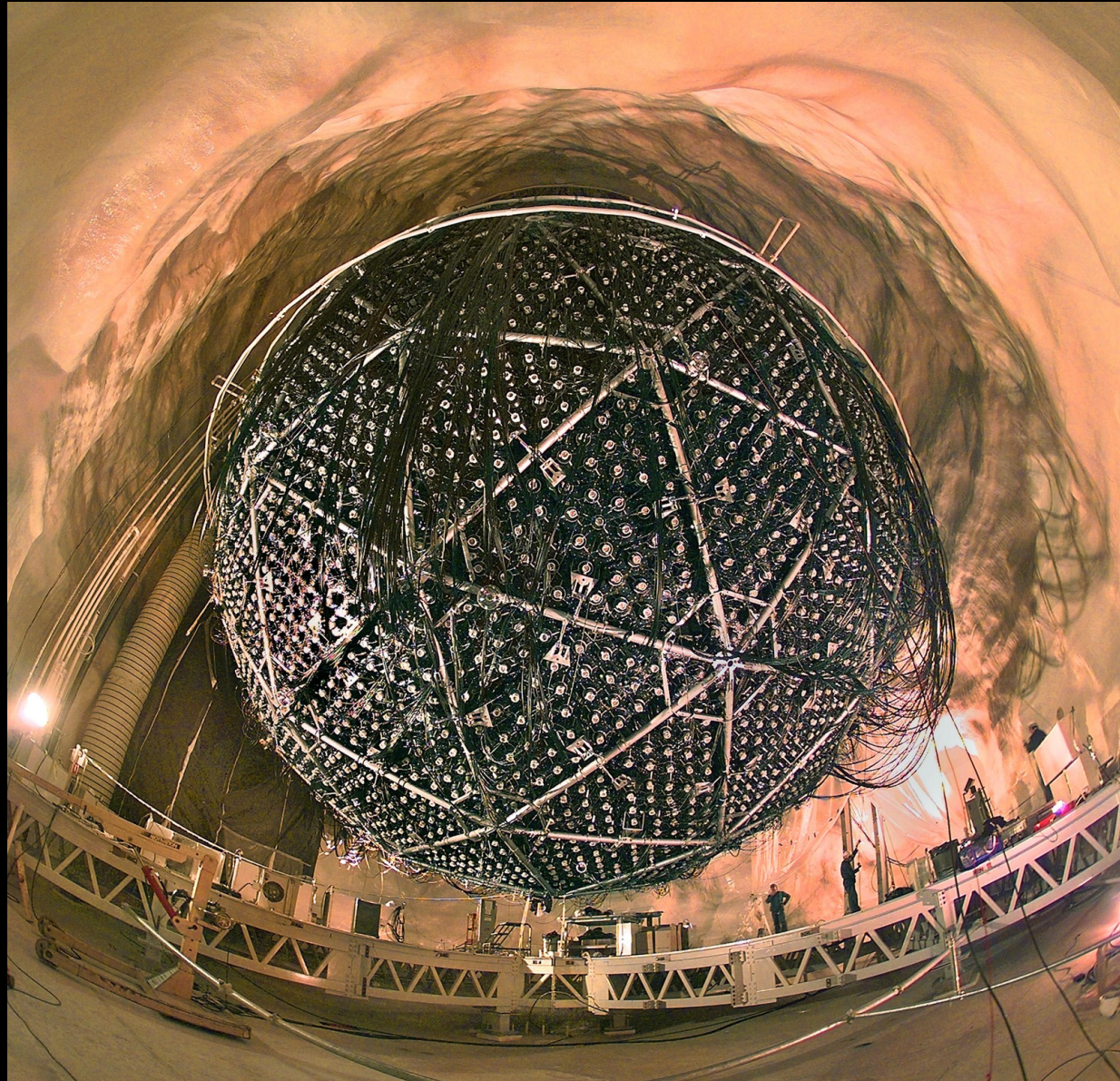
Positron  $e^+$

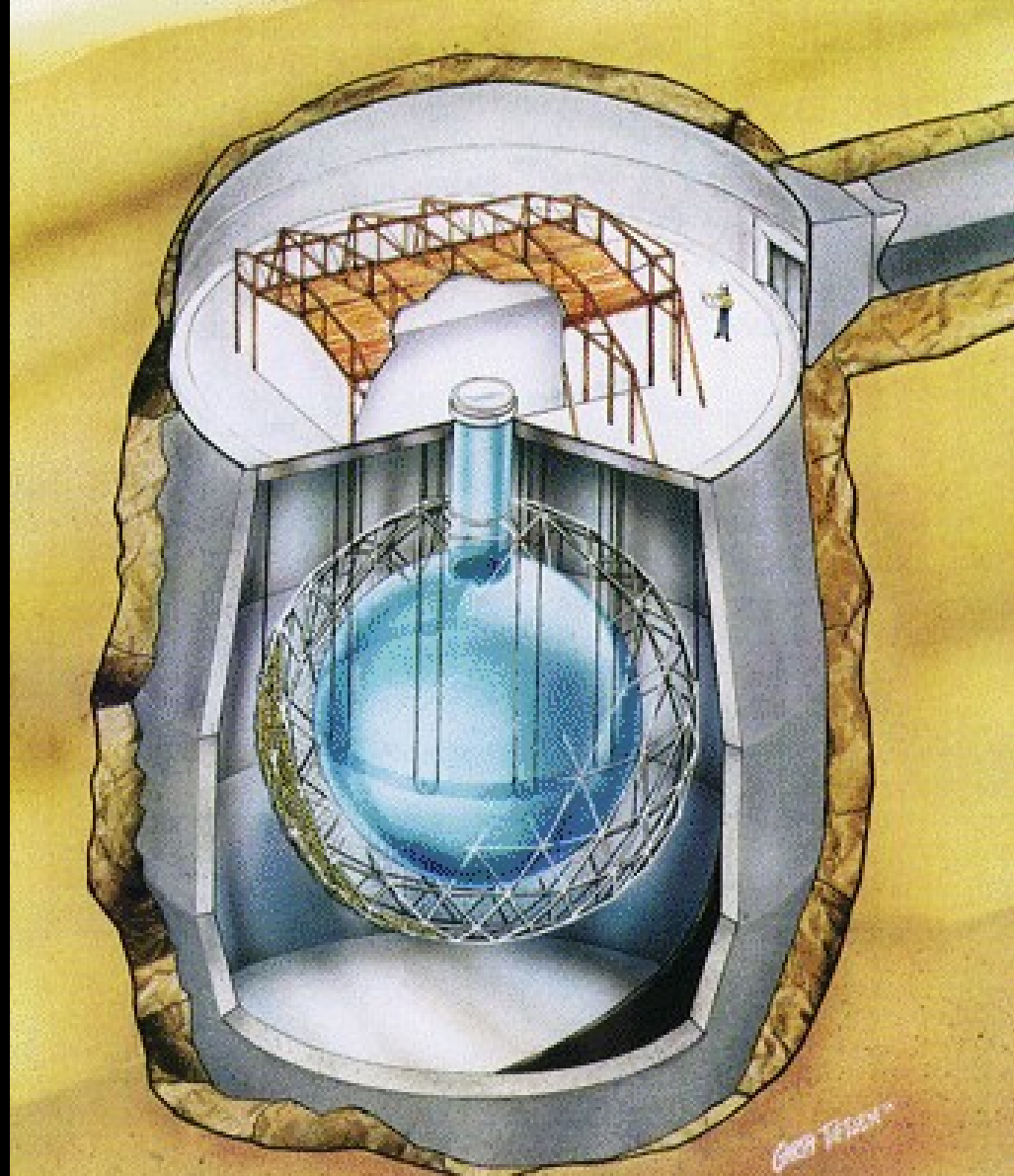
Atmosphere

Photons

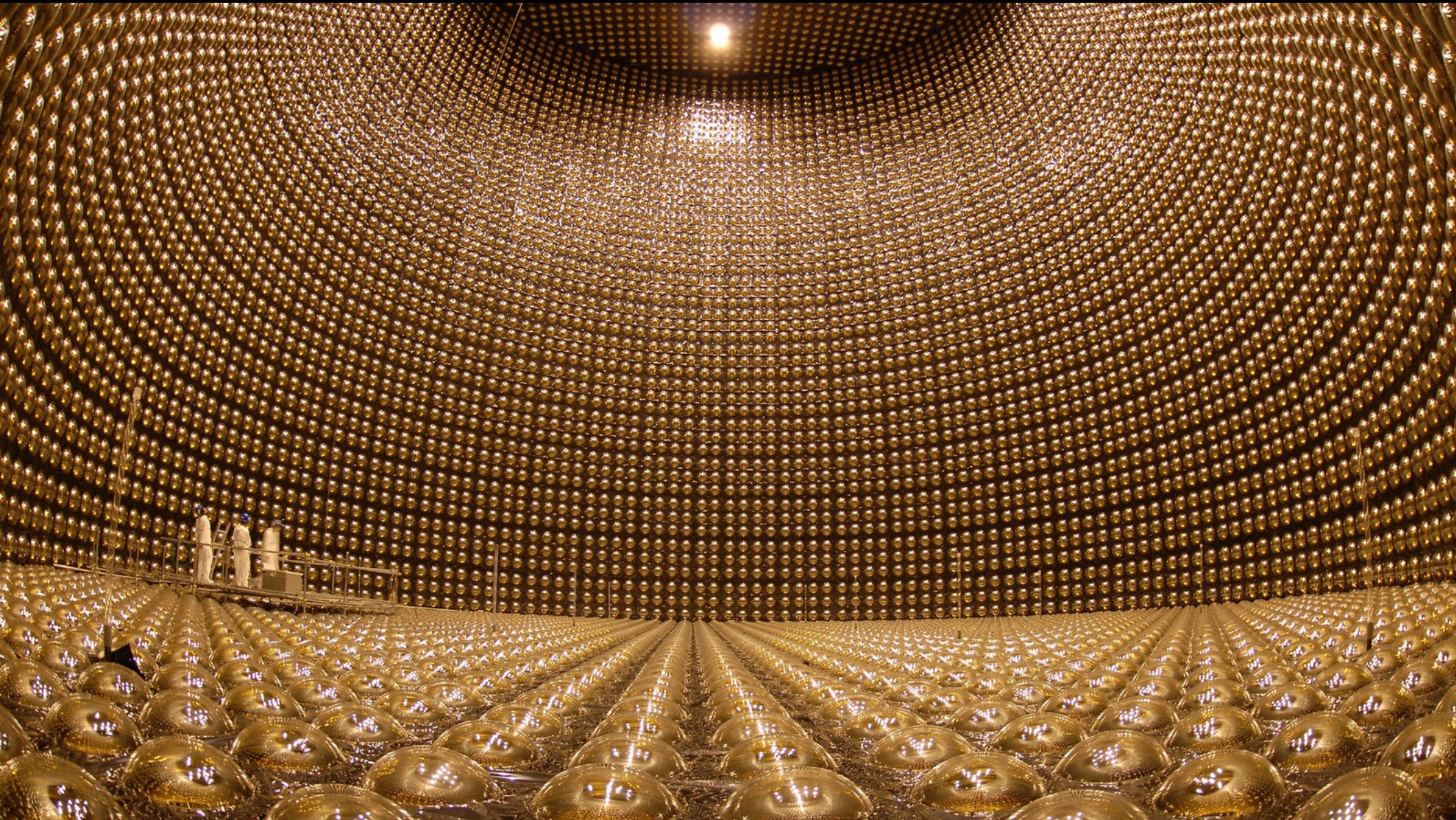


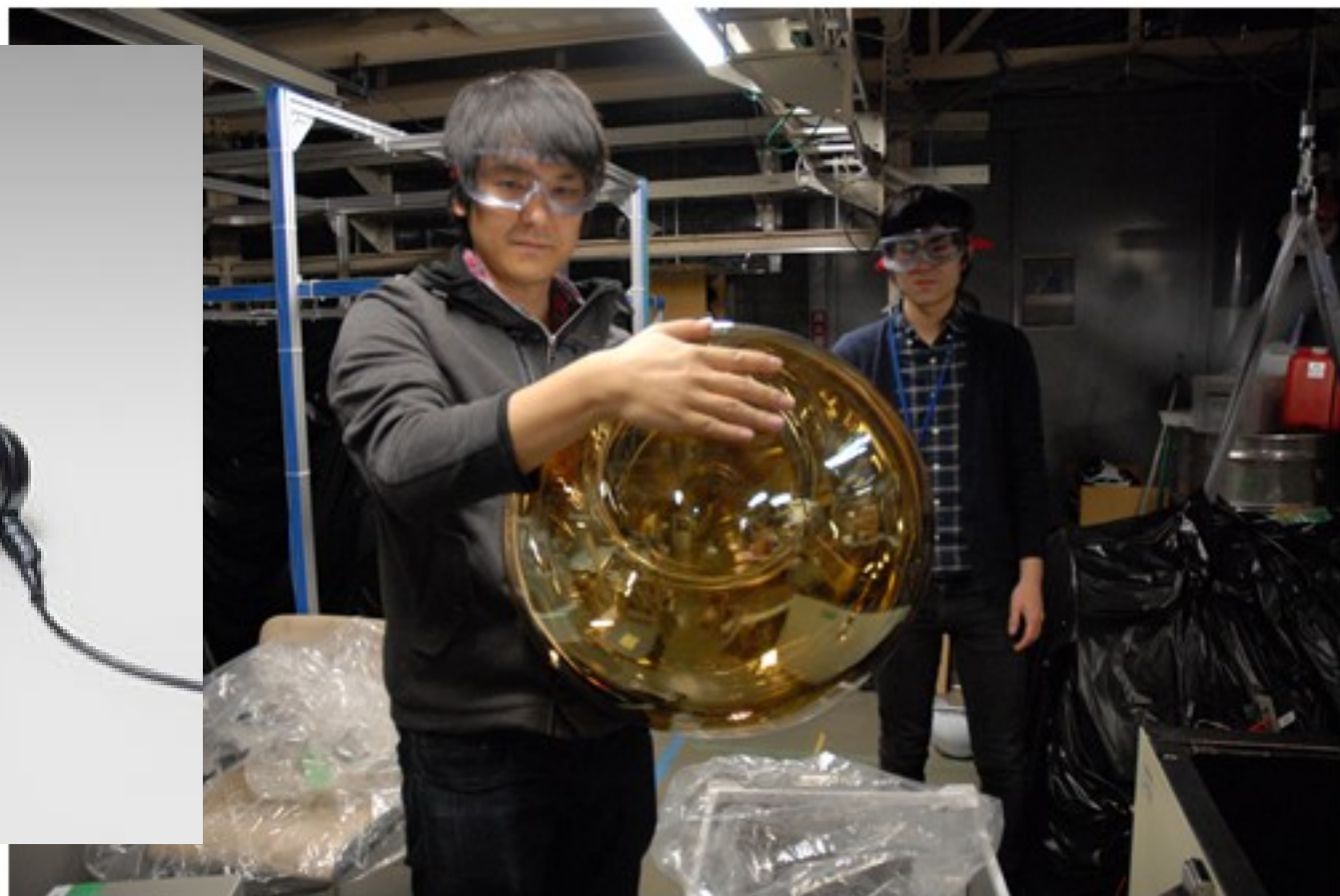


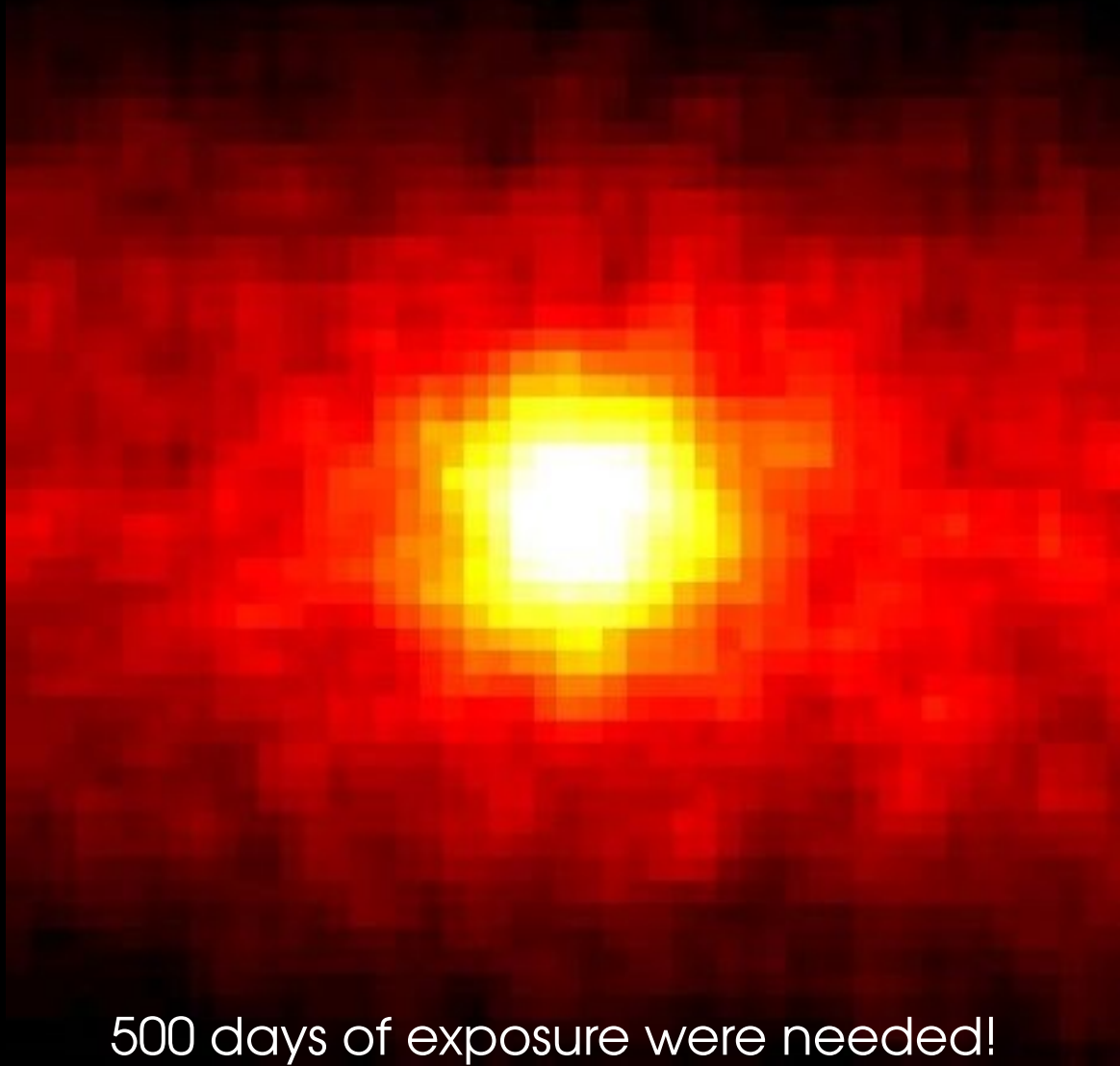




© 2004 TFCO



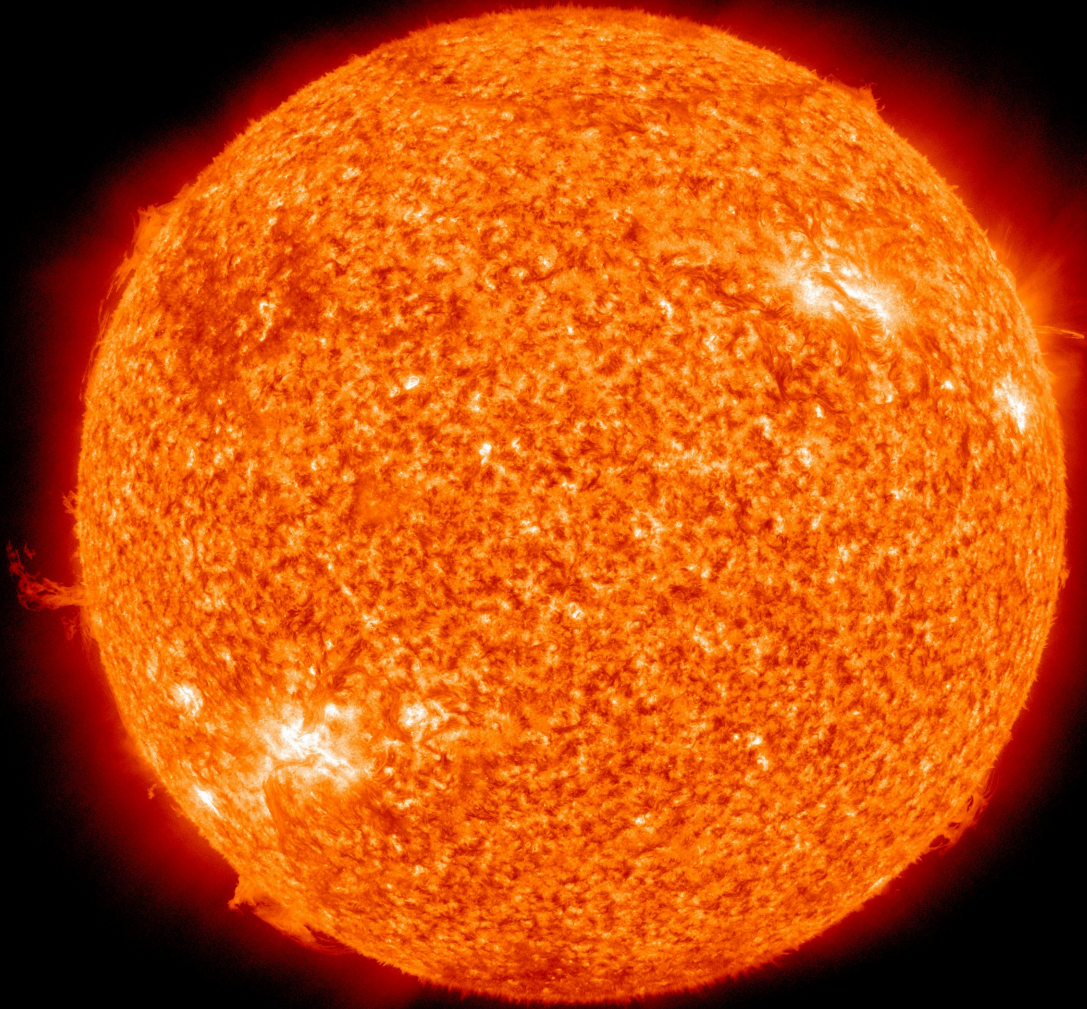




500 days of exposure were needed!

[Super-Kamiokande]

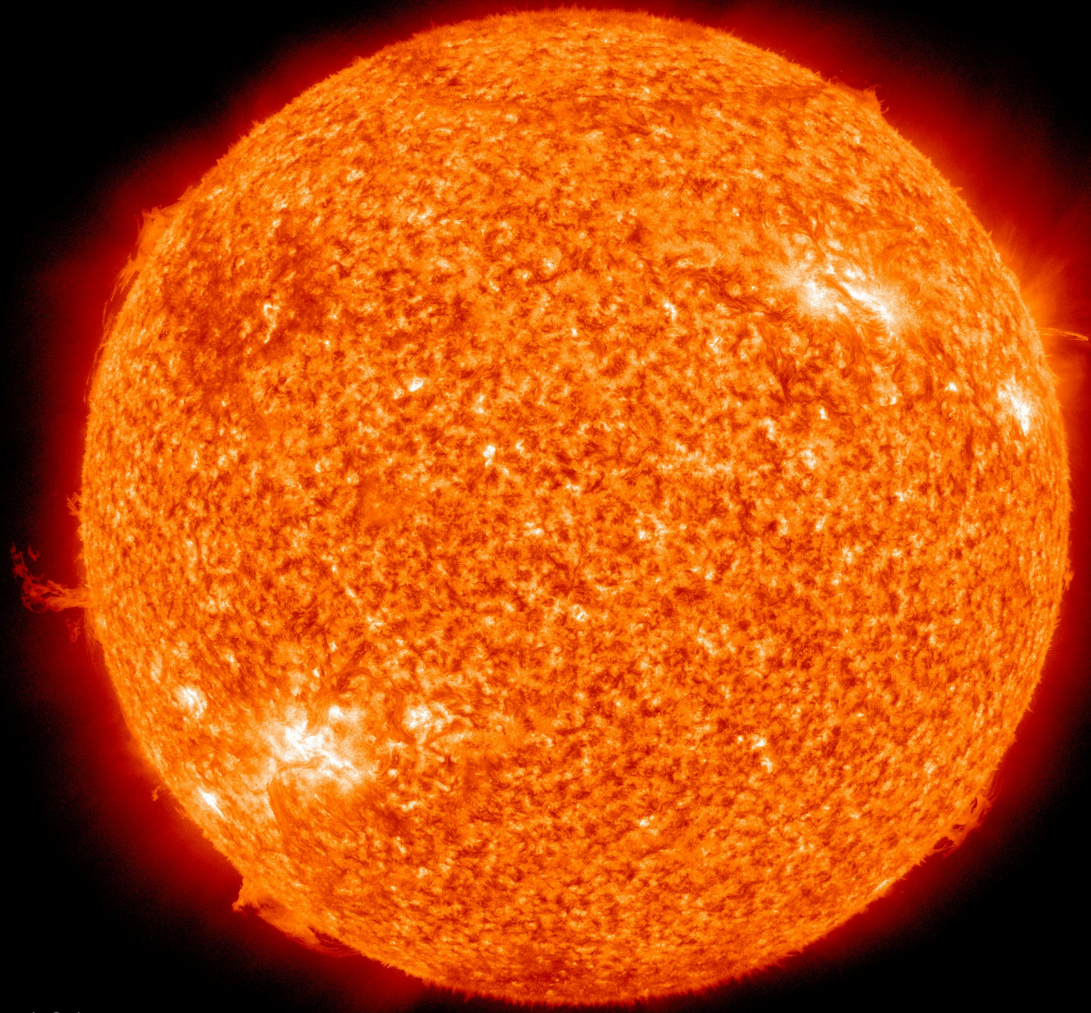
What did we learn?

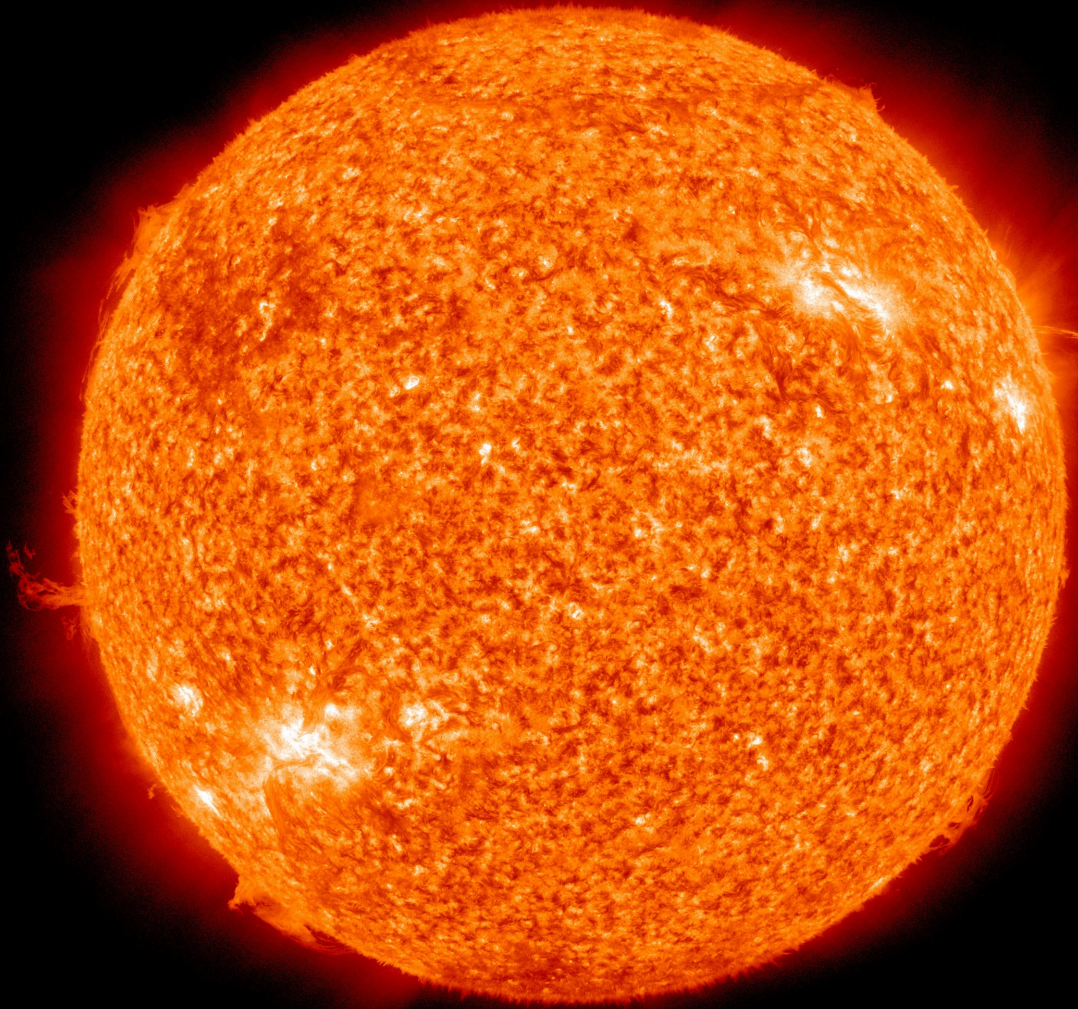


[NASA]

## What did we learn?

- ▶ Nuclear fusion powers the Sun





## What did we learn?

- ▶ Nuclear fusion powers the Sun
- ▶ The  $\nu$  flux strongly depends on the core temperature  $T$ :

$$\nu \text{ flux} \propto T^{24}$$

By detecting  $\nu$ , we found

$$T \approx 15 \text{ million K}$$

## 2002 Nobel Prize in Physics

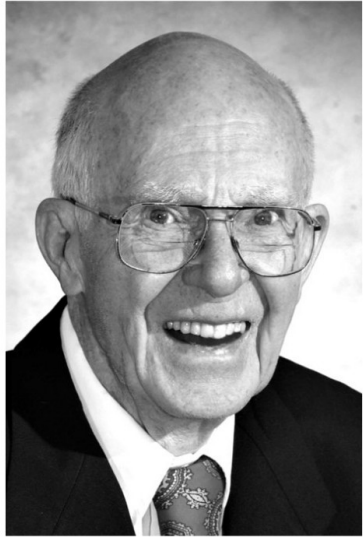


Photo from the Nobel Foundation archive.

Raymond Davis Jr.

Prize share: 1/4



Photo from the Nobel Foundation archive.

Masatoshi Koshihara

Prize share: 1/4



Photo from the Nobel Foundation archive.

Riccardo Giacconi

Prize share: 1/2

“... one half jointly to Raymond Davis Jr. and Masatoshi Koshihara *‘for pioneering contributions to astrophysics, in particular for the detection of cosmic neutrinos’* ”

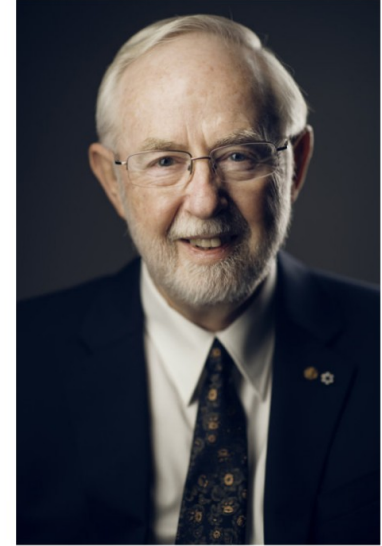
## 2015 Nobel Prize in Physics



© Nobel Media AB. Photo: A. Mahmoud

Takaaki Kajita

Prize share: 1/2



© Nobel Media AB. Photo: A. Mahmoud

Arthur B. McDonald

Prize share: 1/2

“... *‘for the discovery of neutrino oscillations, which shows that neutrinos have mass.’* ”

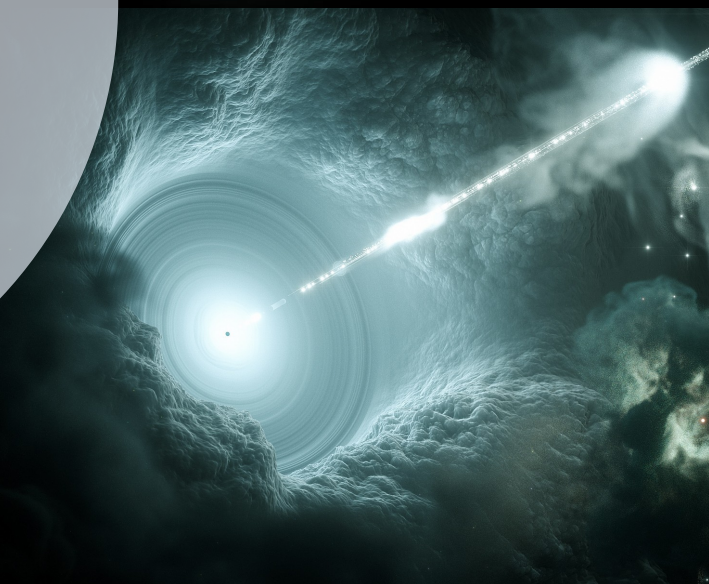
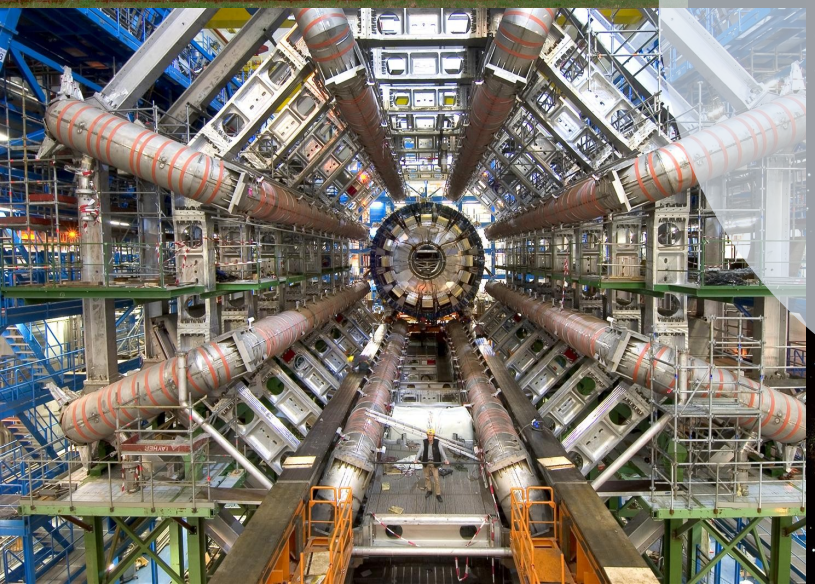
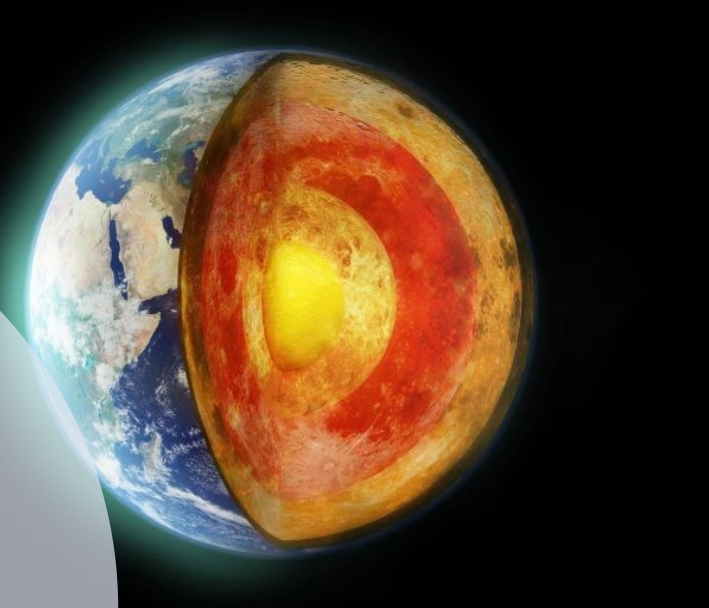
# Congratulations to Dr. Takaaki Kajita

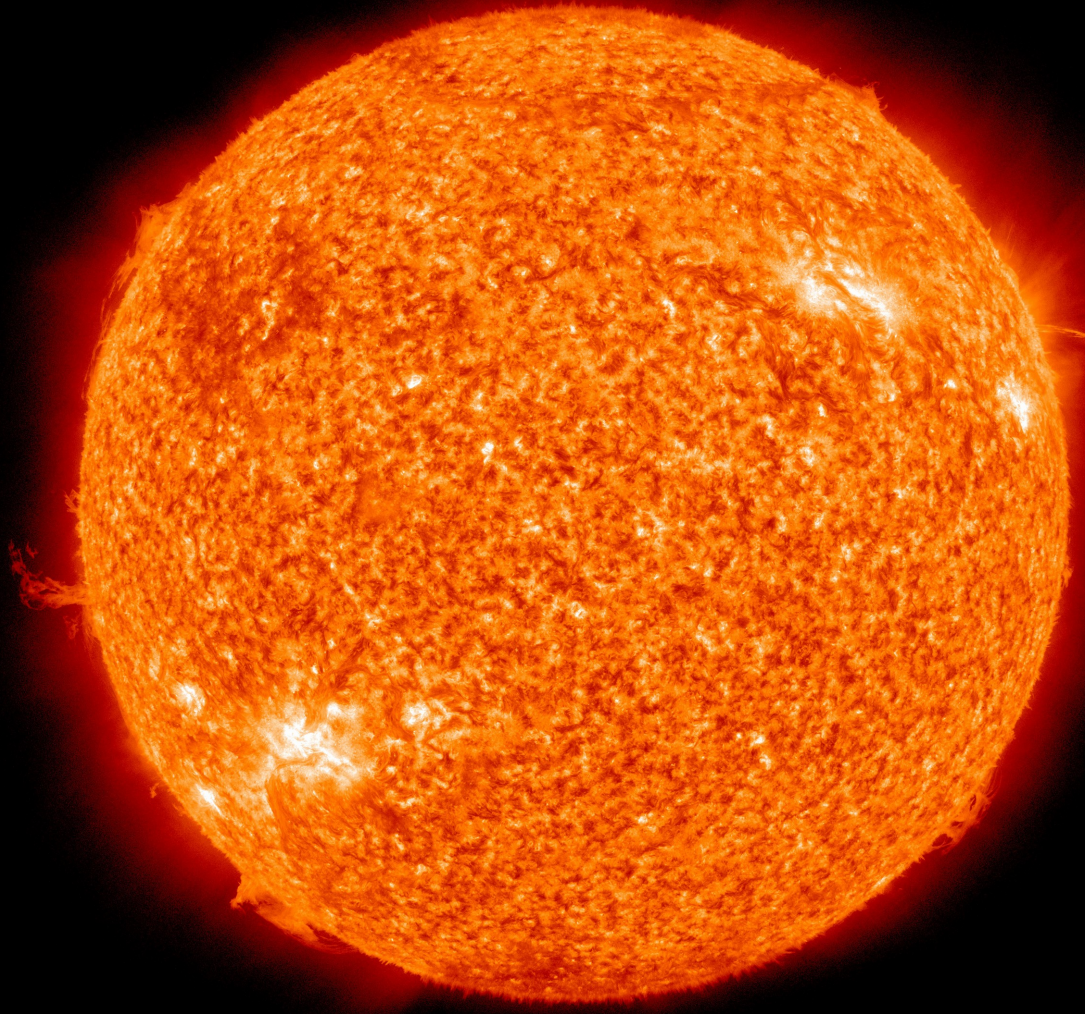
Director of the Institute for Cosmic Ray Research Kashima Campus at The University of Tokyo for being awarded The Nobel Prize in Physics 2011

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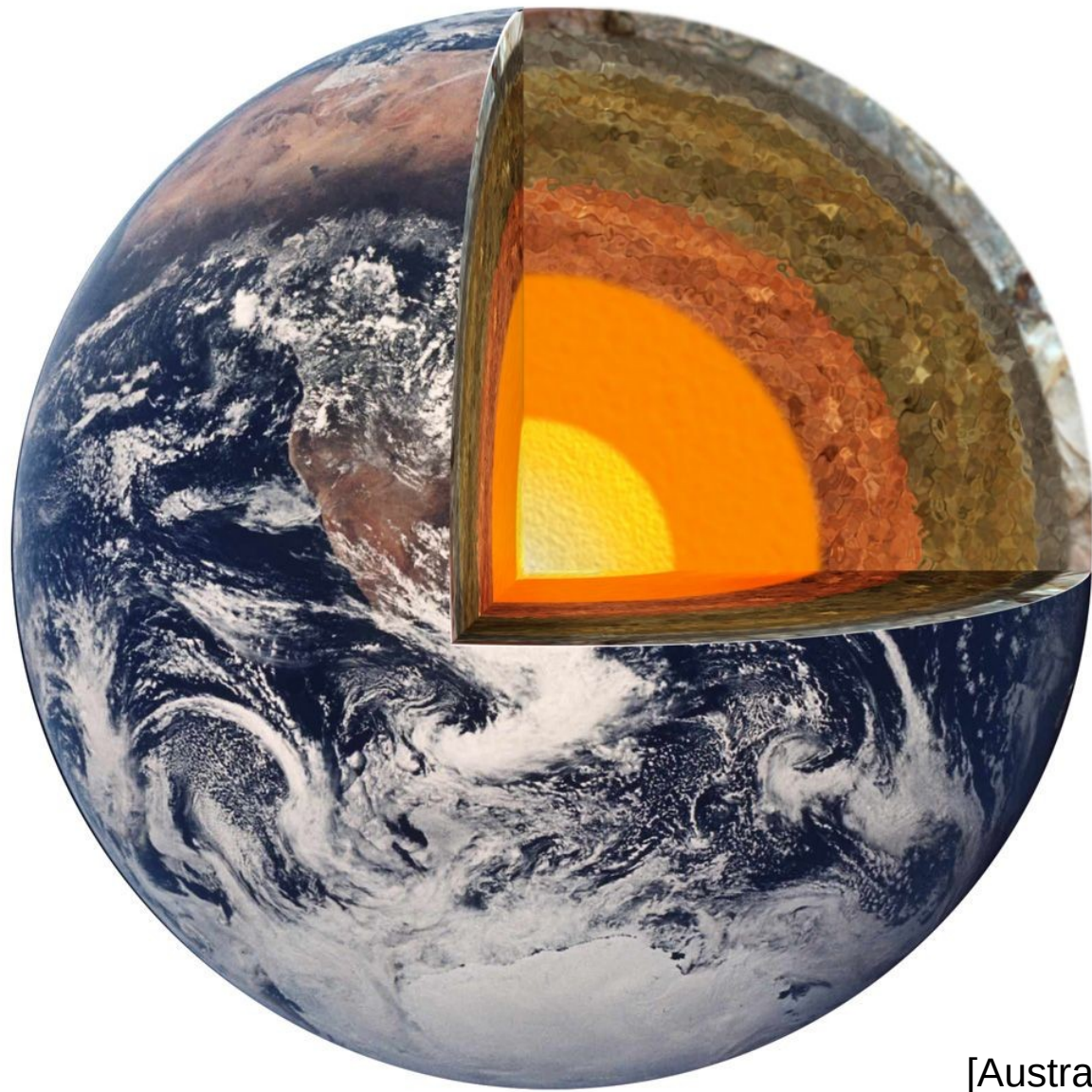




[NASA]



[Avda]



[Australian National University]



[NASA, ESA]



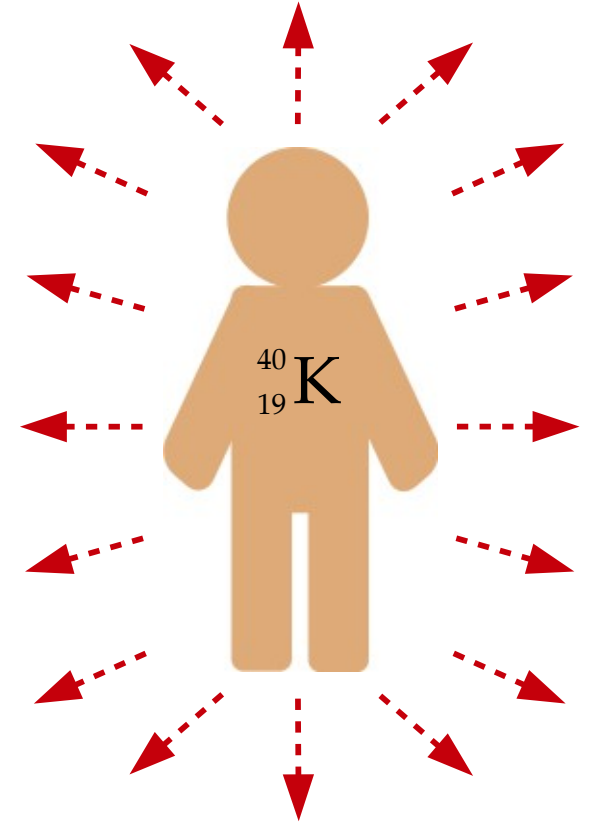
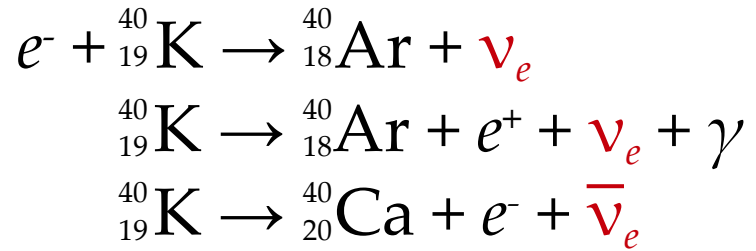
[NASA, ESO]

# Neutrinos are everywhere: even *you* make them!

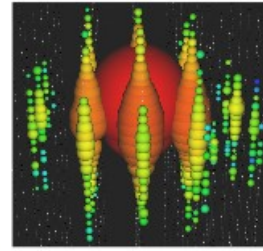
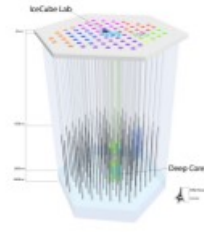
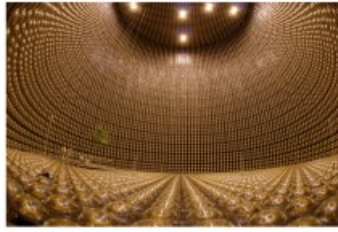


Some of the potassium  
in bananas is radioactive

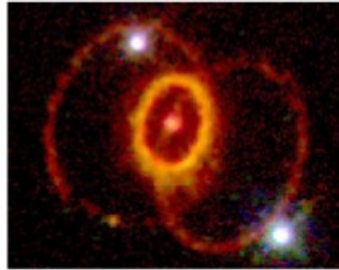
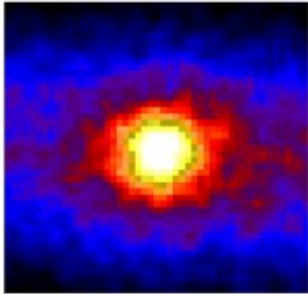
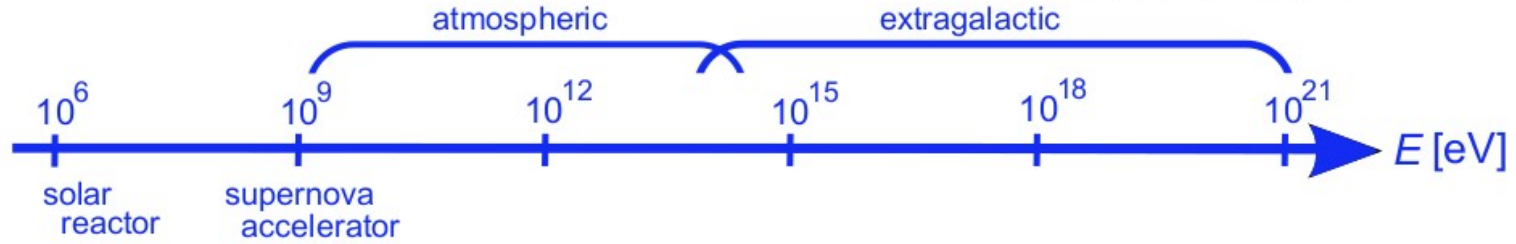
Potassium-40 has a half-life  
of ~ 1 billion years:



4000+ neutrinos emitted each second by a 70-kg person



2013+

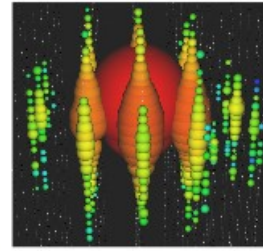
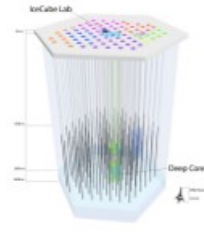
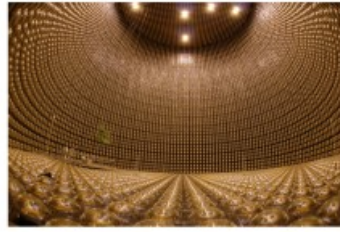


$$10^{12} \text{ eV} \rightarrow$$

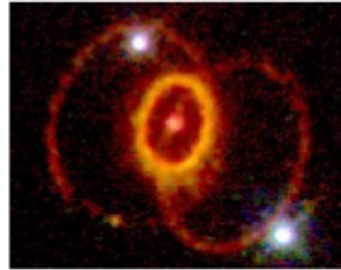
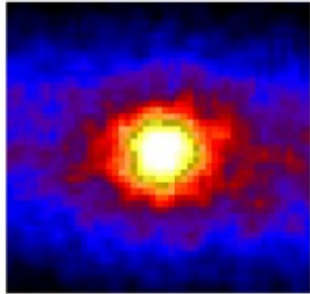
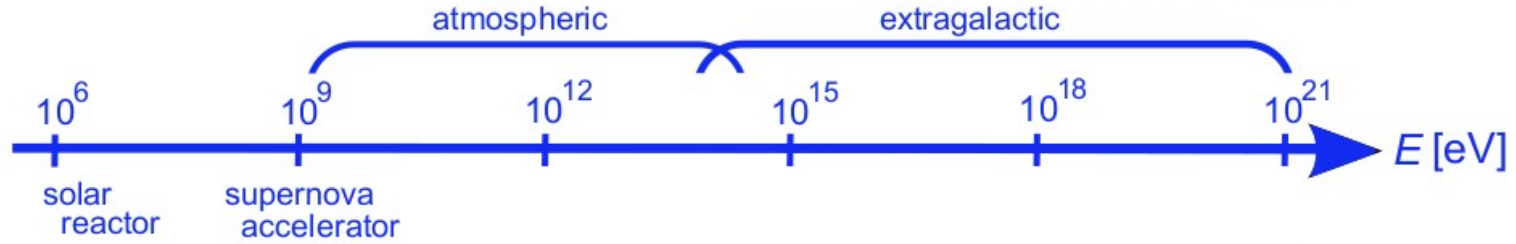


$$6 \times 10^{20} \text{ eV} \rightarrow$$





2013+

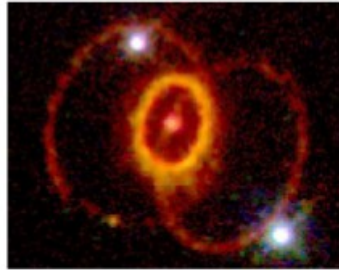
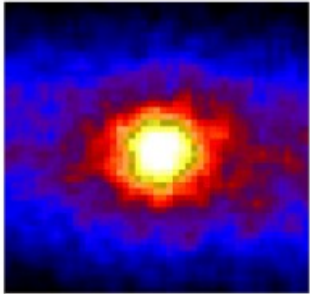
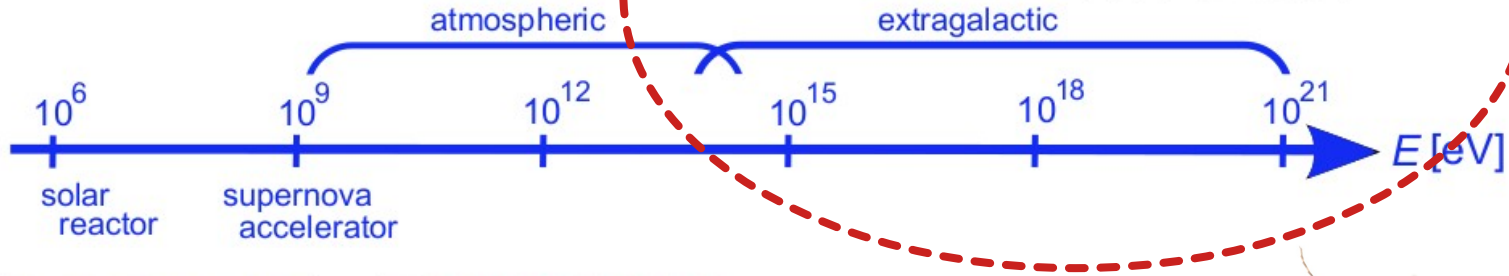
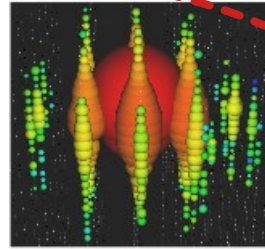
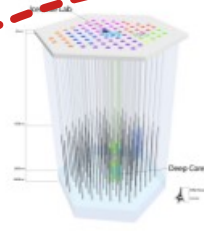


$$10^{12} \text{ eV} \rightarrow$$



$$6 \times 10^{20} \text{ eV} \rightarrow$$



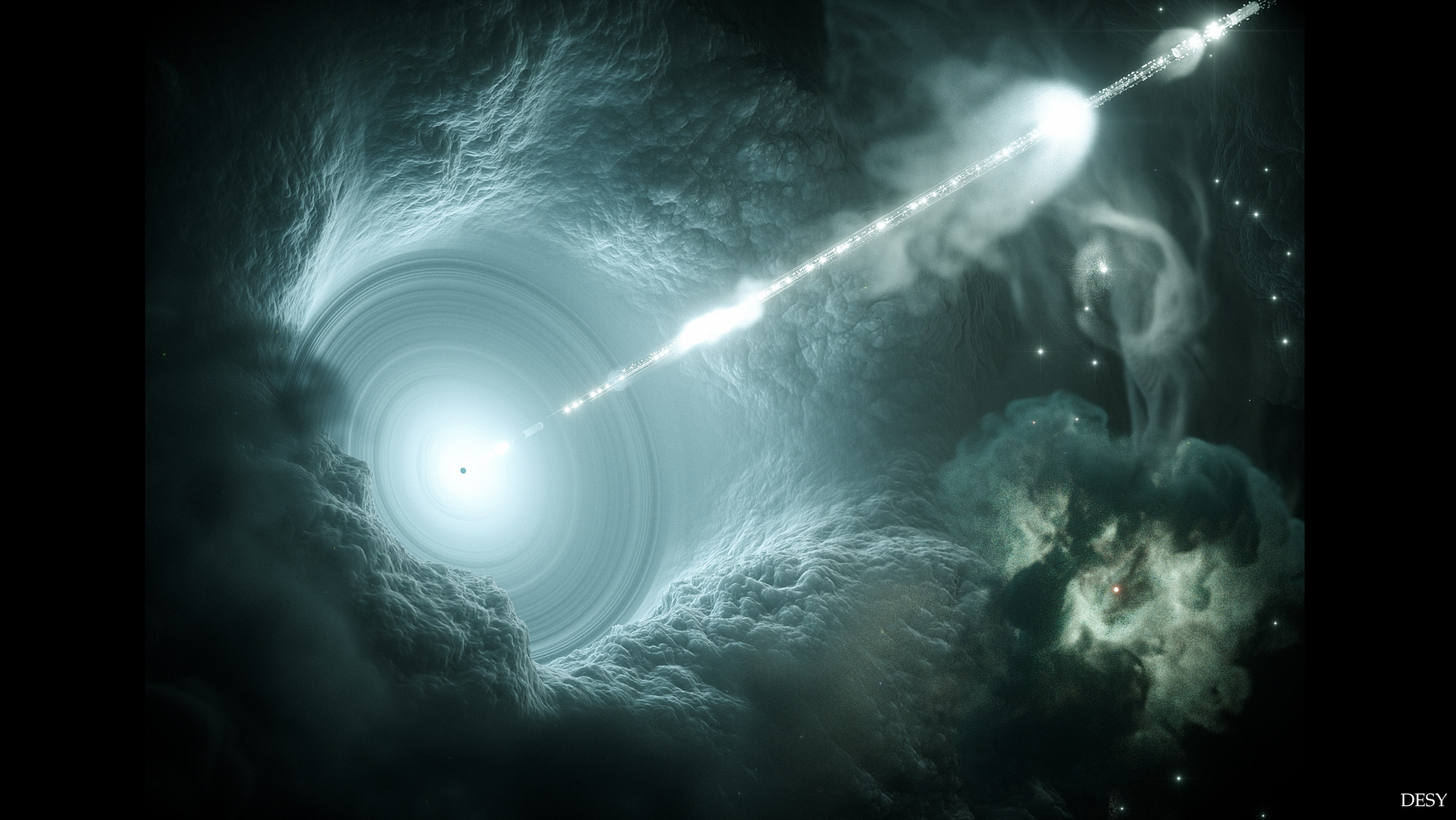


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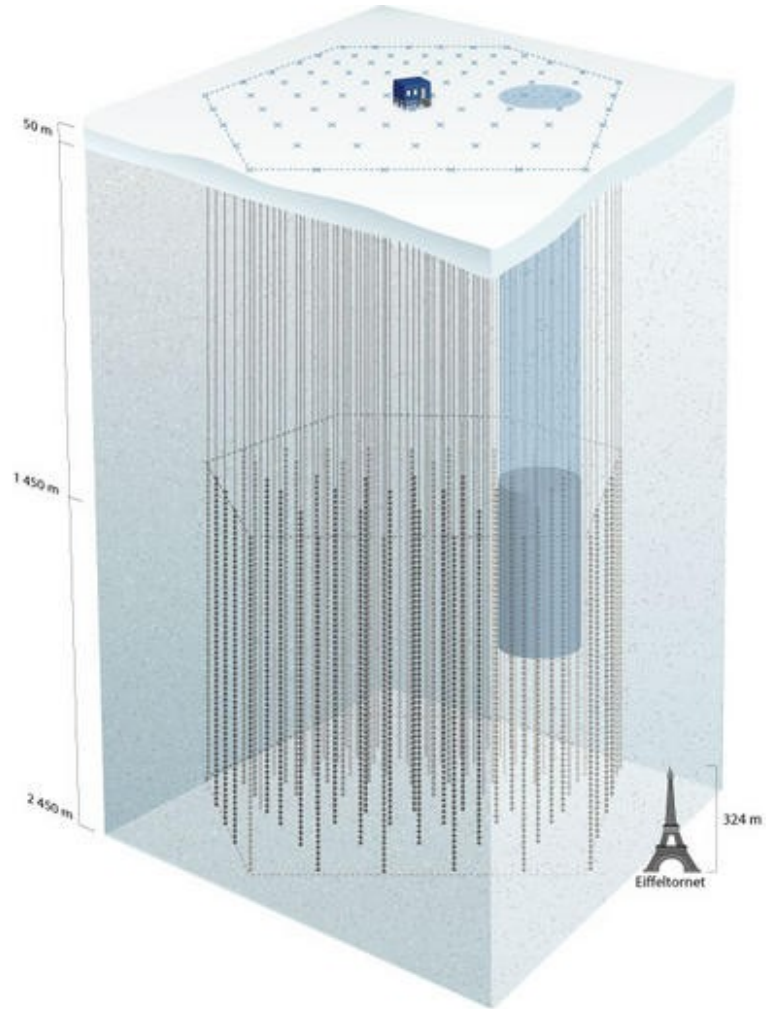
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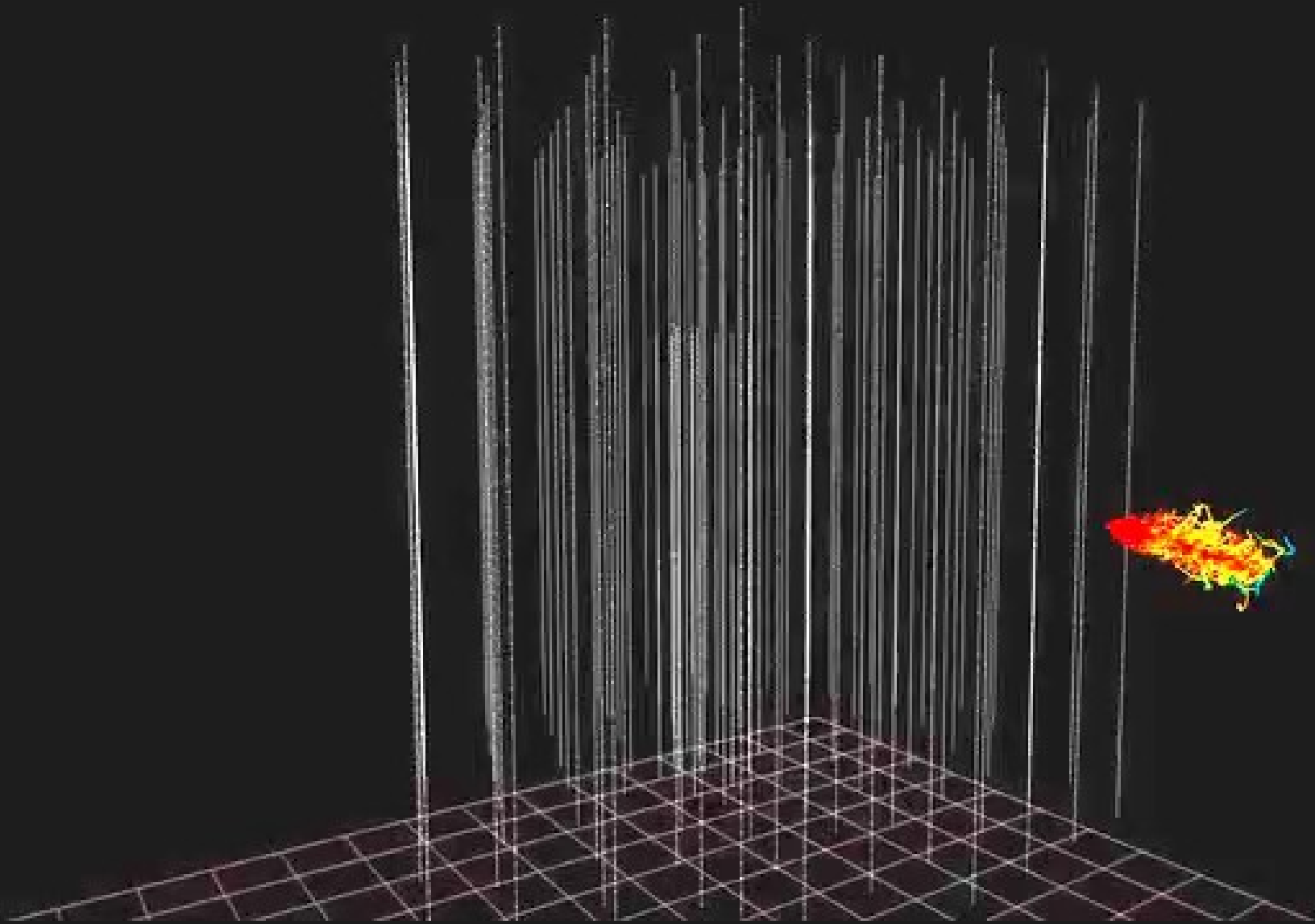


# IceCube: The largest neutrino telescope

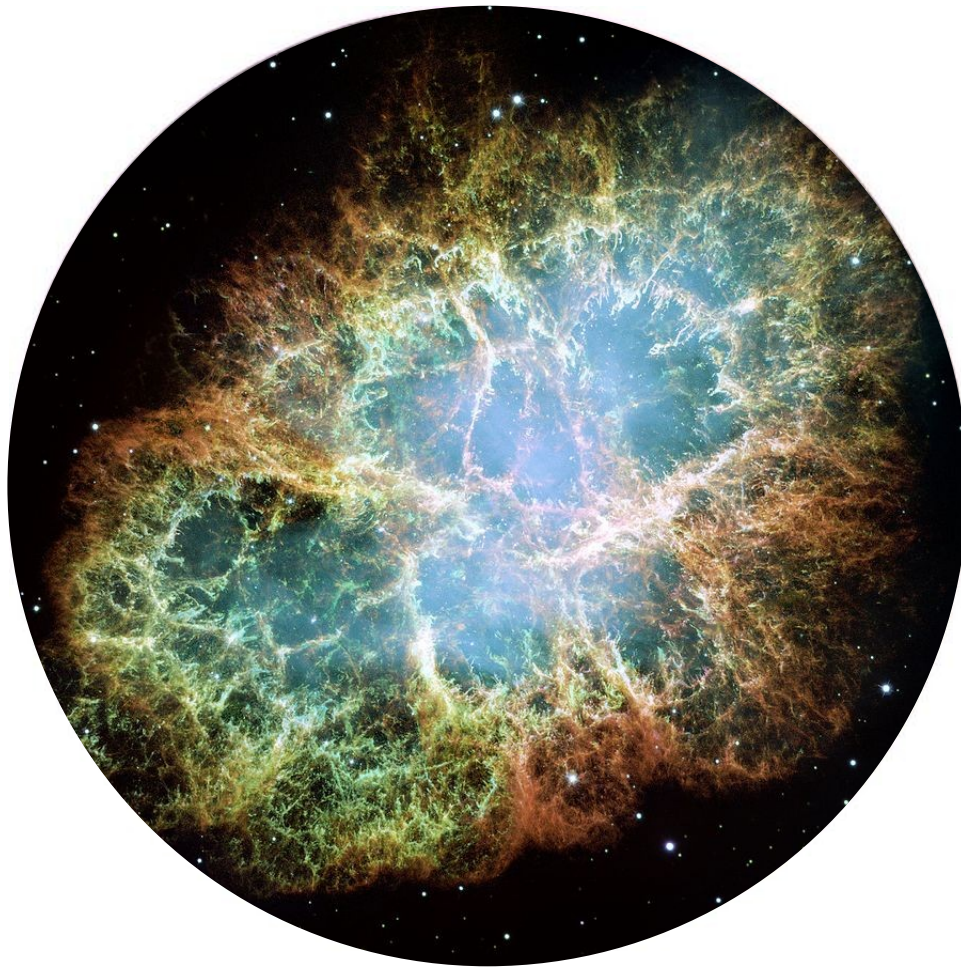


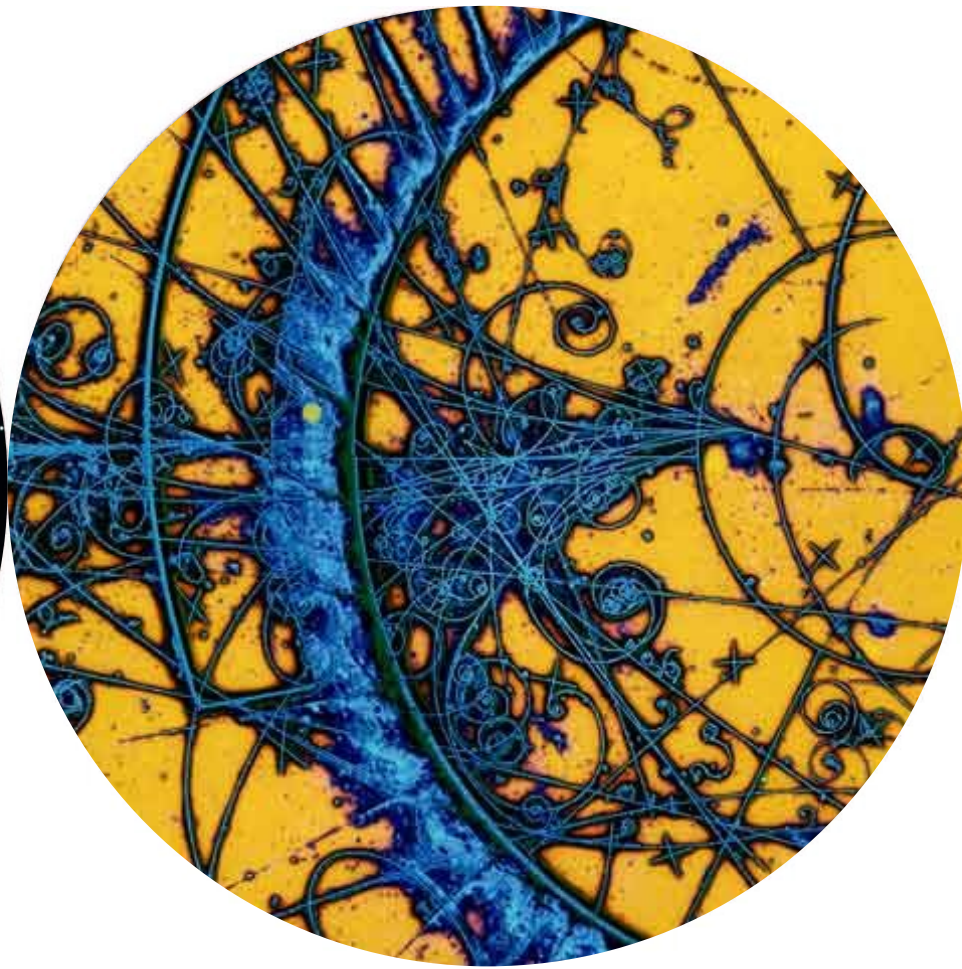
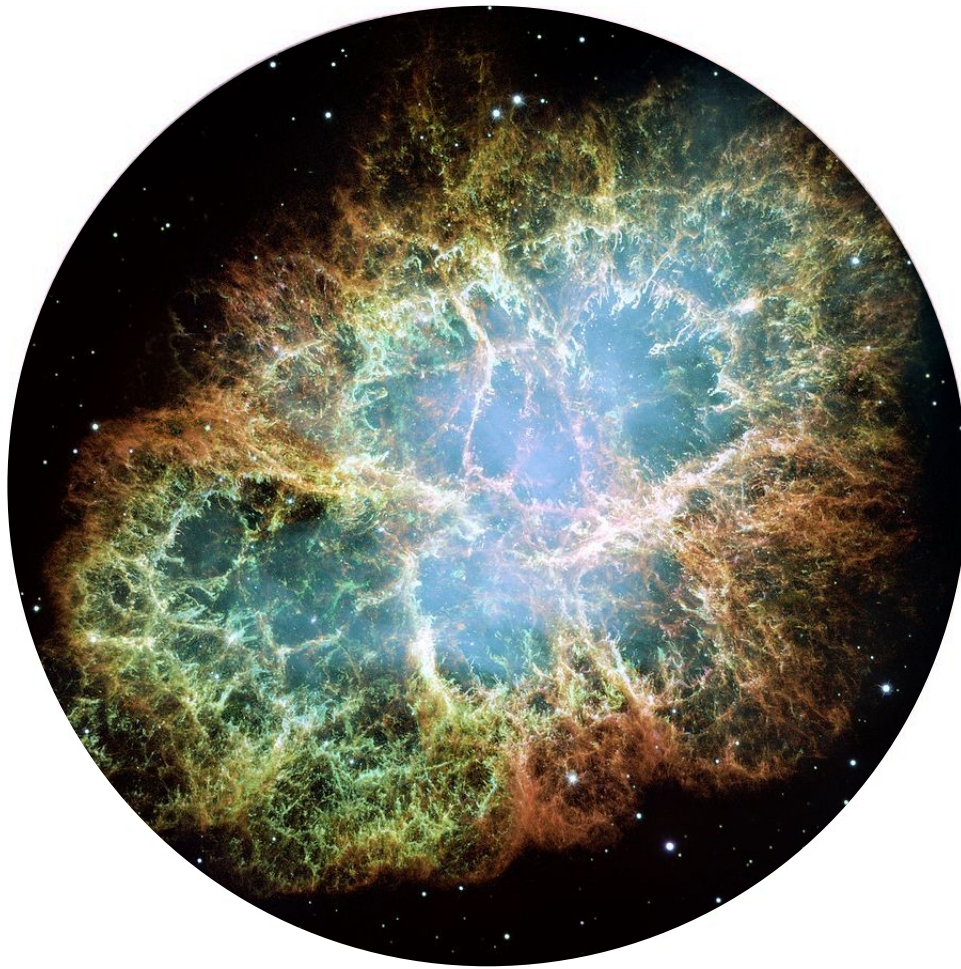
- ▶  $\text{Km}^3$  in-ice Cherenkov detector in Antarctica
- ▶ > 5000 PMTs at 1.5–2.5 km of depth



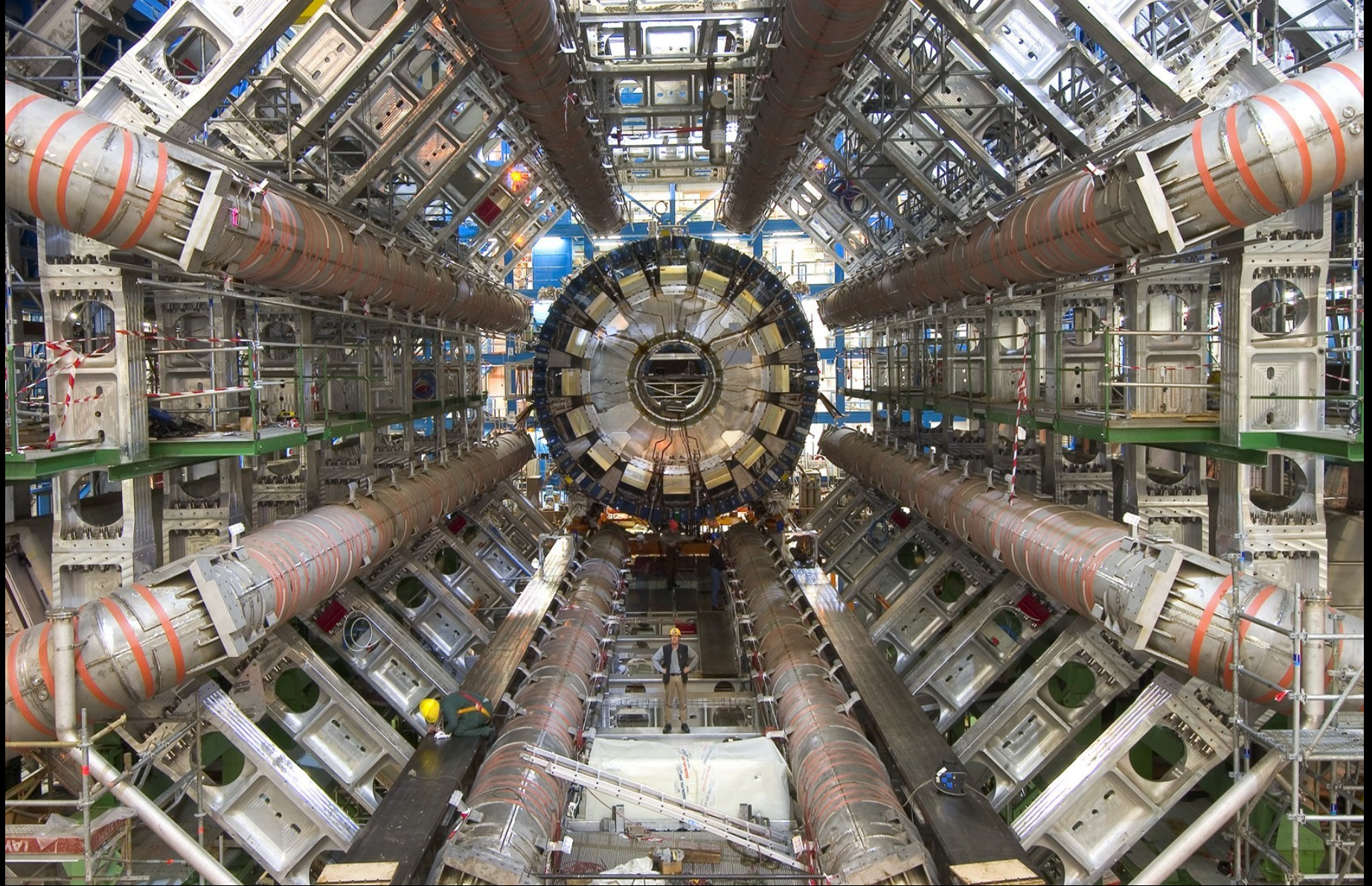


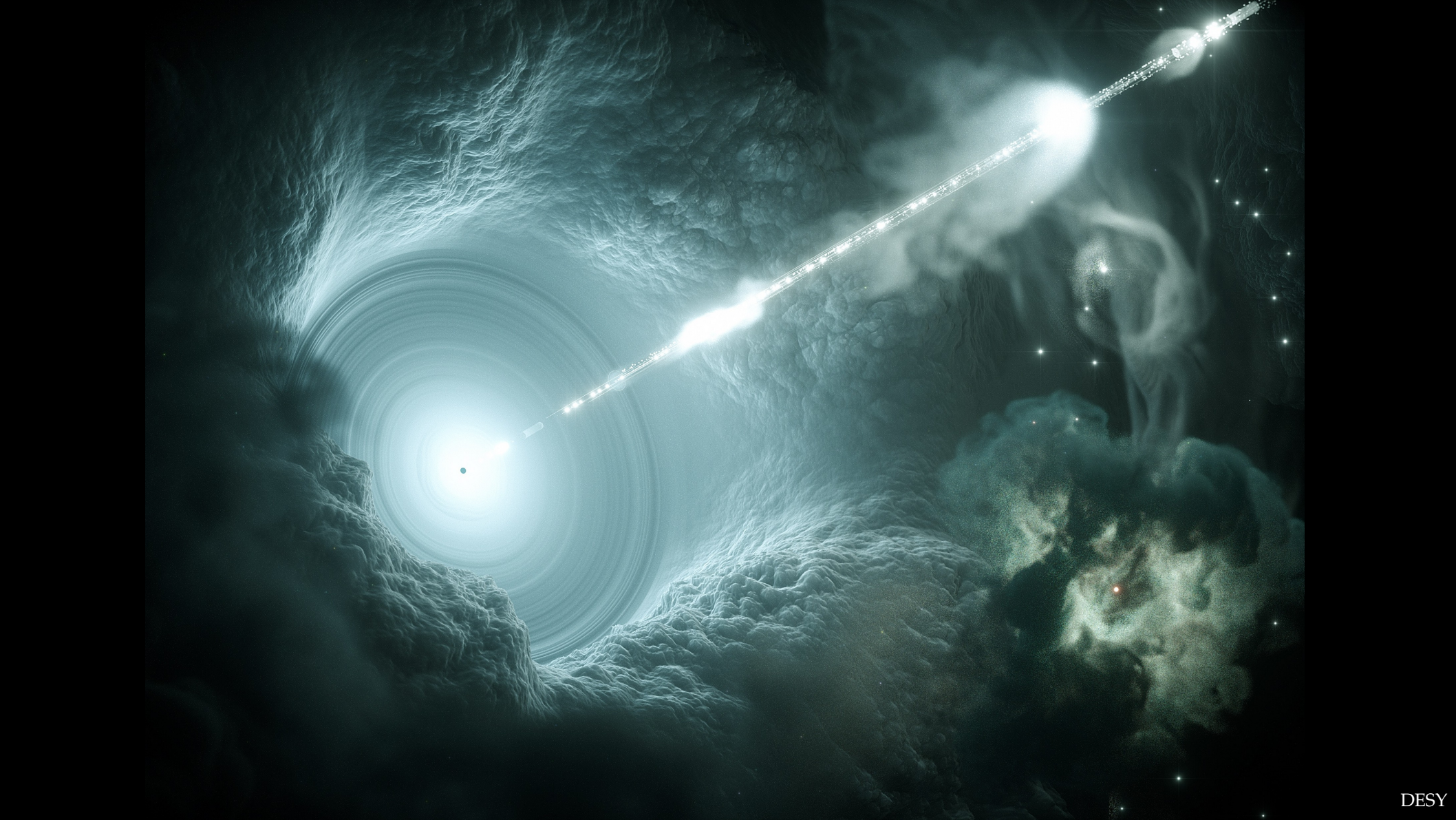


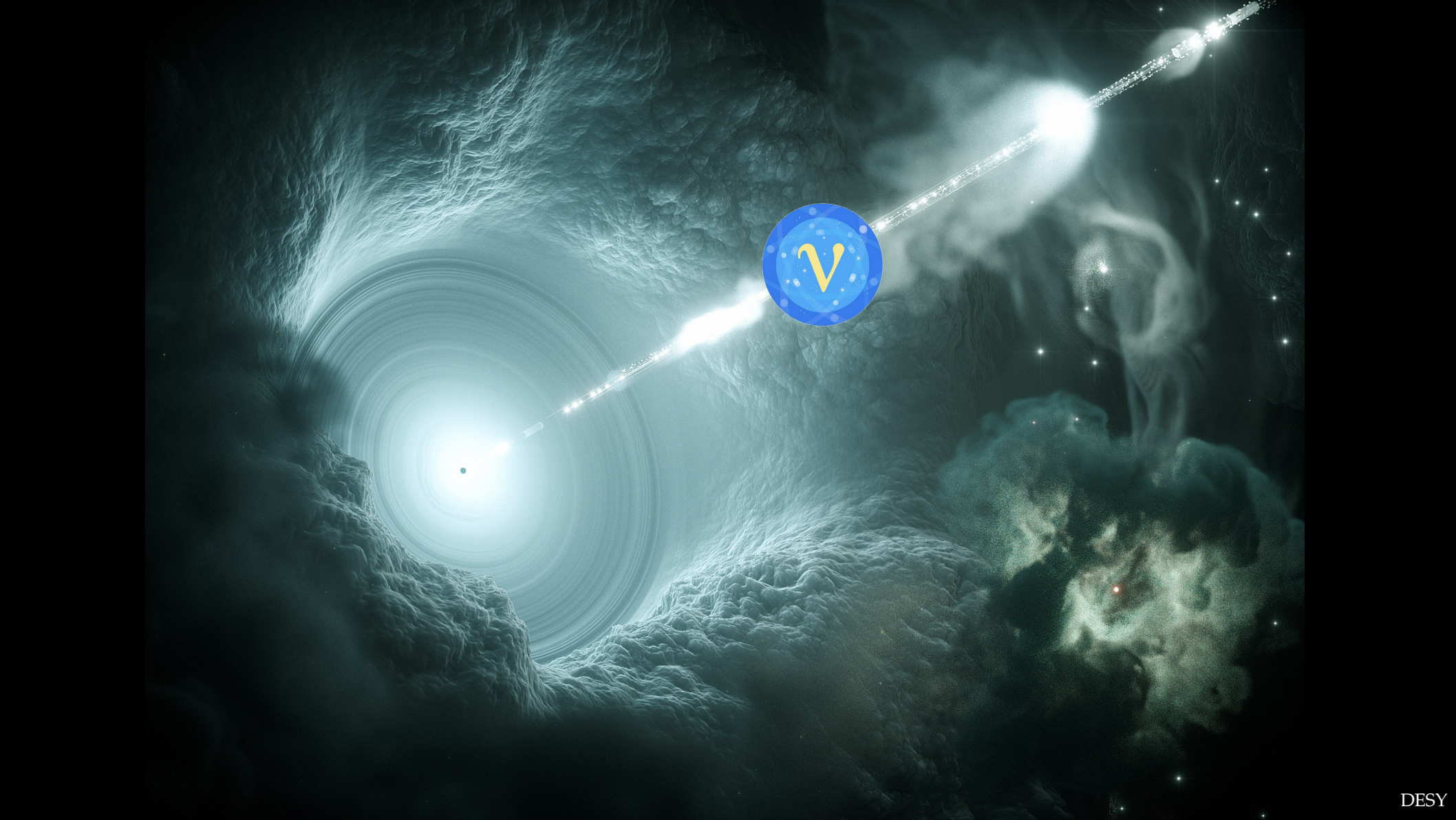

















2021 (*we are here*):  
TeV–PeV  $\nu$   
discovered  
First possible sources



2020s (*we are getting there*):  
More source candidates  
Characterize the  $\nu$  flux precisely

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> 2040:  
????

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# Undergrad & MSc: PUCP, 5+3 yr



- ▶ High Energy Physics Group
- ▶ 4 papers (+ proceedings) – key to apply to PhD
- ▶ 7 local & international conferences and schools
- ▶ 6–12-month research visits during MSc
  - ▶ Meet future colleagues
  - ▶ Meet future mentors
  - ▶ Learn the state of the art
  - ▶ Be exposed to other styles

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## PhD: Würzburg & DESY, Germany, ~3 yr

- ▶ Had a young, motivated PI as supervisor
- ▶ Main work: modeling high-energy astrophysical particle production
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# Working with experimental collaborations

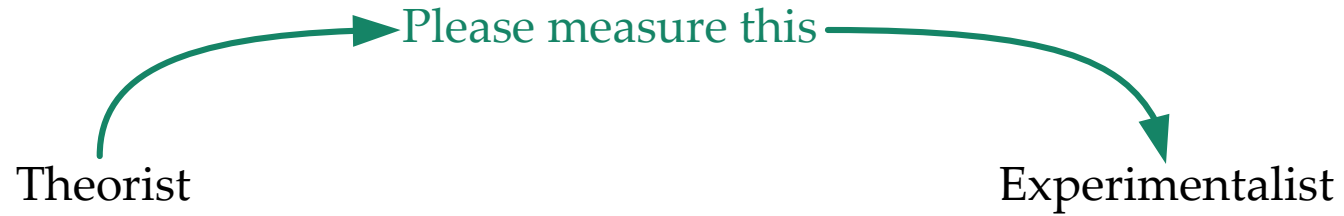
I'm a theorist, but I work in close proximity to experimental collaborations

Theorist

Experimentalist

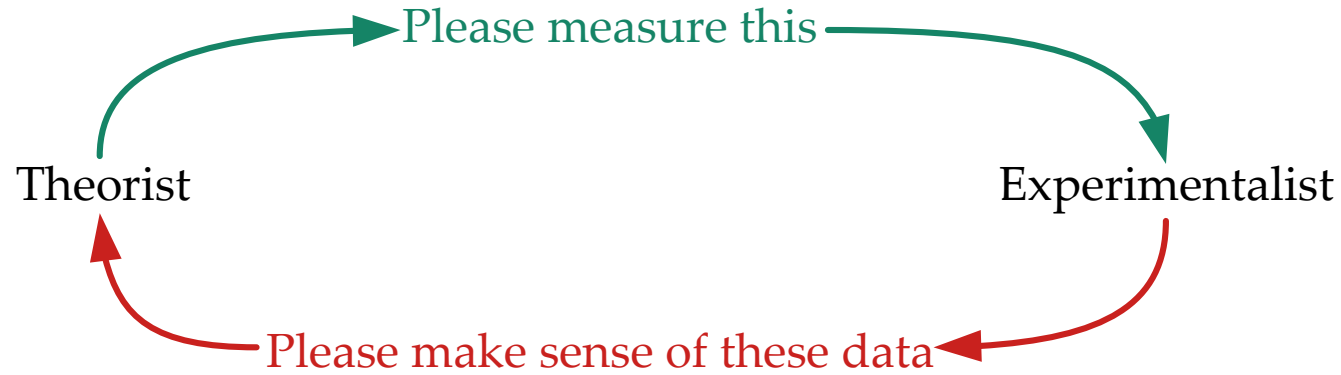
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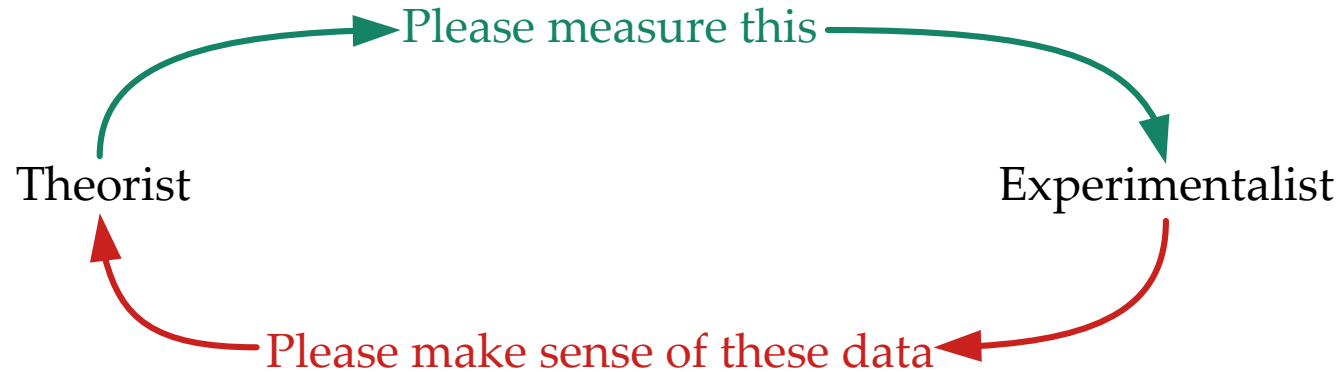
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Plan the next generation of large-scale neutrino telescopes for the coming 10–20 years

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## Postdoc #1: Ohio State, USA, 3 yr

- ▶ Center for Cosmology and Astroparticle Physics
- ▶ **Transitioned to more particle physics**
- ▶ Started work with experimental collaborations
- ▶ Started refereeing papers: improves writing
- ▶ Gave lots of talks (~40 significant ones!)
- ▶ Started building my own researcher identity: experimentally minded theoretical perspective



*Becoming an independent researcher*

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Today

*Becoming an independent researcher*

# Today: Faculty position

- ▶ Assistant Professor, Niels Bohr Institute, U. of Copenhagen, Denmark
- ▶ Starting research grant from the Danish Villum Fonden (~1.5M USD)
- ▶ Building my own research group (students + postdocs)
- ▶ My time is divided between my students, my own research, and grant applications
- ▶ Key skill: Time management (too many things to do, too little time!)
- ▶ Working with students is a rewarding time investment

# Students at NBI

## Undergrads

- ▶ First direct exposure to research
- ▶ Well-defined project
- ▶ ~4 months
- ▶ Topics closely linked to content courses



2018  
Siqiao Mu  
(Caltech)  
*Unitarity bounds of astrophysical neutrinos*  
PRD 2018



2020  
Niels Gustav Willeßen  
*Unitarity bounds of astrophysical neutrinos*  
JCAP 2021



2021  
Jonathan Baltazar  
*Decay of high-energy cosmic neutrinos*



2024  
Franciszek Nawrocki  
*Investigating neutrino oscillations at astrophysical energy scales*



2025  
Julie Harder Gabrielsen  
*A method to compute neutrino oscillation probabilities for arbitrary time-dependent Hamiltonians*



2025  
Asmus Bo Klingemann  
*Neutrino tomography of the Earth in ESSnuSB*

## MSc

- ▶ Project led by student
- ▶ Well-defined goal
- ▶ Solutions defined by the student
- ▶ ~1 year
- ▶ Topics firmly beyond courses
- ▶ Paper



2019–2020  
Charlotte Rosenstrøm  
*Bounds on secret neutrino interactions from high-energy astrophysical neutrinos*  
PRD 2020



2020–2021  
Kjartan Másson  
*Secret interactions of ultra-high-energy neutrinos*



2021–2022  
Marie Hansen  
*Interactions between high-energy cosmic neutrinos and axions*



2022–2024  
Anirudh Bhatnagar  
*Probing quantum-gravity decoherence in high-energy neutrinos from the active galaxy NGC 1068*



2023–2024  
Federico Testagrossa  
*Two-detector flavor sensitivity to ultra-high-energy cosmic neutrinos*



2024  
Andreas Knage  
*Proton-dark matter interaction in the active galactic nucleus NGC 1068*



2024–2025  
Johannes Vos  
*Ultra-high-energy neutrinos from dark matter trapped inside the Earth*



2025  
Aske Matthiesen  
*Flavor measurements in TAMBO*



2025  
Yifei Li  
*Ultra-high-energy neutrino flux models in GRAND*

# Students at NBI

## PhD

- ▶ State-of-the-art research
- ▶ Well-defined general plan
- ▶ Freedom to explore
- ▶ ~3 years
- ▶ Several papers
- ▶ Prepare for a career in academia if desired



2020–2023  
Víctor Valera  
*Pushing neutrino  
physics to the cosmic  
frontiers*  
Undergrad: UNI  
MSc: ICTP Trieste



2021-2025  
Bernanda Telalovic  
*Novel tests of new physic  
with astrophysical neutrinos*  
Undergrad: U. Melbourne  
Msc: U. Tübingen

# Take-aways

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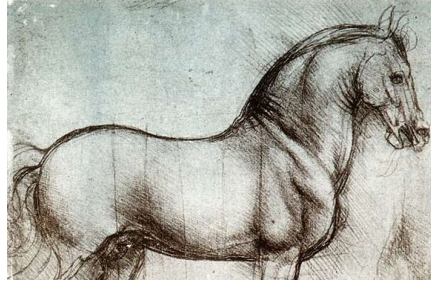
Lascaux, 15000 BC

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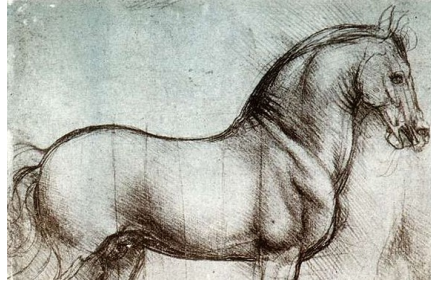
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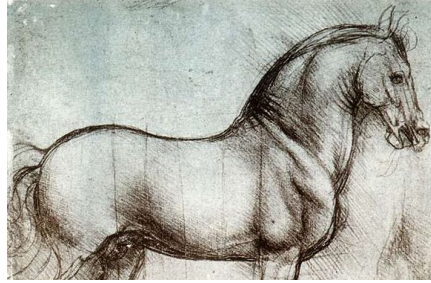
Tiepolo, 1773

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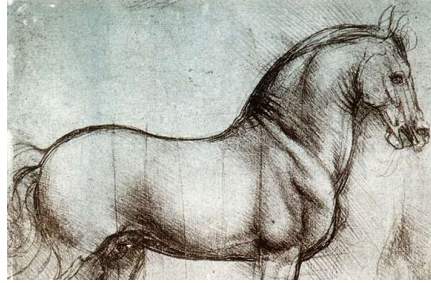
Kandinsky, 1911

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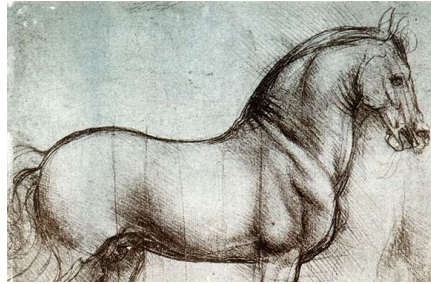
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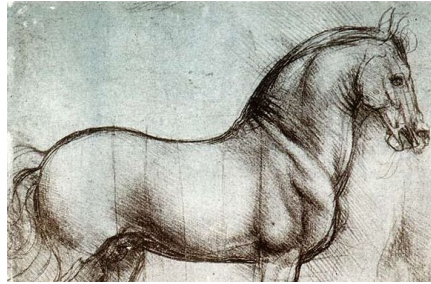
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Learn everything  
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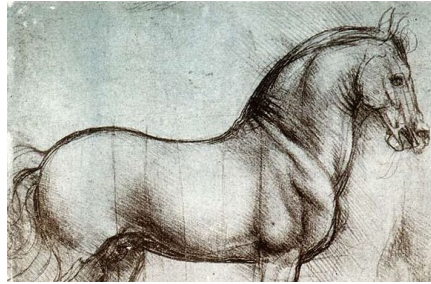
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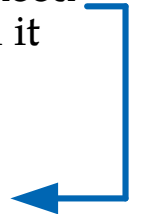
**✓** Yes

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Do research

Learn what you need  
when you need it

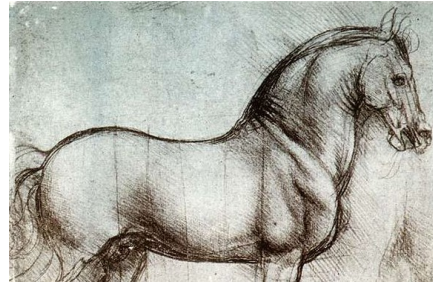


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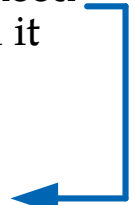
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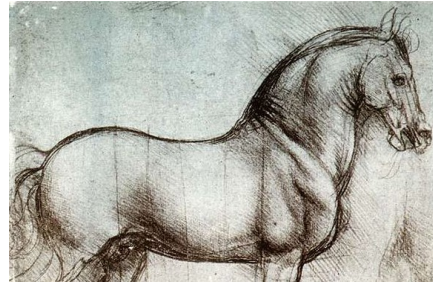
- ▶ Build your own identity as a researchers

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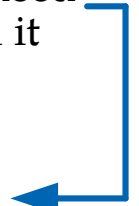
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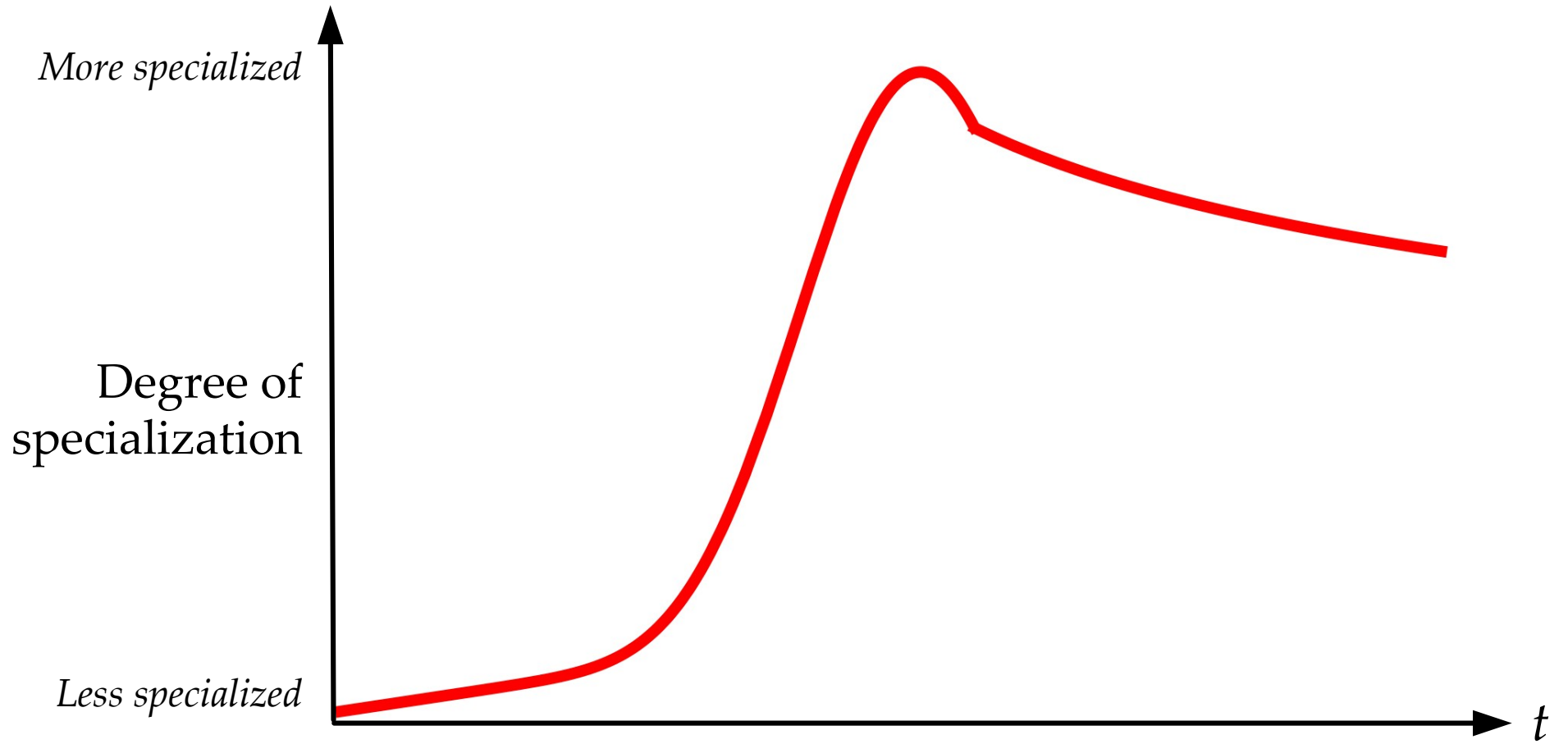
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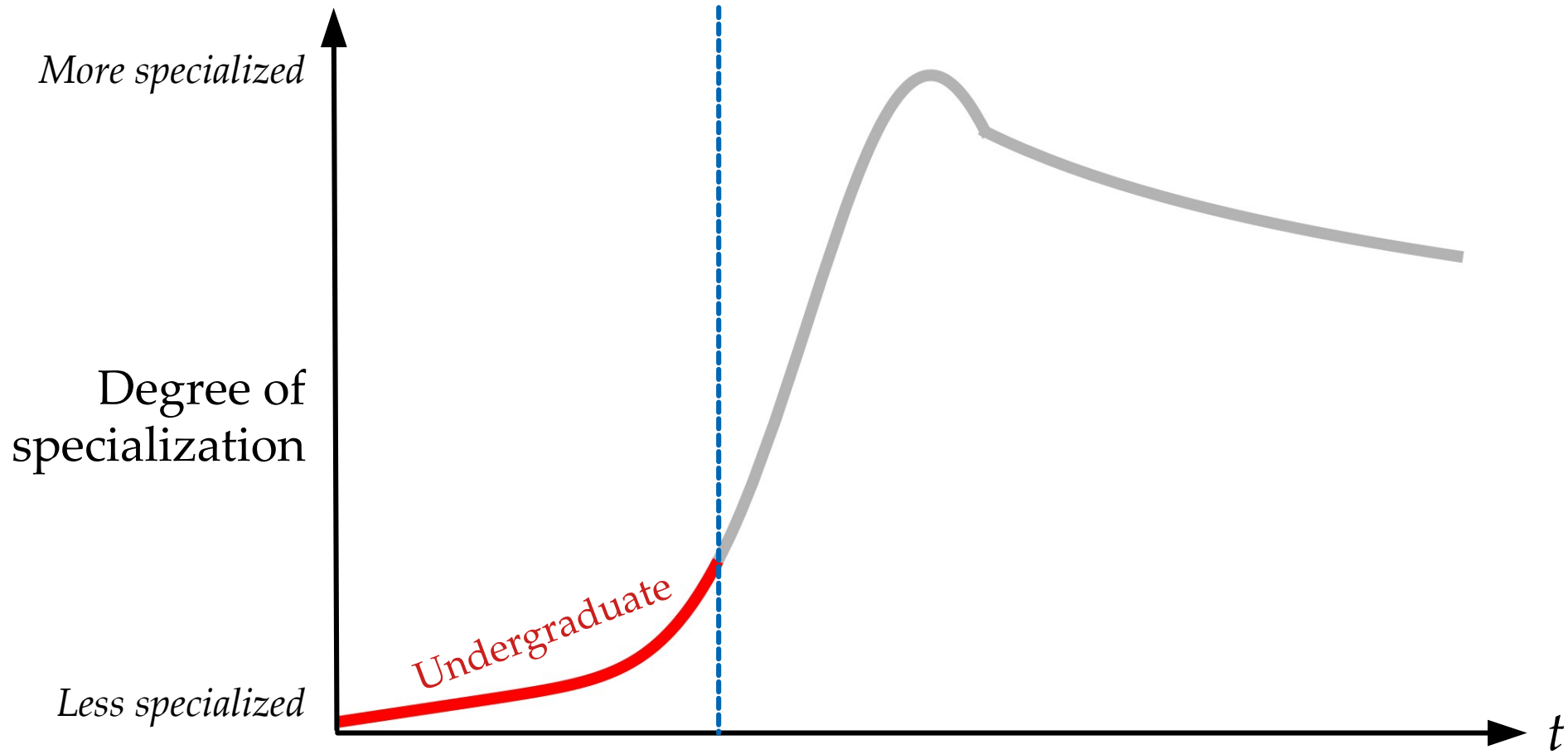


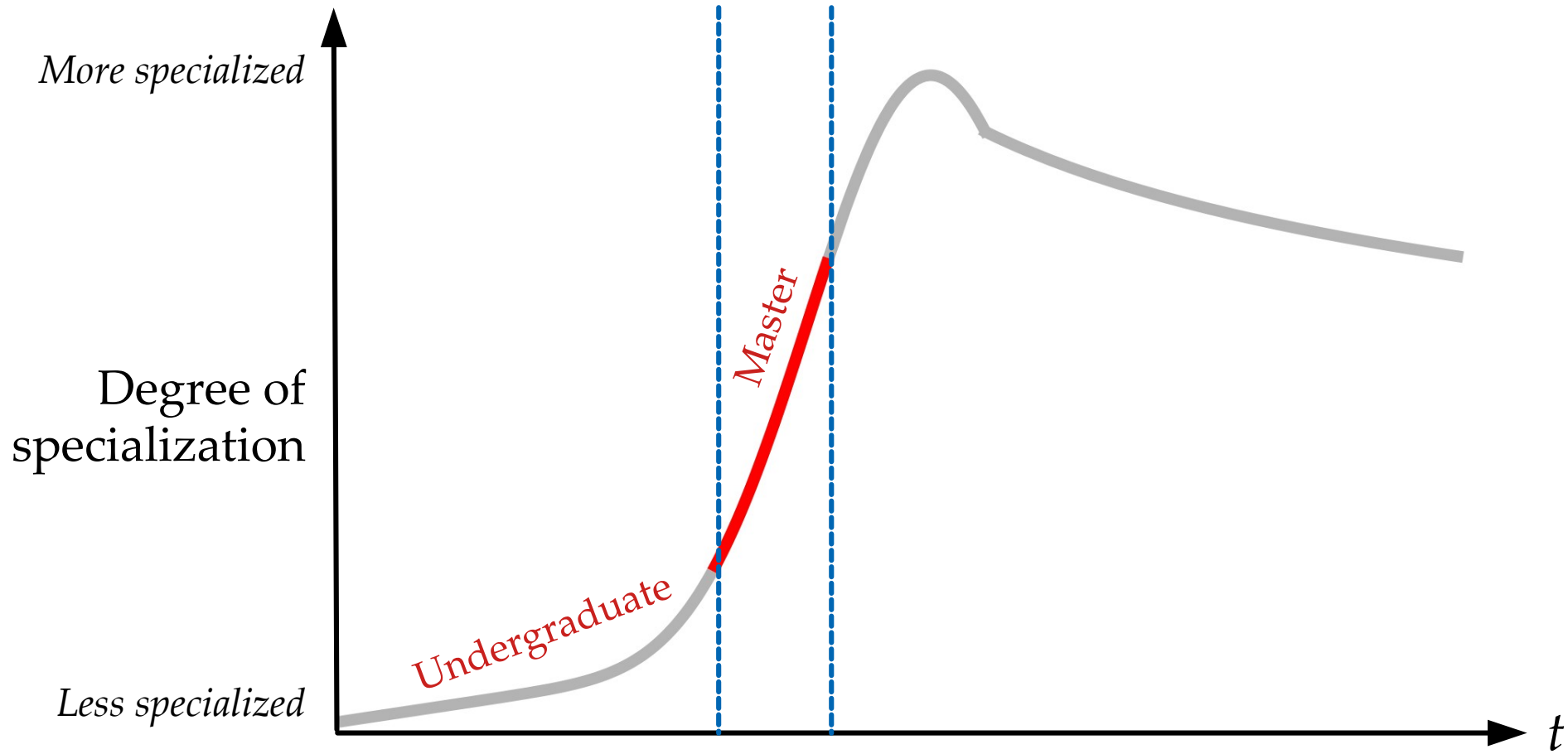
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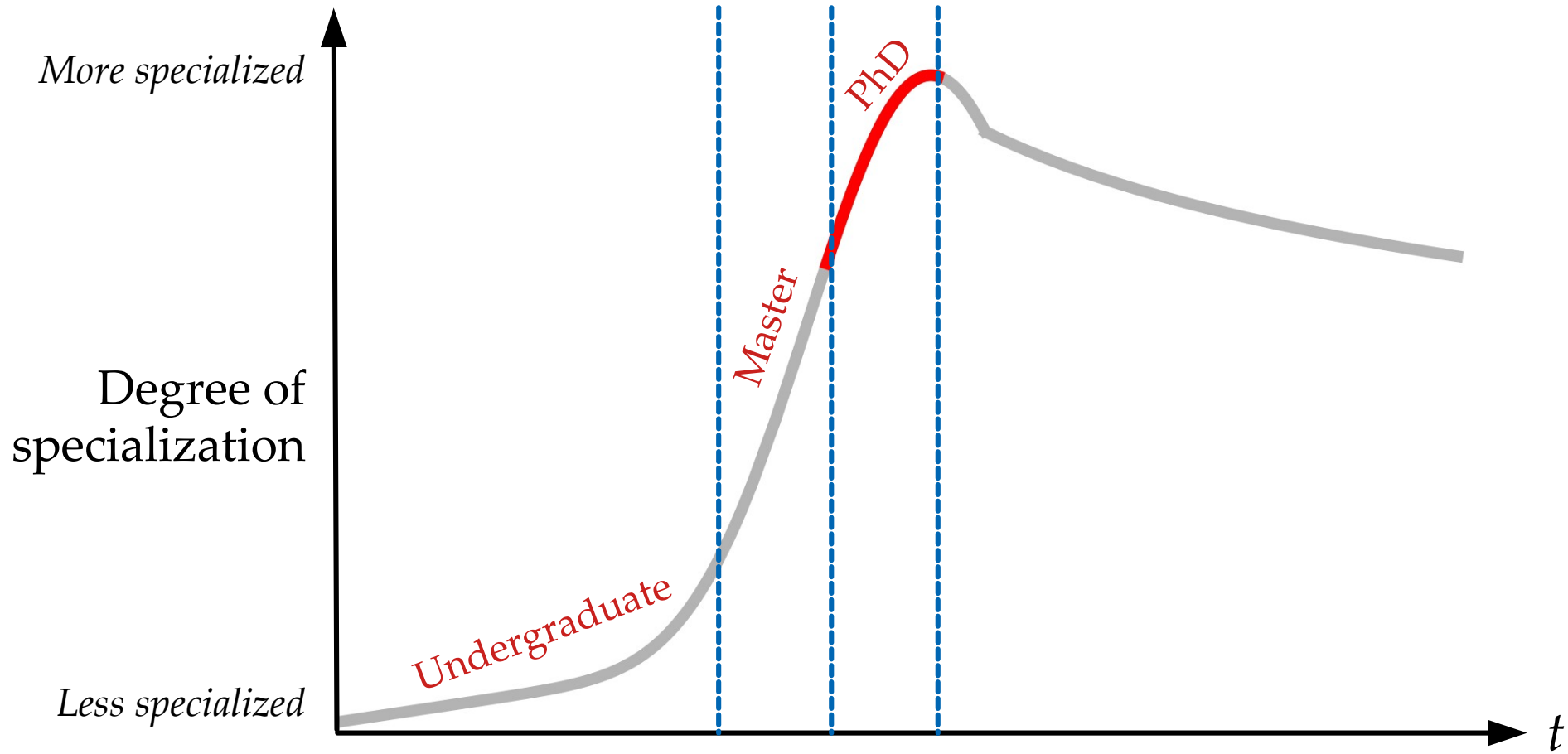
*What interests you? What is your competitive advantage? What do people associate you with?*

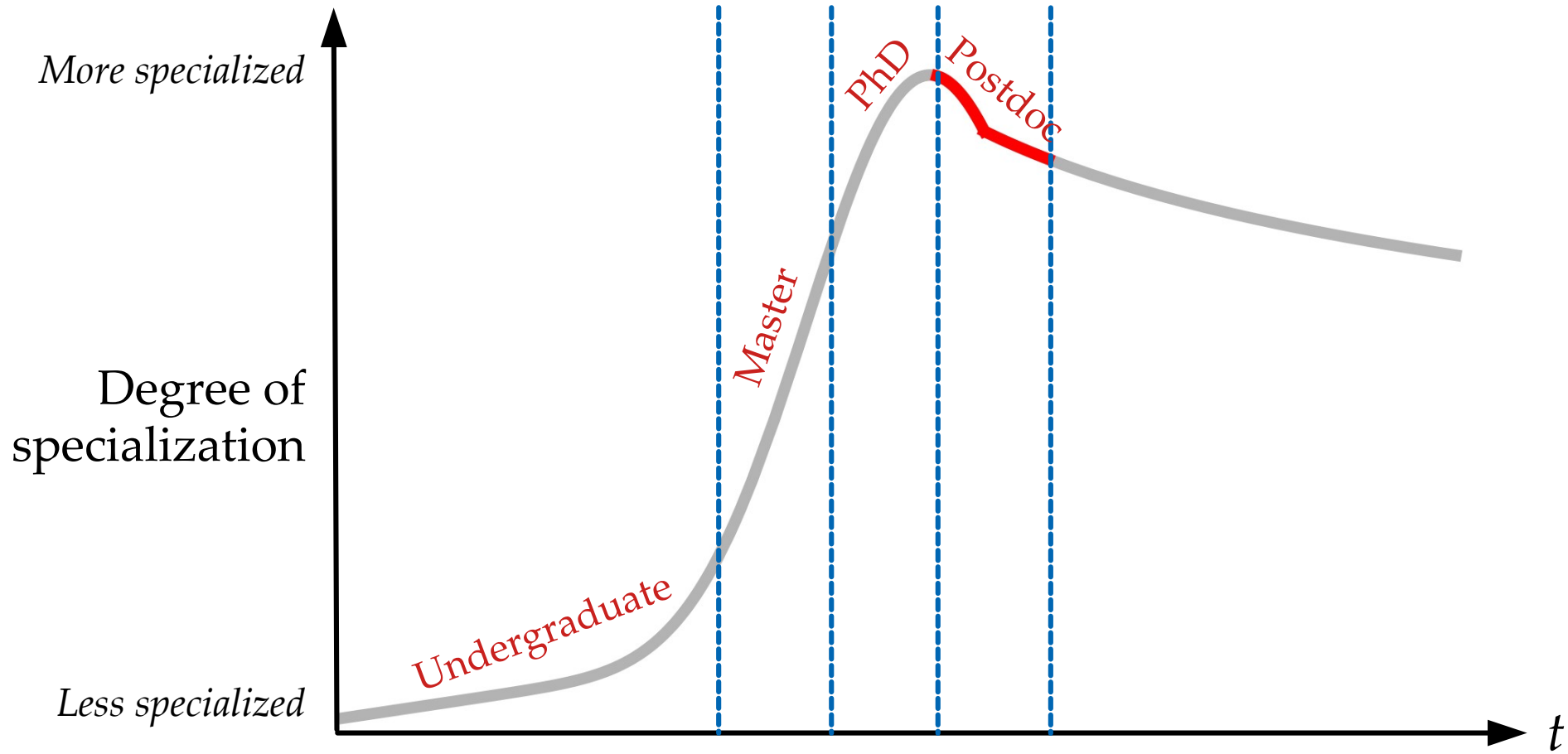


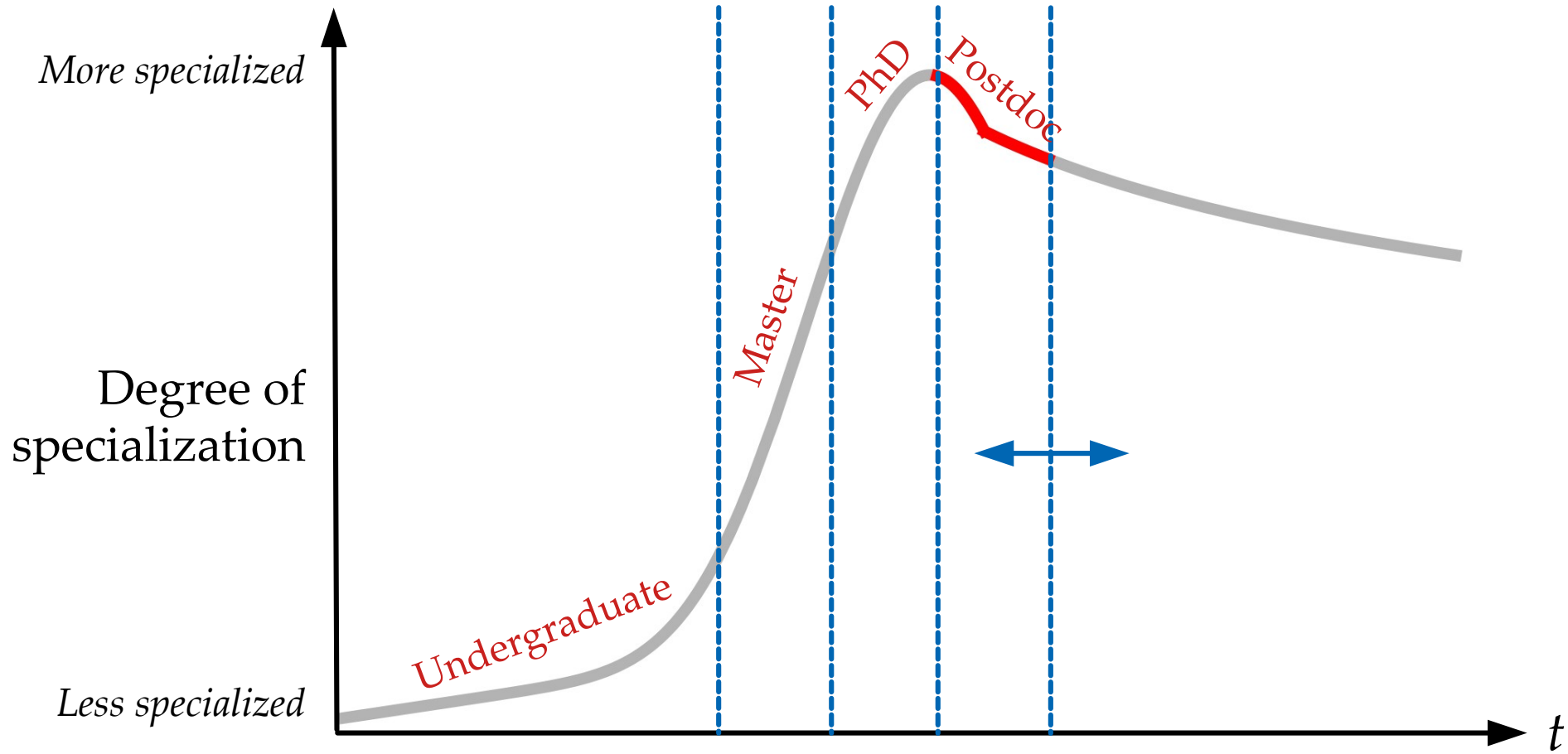


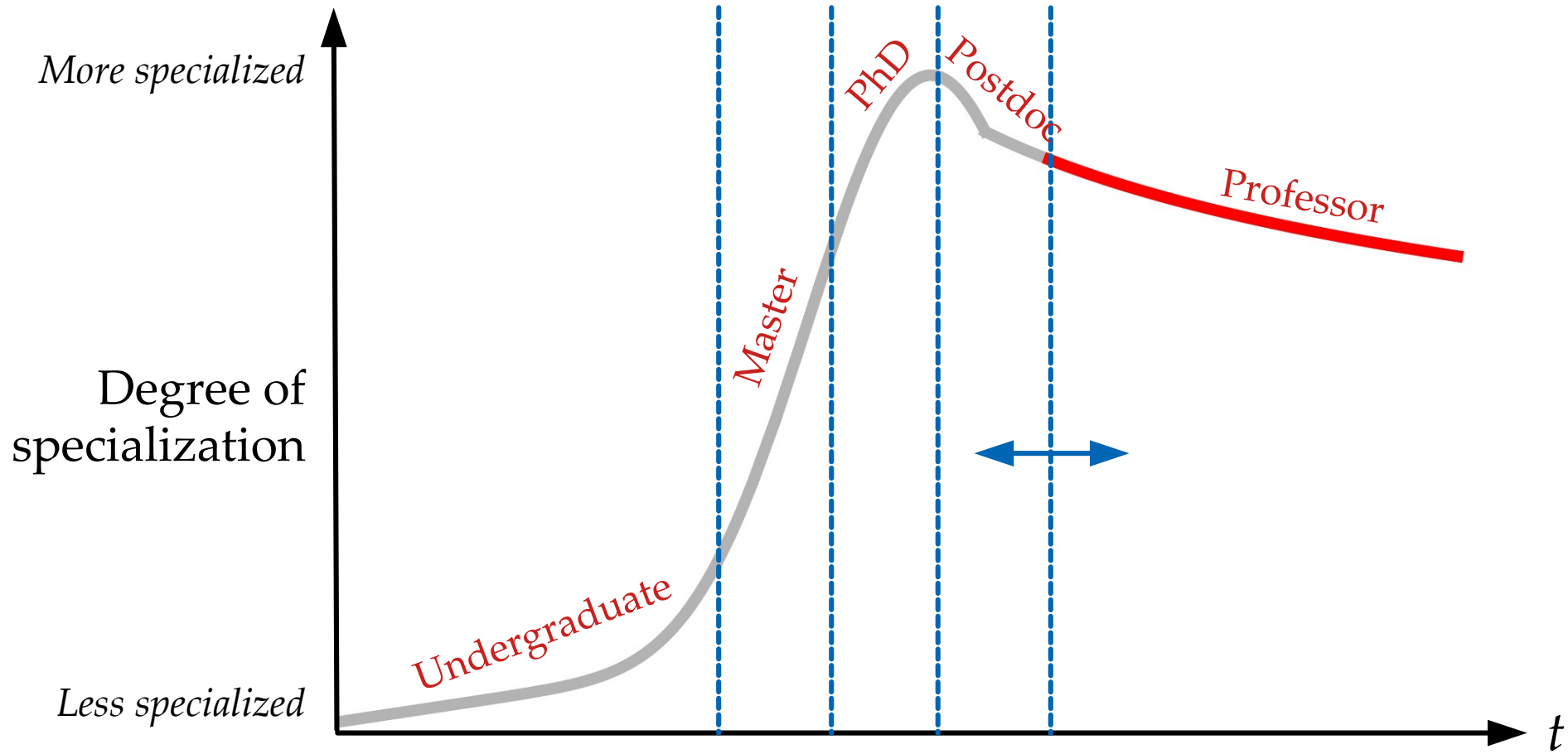


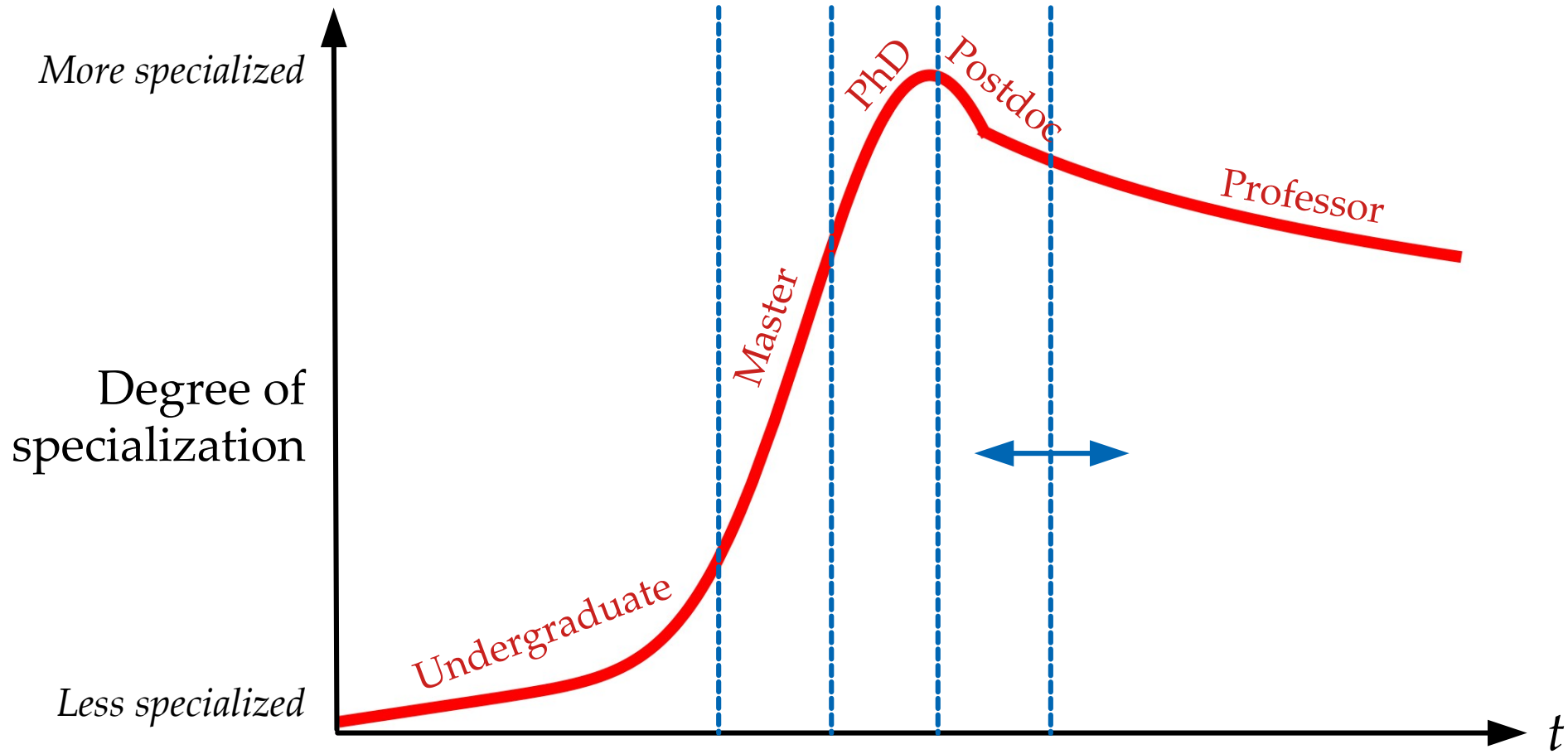




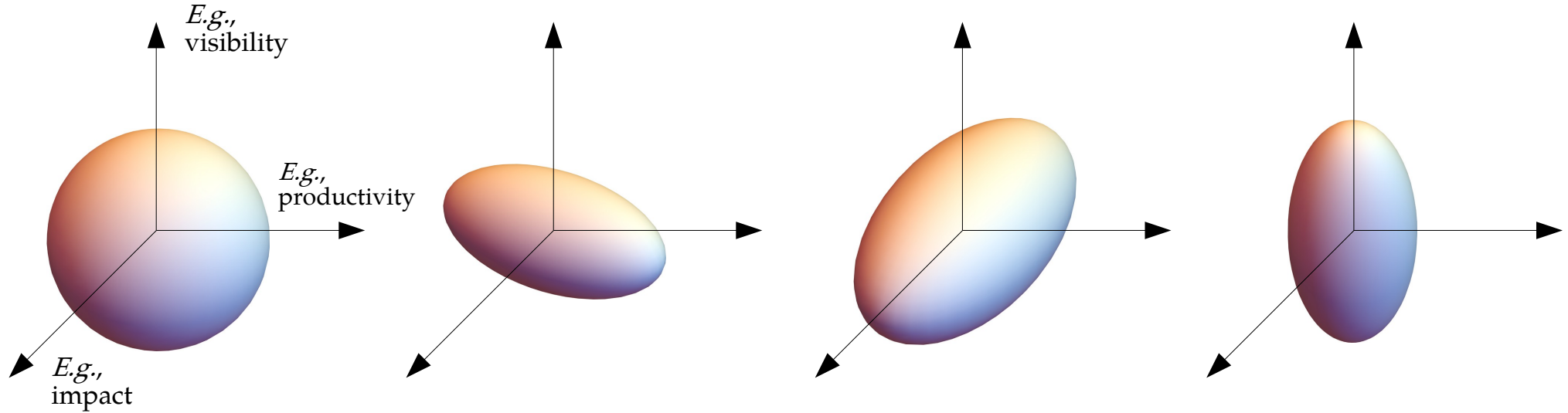








# Academic success is a multidimensional function



$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 + \left(\frac{z}{c}\right)^2 = 1$$

$$\text{Success: } V = \frac{4}{3}\pi abc$$

If any of  $a$ ,  $b$ , or  $c$  are zero,  $V$  is zero

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Most important: Make good science

Second most important: Communicate it effectively!

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*Explain the core of your work in the span of a minute*

Featured in Physics

Editors' Suggestion

Access by University of Copenhagen

[Go Mobile »](#)

## Universe's Worth of Electrons to Probe Long-Range Interactions of High-Energy Astrophysical Neutrinos

Mauricio Bustamante and Sanjib Kumar Agarwalla  
Phys. Rev. Lett. **122**, 061103 – Published 12 February 2019

**Physics** See Synopsis: [Neutrino Probes of Long-Range Interactions](#)

## Deutsches Elektronen-Synchrotron DESY A Research Centre of the Helmholtz Association



2015/04/10

[Back](#)

### Gamma-ray bursts as cosmic particle accelerators

Study provides new insights into the universe's most powerful explosions

This approach can not only explain the observed strong variations in the light curves of gamma-ray bursts. A consequence of this model is also that neutrinos, cosmic rays and gamma-rays must be produced in completely different regions of the jets. This can explain, why the expected flux of neutrinos could not be found. "We expect that the next generation of neutrino telescopes, such as IceCube-Gen-2, will be sensitive to this minimal flux that we're predicting", says Bustamante. In contrast to earlier models, this estimate is more robust and does only weakly depend on the characteristics of individual gamma-ray bursts.

PHYSICS

## Astronomers Propose Huge New Telescope System to Understand the Most Energetic Particles Ever Detected

# GIZMODO

**GIM** Ryan F. Mandelbaum  
10/29/18 4:20PM • Filed to: **GRAND**

14.3K 24 5 [f](#) [t](#) [e](#) [l](#)

"Blazars could maybe make neutrinos in a wide energy range, or maybe it could be something else making these higher-energy neutrinos," Mauricio Bustamante, editor of the experiment's white paper and a postdoc at the Niels Bohr Institute in Copenhagen, told Gizmodo. "We hope it's as interesting as possible."

MIT  
Technology  
Review

Sustainable Energy

## How Neutrino Beams Could Reveal Cavities Inside Earth

Geophysicists want to use neutrinos to 'x-ray' the Earth, a technique that could reveal undiscovered oil fields. But how practical is such a scheme?

by Emerging Technology from the arXiv

Feb 1, 2012

NEUTRINOS | NEWS

## The case of the disappearing neutrinos

15 January 2018

In an additional analysis of six years of IceCube data, Amy Connolly and Mauricio Bustamante of Ohio State University employ an alternative approach which uses 58

IceCube-contained events (in which the neutrino interaction took place within the detector) to measure the neutrino cross-section. Although these events mostly have well-measured energies, their neutrino zenith angles are less well known and they are also much less numerous, limiting the statistical precision.

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NEWS RELEASE 28-OCT-2019

## Giant neutrino telescope to open window to ultra-high-energy universe

SCIENCE CHINA PRESS

Media Contact

Mauricio Bustamante  
mbustamante@nbi.ku.dk

<http://www.scichina.com/>

PHYSICS

## Bizarre Particles Keep Flying out of Antarctica's Ice, and They Might Shatter Modern Physics

"It was clear from the start that if the ANITA anomalous events are due to particles that had propagated through thousands of kilometers of Earth, then those particles were very likely not SM particles," said Mauricio Bustamante, an astrophysicist at the Niels Bohr Institute at the University of Copenhagen, who was not an author on the new paper.

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To our fellow citizens  
Who, for love of truth,  
Take from their own wants  
By taxes and gifts,  
And now and then send forth  
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To forward the search  
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- ▶ Down the road: consider a TEDx event



# Funding opportunities in Peru

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- ▶ Beca 18: For senior high-school students
- ▶ Beca Mujeres en Ciencia: For female senior high-school students
- ▶ Beca Inclusión: For people with disabilities
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[investigacion.pucp.edu.pe/convocatoria/programa-apoyo-iniciacion-en-investigacion-pain](http://investigacion.pucp.edu.pe/convocatoria/programa-apoyo-iniciacion-en-investigacion-pain)
- ▶ Programa de Apoyo a la Investigación para Estudiantes de Posgrado (PAIP): For grad students  
[investigacion.pucp.edu.pe/convocatoria/programa-apoyo-investigacion-para-estudiantes-de-posgrado-paip](http://investigacion.pucp.edu.pe/convocatoria/programa-apoyo-investigacion-para-estudiantes-de-posgrado-paip)
- ▶ Beca Huiracocha: For PhD students  
[posgrado.pucp.edu.pe/beca-fondo/beca-huiracocha](http://posgrado.pucp.edu.pe/beca-fondo/beca-huiracocha)

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Full list of undergraduate scholarships (State, private): [www.pucp.edu.pe/pregrado/becas](http://www.pucp.edu.pe/pregrado/becas)

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**Undergrad:** do not pick topics narrowly (attend PUCP Physics colloquia!)

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▶ Different from reading a non-scientific text

▶ Key skill: single out main results quickly

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# Want more good science content? Start here



How to become a good theoretical physics, by Gerard 't Hooft (1999 Nobel):  
[webspaces.science.uu.nl/~gadda001/goodtheorist/index.html](http://webspaces.science.uu.nl/~gadda001/goodtheorist/index.html)



Coloquios de Física PUCP

Videos of colloquia since 2011 – [sites.google.com/site/fisicapucp](http://sites.google.com/site/fisicapucp)



Científicos.pe

Interviews, resources, opinion articles – [cientificos.pe](http://cientificos.pe)

arXiv.org

arXiv

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How to become a good theoretical physics, by Gerard 't Hooft (1999 Nobel):  
[webspacescience.uu.nl/~gadda001/goodtheorist/index.html](http://webspacescience.uu.nl/~gadda001/goodtheorist/index.html)



Coloquios de Física PUCP

Videos of colloquia since 2011 – [sites.google.com/site/fisicapucp](http://sites.google.com/site/fisicapucp)



Científicos.pe

Interviews, resources, opinion articles – [cientificos.pe](http://cientificos.pe)

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## For the neutrino enthusiasts:



All things neutrino

Curated by Fermilab – [neutrinos.fnal.gov](http://neutrinos.fnal.gov)



Neutrino Unbound

Papers on neutrinos, curated by Carlo Giunti – [www.nu.to.infn.it](http://www.nu.to.infn.it)

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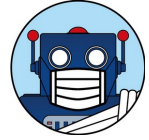
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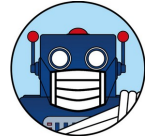
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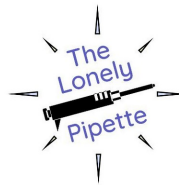
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### Mindscape, by Sean Carroll

In-depth interviews about many topics

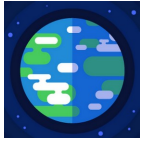


### The Lonely Pipette

Tips from scientists for how to do better science

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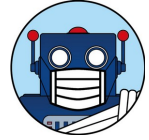
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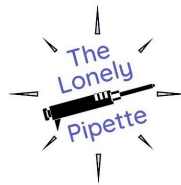
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## Instagram:



### @astrocarlaa

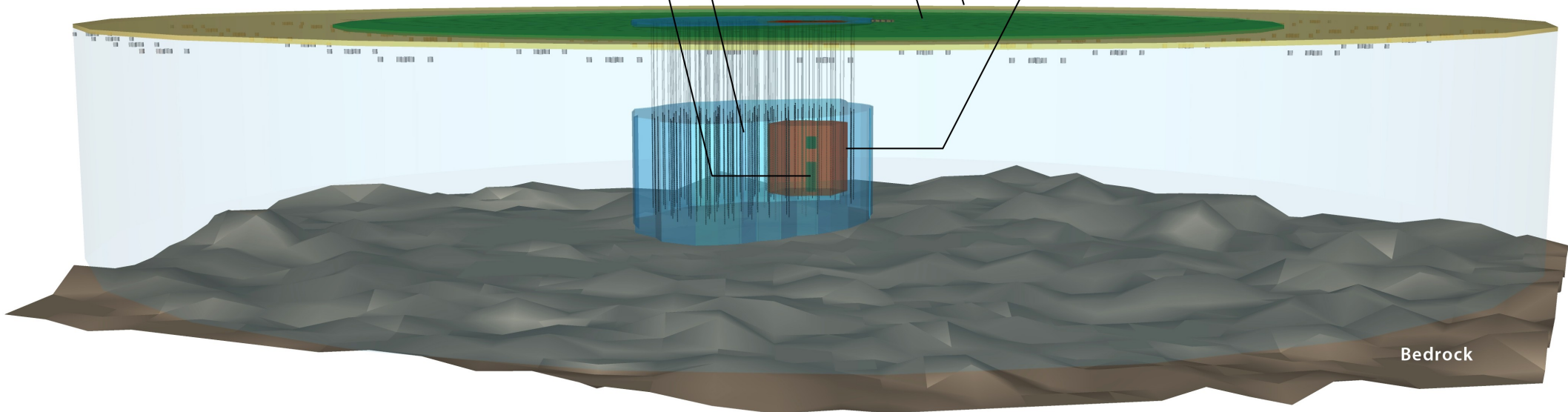
Astronomy and astrophysics, explained

Backup slides

# IceCube-Gen2

arXiv:2008.04323

Radio Array  
Surface Array  
Main Array  
Core (PINGU)  
IceCube-86, IceTop

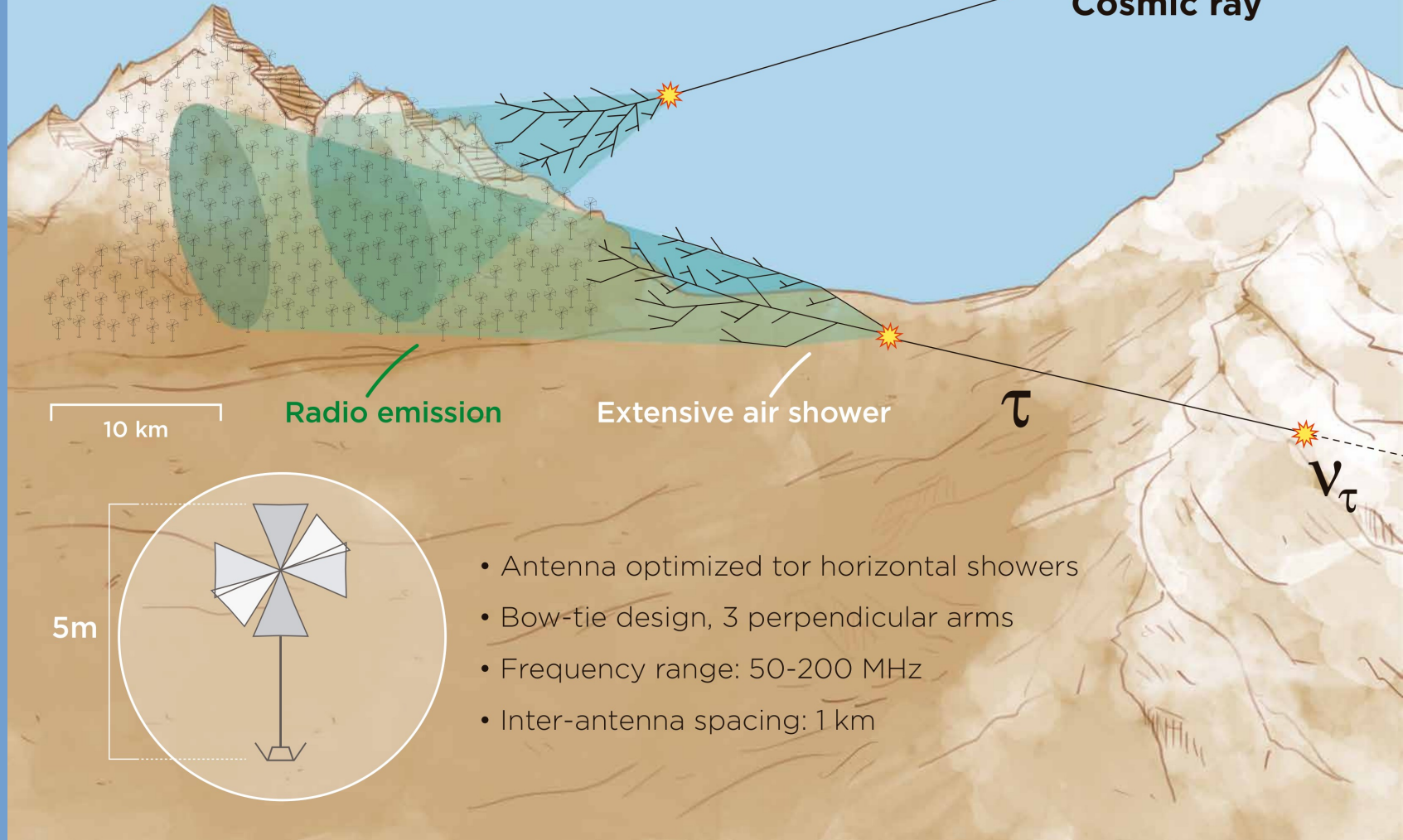


Bedrock



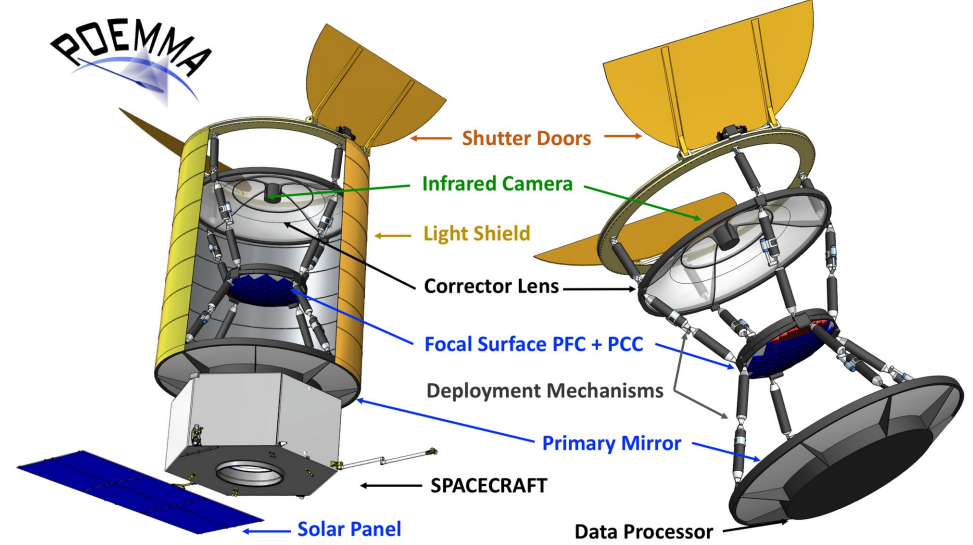
# Giant Radio Array for Neutrino Detection

arXiv:1810.09994

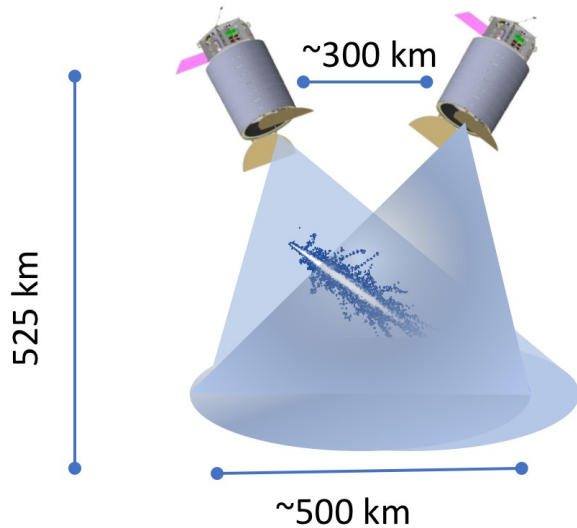


# POEMMA: Probe of Extreme Multi-Messenger Astrophysics

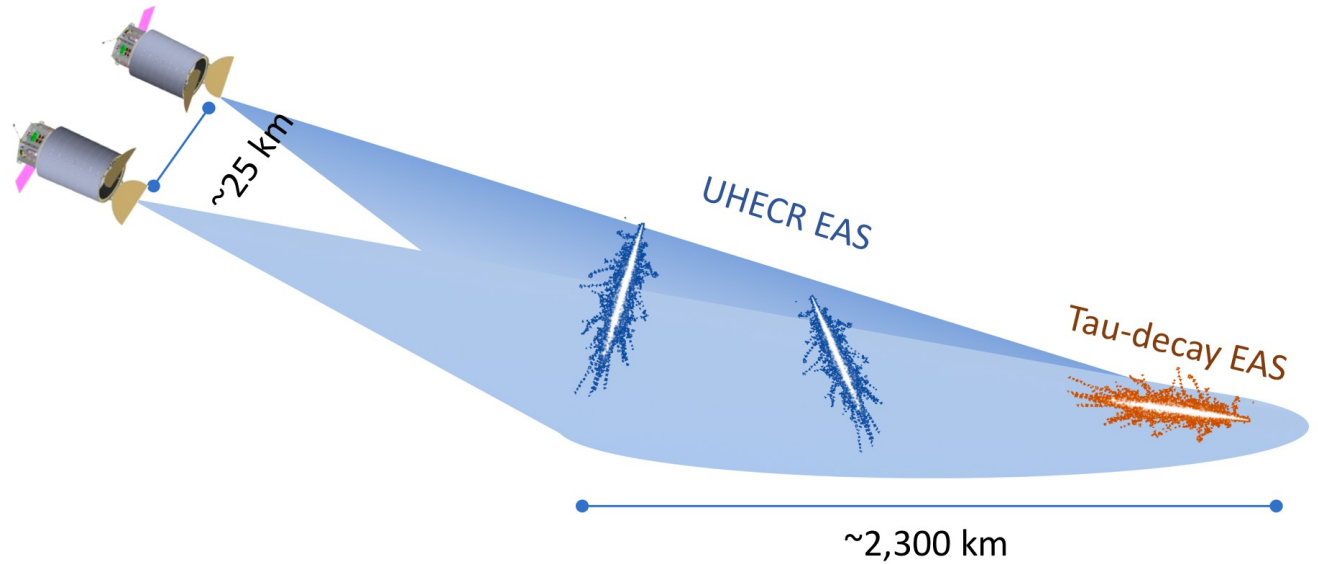
arXiv:2012.07945



POEMMA-Stereo



POEMMA-Limb



# TAMBO

arXiv:2002.06475

AIR SHOWER:

3 - 10 KM LENGTH  
200 M DIAMETER

DECAY

$\tau$

RANGE:  
50 M - 5 KM

ROCK

> 4 KM SHIELDING FROM  
BACKGROUND MUONS

$\nu_{\tau}$

CHARGED-CURRENT  
INTERACTION

~100 M  
SEPARATION

WATER CHERENKOV  
DETECTOR ARRAY

~M<sup>3</sup> EACH

DEEP VALLEY

Find these slides at  
[mbustamante.net/talks](http://mbustamante.net/talks)

Feel free to contact me at [mbustamante@nbi.ku.dk](mailto:mbustamante@nbi.ku.dk)